TELESEISMIC IMAGING: FIELD STUDY IN SOUTHERN ALBERTA AND NUMERICAL SIMULATIONS OF INVERSE SCATTERING

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#### Abstract

This thesis consists of two parts. Part I presents results from a Lithoprobe teleseismic experiment undertaken across southern Alberta. Relative $P$-wave delay-times from 293 earthquakes have been inverted for upper mantle velocity perturbations. The recovered model reveals a high velocity anomaly underlying a substantial portion of the southern Hearne Province to depths of $200-250 \mathrm{~km}$ which is interpreted as the signature of deep-seated lithospheric structure. This result suggests that, contrary to recent tectonic models, the bulk of the lithosphere in this region has remained essentially intact. In particular, it appears unlikely that evidence for extensive lower crustal melting is due to wide-scale lithospheric delamination. However, observed high mantle conductivity may be the result of small volumes of connected hydrous minerals or some other conductive species introduced during subduction that contributed to the construction of a root. Multi-event $S K S$-splitting results yield an average delay-time of $0.82 \pm 0.30 \mathrm{~s}$ and fast polarization direction of $45^{\circ} \pm 8^{\circ}$ which broadly coincides with both the presumed orientation of fossil strain fields related to the ca. 1.8 Ga NW-SE shortening of the Hearne Province and absolute North America plate motion. Processing of receiver functions yields Moho depth estimates which are fairly uniform ( $\sim 38 \mathrm{~km}$ ) beneath northern stations, but show crustal thickening ( $>40 \mathrm{~km}$ ) within the Medicine Hat Block.

Part II investigates the formal inversion of synthetic teleseismic $P$-coda waves for subsurface elastic properties using an asymptotic method which assumes single-scattering. The model comprises an idealized lithospheric suture zone. Two dimensional, pseudo-spectral synthetic seismograms representing a plane $P$-wave incident upon this structure are preprocessed to extract an estimate of the scattered wavefield. These data are employed in a series of experiments that examine the dependence of multi-parameter inversion on a range of input parameters and demonstrate: i) the contrasting sensitivity which forward- and back-scattered waves display to structural recovery; ii) the diminution of the problem null-space accompanied by increased source coverage; iii) improvements in model reconstruction achieved through simultaneous treatment of multiple scattering modes; and iv) the robustness of the approach for data sets with noise levels and receiver geometries that approach those of field experiments.


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## Chapter 1

## Motivation for Thesis

The structure and dynamics of the Earth's lithosphere have a significant impact upon human society by virtue of our reliance on natural resources, and the substantial devastation that is regularly wrought by earthquakes and volcanic activity. These two factors provide much of the impetus for the use of geophysics in understanding shallow crustal structure and processes. However, geophysics, and seismology in particular, are also used in more fundamental studies to characterize deeper portions of the lithosphere and upper mantle. In both fundamental and applied contexts there is a continual drive to develop novel approaches to processing that afford improved resolution for challenging structural targets.

In Canada, Lithoprobe studies, spearheaded by the seismic reflection method, have assembled much detailed information concerning lithospheric structure and evolution across the Canadian landmass. It is worth noting, however, that a majority of the lithospheric plate resides below the crust-mantle boundary and, due to the stronger rheology of mantle materials, it may exert a substantial influence on the near-surface evolution. The use of seismic reflection methods in characterizing the lithospheric mantle is hindered by the limited depth penetration and frequency bandwidth of anthropogenic sources. In contrast, teleseismic waves (seismic waves generated by earthquakes at epicentral distances $>30^{\circ}$ ) do not suffer from these shortcomings to the same degree and allow a more comprehensive characterization of sub-crustal lithospheric structures.

This thesis investigates two different aspects of the use of teleseismic analysis techniques in
the recovery of lithospheric structure. In Part I, the application of travel-time tomography, shear wave splitting, and radial receiver functions to a suite of teleseismic seismograms recorded in the Hearne Province of southern Alberta is described. Geophysical constraints resulting from these analyses are subsequently interpreted to shed light on the tectonic evolution of a region which is hidden from view by the thick sequences of the Western Canada Sedimentary Basin. As such, the work documented in Part I of this thesis presents a passive-source study of the upper mantle of Alberta at a length-scale comparable to or exceeding large-scale refraction profiling (e.g., Deep Probe), and stands as an important contribution to the understanding of the assembly of the Laurentia supercontinent.

Part II of this thesis investigates a new high-resolution teleseismic imaging technique that draws motivation from the increasing availability of broadband, three-component instruments and the seismic migration problem in hydrocarbon exploration. This method provides a rigorous basis for the inversion of teleseismic wavefields recorded on dense receiver arrays for Earth's elastic parameters. I discuss the implementation of this new technique on a set of synthetic data with the objective of assessing the potential and limitations of the inversion methodology in recovering discontinuous model structure. In particular, this work identifies the roles that various parameters play in the inversion process, and demonstrates the feasibility of the approach prior to its application to field data.

## Part I

INTEGRATED TELESEISMIC STUDIES OF THE SOUTHERN ALBERTA UPPER
MANTLE

## Chapter 2

## Introduction and Study Area Overview

Part I overview: Advances in our understanding of tectonic processes and associated lithospheric structure have often relied on remote sensing geophysical methods and, in particular, seismology. In regions where crustal basement rocks are concealed by thick sequences of sedimentry strata, as in southern Alberta, the use of such methods is essential if the tectonic evolutionary history is to be uncovered. The work presented in Part I of the thesis discusses a teleseismic experiment for which the objective was to obtain constraints on the lithospheric structure of southern Alberta, and to better understand the region's tectonic evolution.

### 2.1 Motivation and Research Goal

The formation of the supercontinent Laurentia, the Paleoproterozoic (ca. 2.0-1.6 Ga) landmass which forms the cratonic root of present day North America, is not well understood and numerous questions pertaining to the assembly of its constituent tectonic fragments (e.g., Hoffman [1988]) remain unresolved. One particularly enigmatic region is southern Alberta where sedimentary cover precludes direct investigation of the crustal basement. Some of the first observations relating to the region's tectonic evolution emerged from early studies of sedimentary strata [Deiss, 1941] which indicated an anomalous history of epeirogenic motion. Vertical movements of crustal blocks of the order of 100 's of km in horizontal extent were noted, suggesting a deep, possibly subcrustal, origin. Pioneering seismic reflection studies [Kanasewich and Cumming,

1965; Kanasewich et al., 1969; Clowes and Kanasewich, 1972] and early refraction studies (e.g., Chandra and Cumming, [1972]) further documented dramatic variations in crustal structure between the various basement domains.

During the past decade, research undertaken as part of LITHOPROBE's Alberta Basement Transect has provided better constrained and more detailed interpretations of the tectonic evolution of much of the region's crustal basement (e.g., Eaton et al., [1999]; Ross et al., [2000]; Gorman et al., [2001]; Lemieux et al., [2000]). However, one issue which remains poorly understood is the role of the subcrustal lithosphere in the formation and subsequent evolution of southwestern Laurentia. The limited depth penetration and frequency bandwidth of crustal seismic reflection profiling preclude a comprehensive examination of the region's upper mantle (i.e., to depths exceeding 150 km ; see Eaton et al., [2000]). Large-scale refraction profiling (e.g., DEEP PROBE) has been more successful in constraining upper mantle velocity variations (e.g., Gorman et al., [2001]), although it is still limited by signal penetration in depth. Regional-scale, broadband teleseismic investigation does not suffer from these limitations to the same extent and is able to extend the resolution of velocity structure through the entire upper mantle column. Thus, the objective of the present study is to improve our current understanding of deep-seated lithospheric structure in southern Alberta through the application of a suite of teleseismic processing techniques to data recorded on a portable broadband seismic array (Chapter 3). This information will, in turn, provide additional geophysical contraints on the tectonic evolution of southern Alberta (Chapter 4).

### 2.2 Tectonic Overview of the Alberta Basement

The Hearne Province is one of several microcontinental fragments constituting the Laurentian supercontinent and has been geologically mapped where exposed in the Canadian Shield of Saskatchewan and the Northwest and Nunavut Territories (e.g., Bickford et al., [1994]). A comprehensive understanding of its basement architecture and evolution in southern and central Alberta, western Saskatchewan, and northern Montana has, however, been hindered by the overlying Phanerozoic Western Canada Sedimentary Basin. Some information concerning the region's crustal basement is afforded by geochronological dating of drill core and potential field mapping which reveal a heterogeneous mosaic of predominantly Archean-aged crustal blocks [Ross et al., 1991; see Figure 2.1].

At present, however, there is a lack of consensus concerning the tectonic affiliations of these internal blocks. In particular, the location of the boundary between the Hearne Province and the Wyoming Province to the south is in question. Some authors associate the Medicine Hat Block (MHB) with the Hearne Province (e.g., Ross et al., [1991]) and propose the Great Falls Tectonic Zone as the site of the Hearne-Wyoming suture (e.g., O'Neill and Lopez, [1985]). Other investigators associate the MHB with the Wyoming Province (e.g., Hoffman, [1990]) and interpret the Vulcan Structure to represent the interprovincial suture (e.g., Eaton et al., [1999]). Furthermore, geophysical evidence does not preclude the MHB from once existing as an autonomous Archean crustal fragment that is now delineated by sutures with both the Hearne and Wyoming Provinces at the Vulcan Structure and the Great Falls Tectonic Zone, respectively [Gorman et al., 2001]. The tectonic relation between the MHB and its neighbouring domains is rendered still more enigmatic by observations of differential subsidence along its northern margin as revealed through changes in facies and thickness patterns identified in early stratigraphic studies (e.g., Deiss, [1941]).


Figure 2.1: Simplified tectonic map of the study area (after Ross et al., [1991]) and locations of teleseismic stations. Circles and hexagons represent portable and CNSN stations, respectively. Labelled domains: BH - Buffalo Head, EH - Eyehill High, GFTZ - Great Falls Tectonic Zone, La - Lacombe Domain, LB - Loverna Block, MHB - Medicine Hat Block, Ri - Rimbey Arc, STZ Snowbird Tectonic Zone, Ta - Taltson Magnetic Arc, Th - Thorsby Magnetic Low, VS - Vulcan Structure, W - Wabamun Domain, WB - Wathamun Batholith. Labelled relevant active-source profiles: C7-CAT line 7, S30-SALT line 30, S31-SALT line 31. Black line (with triangles) shows limit of the Cordilleran Deformation Front (DF).

The eastern margin of the Hearne Province is marked by the Trans-Hudson Orogen, a complex Paleoproterozoic orogenic belt that developed during relative convergence of the Superior and Hearne Provinces in the interval 1.9-1.7 Ga. Tectonic activity along this boundary included inward-dipping plate consumption and the subsequent imbrication of continental margin rocks beneath the Hearne Province ca. 1.85-1.78 Ga [Ross et al., 2000]. The province's northwestern boundary is marked by the Snowbird Tectonic Zone, a distinct potential field anomaly which, based on constraints from recent crustal reflection profiling and long period electromagnetic surveys, has been interpreted as a lithospheric-scale Proterozoic suture zone [Ross et al., 1991; 2000]. Tectonic modelling and geochronologic dating of this zone further suggest a southeastward subduction of oceanic crust beneath the Hearne Province coeval with activity along its eastern boundary. The western portion of the composite Hearne-Wyoming system is proposed to have later rifted, thus establishing a passive continental margin which was further modified through the development of the overlying North American Cordillera (e.g., Burchfiel et al., [1992]). For more detail the reader is referred to Ross et al. [2000], Eaton et al. [1999], and references therein.

### 2.3 Previous Geophysical Coverage

Southern Alberta has long been a focus of geophysical study. Early investigation of the Vulcan Structure, including one of the first crustal seismic reflection profiles, led to the initial interpretation of this domain as a Precambrian rift on the basis of vertical offset in deep reflecting horizons and a pronounced linear anomaly in gravity and magnetic map signatures which suggested an eastwest oriented graben structure [Kanasewich et al., 1969]. Further geophysical constraints were supplied by Clowes and Kanasewich [1972] and Chandra and Cumming [1972] who presented evidence for varying Moho topography based on significant lateral changes in lower crustal reflection horizons, and spatial correlations between high-velocity zones at shallow crustal levels
and Bouguer gravity highs, respectively.

In the past decade, an improved understanding of the region's architecture and tectonic evolution has emerged through the multidisciplinary Lithoprobe Alberta Basement Transect. The Central Alberta Transect (CAT), a collection of ten rectilinear, predominantly east-west profiles at $\sim 53^{\circ} \mathrm{N}$, displays crustal-scale imbrication of central and western portions of the Hearne Province with northwest vergence [Ross et al., 1995]. The Trans-Hudson Orogen Transect (THOT), which profiles the northeastern sections of the Hearne Province, recorded vergence in an opposing sense [Lucas et al., 1993; Lewry et al., 1994]. Collectively, these two profiles document the penetrative nature of the crustal-scale Paleoproterozoic shortening of the Hearne Province [Ross, 1997]. Southern Alberta Lithosphere Transect (SALT) profiles, which cross from the southern Loverna Block into the northern MHB, show crustal thickening within and southward of the Vulcan Structure, and an interpreted south-verging underthrusting of the Vulcan Structure by the Loverna Block [Eaton et al., 1999]. Examination of data from the Southern Alberta Refraction EXperiment (SAREX), a north-south profile situated east of the SALT lines ( $\sim 110^{\circ} \mathrm{E}$ ), reveals wavy undulations in the Loverna Block's velocity structure and a thick high velocity, lower crustal layer below the MHB [Clowes et al., 2001].

Subcrustal lithospheric structure beneath southern Alberta has been illuminated by the Vibroseis Augmented Listen Time (VAuLT) experiment, a set of seismic reflection profiles which span southern Loverna Block, the Vulcan Structure, and northern MHB between $\sim 113^{\circ}$ and $114^{\circ} \mathrm{E}$ [Eaton et al., 2000]. Processed VAuLT sections document south-dipping reflectivity into the upper mantle under the Loverna Block and Vulcan Structure that becomes subhorizontal beneath the MHB. This reflectivity has been interpreted to arise from compositional layering and (or) zones of ductile deformation within the mantle [Eaton et al., 2000]. To the east of the VAuLT lines, these three domains have been investigated by the continental-scale refraction experiment DEEP

Probe. The resulting profiles document two north dipping reflectors in the mantle that are interpreted to be associated with ancient subduction zones, one located to the north of the MHB, the other to the south [Gorman et al., 2001]. The interpretation of a north-dipping subduction-suture zone at the Vulcan Structure may, however, be somewhat inconsistent with the south-verging reflectivity patterns noted in SALT and VAuLT profiles [Eaton et al., 1999; 2000]. Mantle velocity structure elucidated by DEEP PROBE refraction profiling documents a marginally faster subcrustal mantle beneath the Hearne Province $\left(8.2 \mathrm{~km} \mathrm{~s}^{-1}\right)$ relative to the MHB $\left(8.1 \mathrm{~km} \mathrm{~s}^{-1}\right)$. This difference, however, diminishes at depths greater than $\sim 80 \mathrm{~km}$.

The lithospheric mantle of southern Alberta has also been characterized at greater scales using long-period seismology. Surface wave tomography [van der Lee and Nolet, 1997; Frederiksen et al., 2001], which recovers smooth three-dimensional (3-D) regional velocity structure at length-scales greater than $\sim 200 \mathrm{~km}$, shows southern Alberta to lie within the transition between the slower ( $\sim-8 \%$ ) North American Cordillera in the west and the faster ( $\sim 8 \%$ ) Canadian Shield region in the east. In addition, diminishing seismic velocities are noted in both of these investigations to the south of the $49^{\circ} \mathrm{N}$ parallel. Lithospheric conductivity structure across the Loverna Block and Proterozoic terranes to the northwest is constrained by the inversion of transverse magnetic mode magnetotelluric data [Boerner et al., 1999]. Mantle conductivity profiles appear to be dominantly 2-D and characterized by anomalously high values beneath the Loverna Block which have been interpreted to reflect tectonically induced metasomatism.

### 2.4 Author's Contribution

The research presented in Part I of this thesis has been submitted to the Canadian Journal of Earth Sciences in the form of a manuscript entitled Integrated teleseismic studies of the southern Alberta
upper mantle on behalf of myself as principal author and my co-authors Michael Bostock, Charly Bank, and Robert Ellis. Michael Bostock and Robert Ellis conceived, designed, and oversaw the operation of the field experiment. The data set used in the present analyses was collected by field technician Aaron Webb between July 1998 and June 1999. The data were processed by myself during the same time period with Charly Bank rendering invaluable technical assistance. All of the analysis and tectonic interpretation presented in Chapters 3 and 4 was carried out by myself, although co-authors have been involved, to varying degrees, in consultative roles.

## Chapter 3

## Teleseismic analysis: data, techniques, and results

Chapter overview: This chapter begins with a brief summary of some of experimental data aquisition parameters and the teleseismic event data set. This is followed by a discussion of the data processing for each of three analyses: i) travel time tomography; ii) shear wave splitting; and iii) radial component receiver functions.

### 3.1 Data Acquisition

The data considered in this experiment were recorded between July 1998 and June 1999 on a portable array of 9 broadband (i.e., $\sim 0.033-5.0 \mathrm{~Hz}$ frequency) seismographs and two permanent stations from the Canadian National Seismic Network (CNSN) at Waterton Lakes (WALA) and Edmonton (EDM), Alberta (see Figure 2.1). The array was oriented approximately perpendicular to the inferred basement strike with an interstation spacing between 40 km and 80 km . This aperture and station interval afford resolution in tomographic studies between 66 and 400 km depth in the mantle, and permit along-array profiling of upper mantle anisotropy, Moho depth, and effective Poisson's ratio of the crustal column. The array was well situated with respect to global seismicity; specifically, it falls within $100^{\circ}$ epicentral distance from subduction zones of the northwestern Pacific, Central and South America, and much of the mid-Atlantic ridge, and Alpine-Himalaya belt. The epicentres of the earthquakes used in the analyses are plotted in Figure 3.1, and a full catalogue of events is given in Appendix A.


Figure 3.1: Distribution of the sources used in the analyses of Part I of the thesis. Circles and stars represent events used in $P$-wave tomography and shear-wave splitting, respectively. Events employed in receiver function analysis are comprised of a subset from both categories.

### 3.2 Travel Time Inversion

### 3.2.1 Method

Teleseismic body-wave travel times were inverted to recover a smooth, three-dimensional (3-D) model of $P$-velocity perturbations in the upper mantle beneath the area of study. The tomographic
method of VanDecar [1991] was employed and can be summarized as follows. Optimum delay times are determined using a multi-channel cross-correlation technique that improves visual first $P$-arrival picks by a least-squares minimization of the inconsistency between cross-correlation derived relative delay times for all pairs of stations recording a given event [VanDecar and Crosson, 1990; VanDecár, 1991]. These travel-time residuals are subsequently inverted for velocity perturbations with respect to the iasp91 radial earth model [Kennett and Engdahl, 1991], beneath the array. Velocity perturbations are parameterized over a regular grid of knots every $1 / 3^{\circ}$ in latitude, $1 / 2^{\circ}$ in longitude, and 33 km in depth using splines under tension to smooth variations between knots. Robust linear inversion is performed using conjugate gradients (e.g., Hestenes and Stiefel, [1952]) combined with iterative downweighting of large residuals [Bostock and VanDecar, 1995] to simultaneously solve for slowness perturbations, station-time corrections (associated with e.g. topography), and event mislocation. Regularization is enforced by a damped least-squares procedure in which a combination of the first and second derivatives (flattening and smoothing) is minimized. Further information concerning the method is detailed in VanDecar [1991].

### 3.2.2 Data Set

The raw tomographic data set comprised 1565 visual travel time picks from 293 events with source magnitudes between $m_{b}=4.4$ and $m_{b}=6.6$. These times were measured primarily for direct $P$, although some core diffracted $P$-phases were also included. Cross-correlation derived relative delay times were characterized by a standard deviation of 0.023 s . The majority of events (162 out of 293) occurred in the northwestern Pacific and Latin American subduction zones, and lie within $15^{\circ}$ of the great circle passing through the array.

### 3.2.3 Results

To evaluate the resolution afforded by the data set, a "checkerboard" test was performed using a synthetic model in which every fifth knot in latitude and fourth knot in longitude at three depths ( $133 \mathrm{~km}, 300 \mathrm{~km}, 466 \mathrm{~km}$ ) was assigned an alternating $\pm 5 \%$ slowness perturbation. Synthetic travel time residuals for the source-receiver combinations represented in the real data set were then computed, corrupted by additive Gaussian noise with standard deviation $\sigma=0.1 \mathrm{~s}$, and used as input to the inversion procedure described above. Results of the resolution test are presented in Figure 3.2. The top 66 km of the recovered model are masked out due to the incorporation of crustal and shallow mantle structure within station correction terms [VanDecar, 1991]. Moreover, model regions characterized by poor ray coverage and resolution are also masked out. The important features of the model, shown in Figures $3.2 \mathrm{~b}, \mathrm{~d}, \mathrm{f}, \mathrm{h}$, may be summarized as follows: 1 ) anomalies are smeared across the array axis indicating that resolution is poorer across-axis than along-axis; ii) substantial vertical smearing exists along dominant ray directions especially near the ends of the array; and iii) recovered peak anomaly magnitudes underestimate the maximum slowness perturbations of the synthetic model. Overall, resolution in the upper $300-400 \mathrm{~km}$ is good, and we can expect to resolve any larger-scale mantle velocity variations underlying the major crustal domains.

Results from inversion of the real travel time data are presented in Figure 3.3 as a series of horizontal depth slices and a vertical profile A-A' through the preferred model. This model sits at the corner of the tradeoff curve between data misfit and model variance as constructed through a number of inversions involving varying combinations of candidate flattening and smoothing values. Iterative inversion of the data set led to a $66 \%$ reduction in the travel time residuals (r.m.s. reduction from 1.75 s to 0.59 s$)$. The recovered slowness anomalies display a range of $\pm 1.7 \%$ with


Figure 3.2: Results from a synthetic resolution test. Test involved three layers of alternating spikes with slowness perturbations of $\pm 5 \%$. a) and b) synthetic model and recovered structure for spike layer at 133 km depth; c) and d) as in panels a) and b) but at 300 km depth; e) and f) as in panels a) and b) but at 466 km depth; $g$ ) and h) depth cross section along receiver array between points A and $\mathrm{A}^{\prime}$.


Figure 3.3: Results from travel-time tomography. a) 100 km depth; b) 150 km depth; c) 200 km depth; d) 300 km depth; e) depth cross section along array between points A and A'. Domain names given in Figure 2.1. Receiver locations used in this study shown as white triangles.
respect to the $\operatorname{iasp} 91$ reference model. The most prominent feature imaged is the $\sim-1.7 \%$ slowness anomaly centred beneath the Loverna Block. This structure is truncated by slightly positive slowness perturbations to the south beneath parts of the Vulcan Structure and northern MHB and to the northwest where the lowest velocities are imaged below the Wabamun domain. Based on resolution tests, the localization of the largest perturbations within the top 300 km is probably robust and reflects lithospheric structure. Velocity anomalies at greater depths are, however, most certainly smeared.

### 3.3 Shear Wave Splitting

### 3.3.1 Method

Anisotropy in the continental mantle lithosphere is generally thought to develop through preferred orientation of olivine crystals in reponse to ductile deformation (e.g., Silver and Chan [1991]). The teleseismic $S K S$-phase affords a means of characterizing this anisotropy through the birefringence (or splitting) of $S$-waves in anisotropic material. Under the assumption of a single homogeneous layer of anisotropic mantle, the splitting of an incident $S K S$-wave into two orthogonally polarized waves may be parameterized by the polarization direction, $\phi$, of the fast wave and the delay time, $\delta t$, between the two split arrivals. Parameters $\phi$ and $\delta t$ are thereby representative of the orientation and degree of anisotropy, respectively.

The analysis method employed in the present study is adopted from Silver and Chan [1991] and is illustrated in Figure 3.4. Windows about the $S K S$-arrival on the horizontal (north and east) traces are selected (Figures $3.4 \mathrm{a}, \mathrm{b}$ ) and the waveforms are subsequently rotated into radial and transverse components (Figures 3.4c,d). These components are rotated through a trial angle $\phi$ (Figures $3.4 \mathrm{e}, \mathrm{f}$ ) and shifted by a time delay $\delta t$ (Figures $3.4 \mathrm{~g}, \mathrm{~h}$ ) in an attempt to correct for the


Figure 3.4: Example of shear wave splitting processing. Event 98/09/02 08:37.29 recorded at station AB04 (back azimuth 297.87, epicentral distance 104.19). a) north (solid line - SL) and east (dashed line - DL) component $S K S$ arrival picks; b) particle motion diagram of panel a); c) and d) as in panels a) and b) but rotated into radial (DL) and transverse (SL); e) and f) are components shown in panels c ) and d) rotated by trial $\phi$ into fast (DL) and slow (SL) components; g ) and h) are components shown in panels e) and f) shifted by trial time $\delta t$; i) and j) are the anisotropy corrected components shown in panels $g$ ) and $h$ ) rotated back into radial (SL) and transverse (DL) components; $k$ ) contour map representing minimization of transverse component energy of panels i) and $j$ ) as function of candidate $[\phi, \delta t]$ values. Global minimum shown as triangle; quoted error value outlined by enveloping nearest contour.
magnitude and direction of anisotropy manifest in the recorded signals. From these 'corrected' traces (Figures 3.4i,j) estimates of two functions (the energy on the transverse component and the second eigenvalue of the covariance matrix formed from the two split waves; see, e.g., Silver and Chan, [1991]) are obtained to serve as measures of the success of a particular $\phi, \delta t$ pair in accounting for anisotropy as manifest on the transverse component. This procedure is then repeated through iteration over a range of candidate $(\phi, \delta t)$ values to find the two-parameter pair which minimizes these energy measures (Figure 3.4k). This method has been extended through a simultaneous accommodation of multiple measurements to improve the accuracy of the splitting parameter estimates [Wolfe and Silver, 1998] and this extension has been adopted here.

### 3.3.2 Data Set

A total of 14 earthquake sources yielding 31 station-events were used in the $S K S$-splitting analysis. Contained within this suite of events were six $m_{b}>5.8$ earthquakes procured for stations EDM and WALA from the CNSN archives. Stations $\mathrm{AB} 02, \mathrm{AB} 03, \mathrm{AB} 06$, and AB 10 failed to record any $S K S$-phases at sufficient signal-to-noise levels to merit investigation. An examination of the epicentral distance coverage of the fourteen earthquakes used in the present analysis (shown as stars in Figure 3.1) indicates that the large majority of events originated from the western Pacific (back azimuth $\sim 230-300^{\circ}$ ), with a single suitable earthquake from South America (back azimuth $\sim 140^{\circ}$. The limited geographical sampling of these events does not permit evaluation of the dependence of splitting parameters on back azimuth; thus, it is necessarily assumed that lithospheric mantle anisotropy is homogeneous, with hexagonal symmetry and horizontal symmetry axis.

### 3.3.3 Results

Results from multi-event SKS splitting analysis are presented in Table 3.1 and Figure 3.5. The average value of $\phi$ over the entire array is $45 \pm 8^{\circ}$, with individual results ranging between $39 \pm 8^{\circ}$ and $55 \pm 8^{\circ}$. The fast direction at station AB 08 is rotated slightly clockwise of the relatively uniform northeast polarization direction calculated at other northern stations of the array, but this may reflect larger errors due to the small number of observations (2) at this station. Variations


Figure 3.5: Results (direction and magnitudes) of shear wave splitting analysis. Ovals and dotted line indicate estimated uncertainty ranges and absolute plate motion direction [Minster and Jordan, 1978], respectively. Domain names given in Figure 2.1.
in splitting magnitude recorded at the six northern reporting stations range moderately uniformly between $0.70 \pm 0.30 \mathrm{~s}$ and $1.05 \pm 0.38 \mathrm{~s}$ (average $0.82 \pm 0.30 \mathrm{~s}$ ). The remaining station, WALA, displays a slight increase in splitting delay time magnitude ( $1.20 \pm 0.30 \mathrm{~s}$ ) and a clockwise rotation in polarization angle $\left(\phi=53 \pm 7^{\circ}\right)$ possibly reflecting its proximity to Cordilleran deformation. Shear-wave splitting measurements for EDM and WALA have also been published by Bostock and Cassidy [1995] who obtain $33^{\circ}$ and 0.6 s , and $37^{\circ}$ and 0.9 s for these two stations, respectively (no uncertainty estimates given). Assuming uncertainties comparable to those of this study, these measurements are found to agree to within statistical error.

Table 3.1: Station locations and results of shear wave splitting and receiver function analyses.

| Stat | lat <br> ${ }^{\circ} \mathrm{N}$ | lon <br> ${ }^{\circ} \mathrm{W}$ | Split <br> data | $\phi$ <br> $\left[{ }^{\circ}\right]$ | $\delta t$ <br> $[\mathrm{~s}]$ | Moho <br> data | Moho depth <br> $[\mathrm{km}]$ | Poisson's Ratio |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| WALA | 49.06 | 113.91 | 8 | $53 \pm 7$ | $1.20 \pm 0.30$ | 34 | $50.0 \pm 2.0$ | $0.26 \pm 0.02$ |
| AB01 | 49.77 | 111.43 | 2 | $40 \pm 13$ | $0.80 \pm 0.23$ | 19 | $42.5 \pm 2.0$ | $0.29 \pm 0.02$ |
| AB02 | 50.34 | 111.73 | - | - | - | - | - | - |
| AB03 | 50.73 | 112.06 | - | - | - | 10 | $38.0 \pm 2.0$ | $0.30 \pm 0.03$ |
| AB04 | 51.10 | 112.68 | 4 | $42 \pm 7$ | $0.70 \pm 0.30$ | - | - | - |
| AB05 | 51.54 | 113.04 | 3 | $47 \pm 7$ | $0.80 \pm 0.28$ | 37 | $37.8 \pm 2.0$ | $0.30 \pm 0.02$ |
| AB06 | 52.05 | 113.44 | - | - | - | 3 | $38.0 \pm 3.0$ | $0.30 \pm 0.03$ |
| AB08 | 52.87 | 114.34 | 2 | $55 \pm 5$ | $0.75 \pm 0.30$ | 5 | $37.8 \pm 3.0$ | $0.28 \pm 0.03$ |
| EDM | 53.22 | 113.35 | 8 | $43 \pm 7$ | $1.05 \pm 0.38$ | 32 | $37.3 \pm 2.5$ | $0.28 \pm 0.03$ |
| AB10 | 53.81 | 115.21 | - | - | - | - | - | - |
| AB11 | 54.12 | 115.76 | 4 | $39 \pm 5$ | $0.80 \pm 0.30$ | 48 | $37.8 \pm 2.0$ | $0.29 \pm 0.02$ |

### 3.4 Receiver Function Analysis

### 3.4.1 Method

Receiver function analysis exploits information contained in the teleseismic $P$-wave coda to study discontinuities in the crust and upper mantle (e.g., Langston, [1979]). The approach adopted here is illustrated in Figure 3.6 and follows the method of Bostock [1998] which may summarized as follows. The three components of ground displacement are transformed into estimates of the $P$, $S V$, and $S H$ wavefields using estimates of near-surface $P$ - and $S$-velocities and the wavefront slowness. Wavefields from sources of similar epicentral distance and back azimuth (Figures 3.6a,b,c) are processed together in a simultaneous least-squares deconvolution of $S$-components by $P$ components to estimate the Earth's impulse response (Figure 3.6d).

The receiver functions can be further processed to yield estimates of crustal thickness and Poisson's ratio [Zhu and Kanamori, 2000]. Employing an average crustal $P$-wave velocity of 6.5 $\mathrm{km} \mathrm{s}^{-1}$, the binned receiver function amplitudes, at times corresponding to the direct (or Moho) conversion and the two first-order crustal reverberations (weighted by factors of $0.5,0.3$, and -0.2 respectively), are summed over a range of candidate Moho depths and crustal Poisson's ratios to locate the two-parameter combination which results in maximum coherence (Figure 3.6e). Uncertainty estimates in the present analysis are determined from the contour level of these stacks at $90 \%$ of the maximum value.

### 3.4.2 Data

The data chosen for receiver function analysis at each station consisted of up to 48 seismograms (see Table 3.1) with individual traces selected on the basis of signal-to-noise levels. These events


Figure 3.6: Examples of receiver function processing for station EDM. a) location of bins as a function of back azimuth (.) and epicentral distance ( + ); b) number of events per bin; c) corresponding locations of bins in panels a) and b); d) sample receiver functions with best fit locations of direct $P-S$ conversion and two first-order free-surface multiples (x); e) summed energy magnitude as function of candidate Poisson's ratio and Moho depth values with oval representing boundary of values falling with $90 \%$ of global maximum.
represent subsets of the suite of earthquakes employed in $P$-wave travel-time inversion and (or) shear-wave splitting analyses.

### 3.4.3 Results

Estimates of Moho depth and bulk crustal Poisson's ratio obtained from receiver function analysis for eight stations are presented in Table 3.1. Moho depths at the six northern reporting stations are fairly uniform exhibiting an average value of $\sim 38 \pm 2 \mathrm{~km}$. Stations AB01 and WALA, located in the MHB, display a deeper Moho signature at $42.5 \pm 2 \mathrm{~km}$ and $50 \pm 2 \mathrm{~km}$, respectively. Poisson's ratio estimates are also fairly uniform with an average value ( $0.28 \pm 0.02$ ). Station WALA again deviates from the regional trend with a result that is slightly lower than those to the north and east. Parameter values at both CNSN stations (EDM and WALA) have been checked against a much larger selection of events ( $>100$ ) from CNSN archives with similar results (Bank, Ph.D. thesis in preparation).

## Chapter 4

## Implications for southern Alberta's tectonic history

Chapter overview: This chapter discusses contraints on southern Alberta lithospheric structure and evolution that have arisen from the analyses presented in Chapter 3. In particular, aspects of the tectonic history of the Loverna and Medicine Hat Blocks, the origin of lithospheric anisotropy, and Moho depth and bulk crustal Poisson's estimates beneath individual stations are examined.

### 4.1 Discussion

The inversion of $P$-wave traveltime data from south-central Alberta has imaged significant lateral variations in mantle velocity structure. These variations appear to correlate with inferred locations of crustal basement domains, and are most prominent within the top $100-200 \mathrm{~km}$ of the mantle. In Figure 4.1, we plot the relative slowness perturbation at 133 km depth as an indication of velocity trends beneath the array. The lowest velocities (highest slownesses) are identified towards the northern (Wabamun Domain) and southern (MHB) ends of the array with highest velocities (lowest slownesses) located beneath the Loverna Block. Localization of the strongest velocity anomalies to the upper $\sim 200 \mathrm{~km}$ of the mantle is taken to reflect the preservation of fossil heterogenity in a relatively rigid lithosphere. Accordingly, the highest velocities beneath the Loverna Block are interpreted to be structurally significant, and representative of the thickest lithosphere documented along the profile.


Figure 4.1: Along-array profile of $P$-wave velocity model at 133 km depth. Documented profile contained within A-A' of Figure 3.3. Domain abbreviations as in Figure 2.1.

These observations present some problems for a recent model [Ross et al., 2000] of mantle evolution beneath the Hearne Province. In this model, the contemporaneity of two inwarddipping subduction zones, one involving young buoyant oceanic lithosphere to the northwest and the other continental lithosphere to the east, led to a "tectonic vise" in which a relatively compliant Hearne Province was compressed between two more rigid lithospheric blocks. This configuration enabled underthrusting-related mechanical interaction with the sub-Hearne lithosphere causing shortening and erosion of the mantle, and possibly fostering the growth of the continental root. Soon afterwards, the newly thickened lithosphere is postulated to have been thinned or removed via convective erosion or delamination leading to asthenospheric decompression and crustal melting arising from increased heat flow. This latter inference is partly based on mantle conductivity
profiles derived from magnetotelluric measurements which indicate substantially more conductive mantle beneath the Archean Hearne province than the Proterozoic domains to the northwest [Boerner et al., 1999; 2000]. The increased conductivity is thought to be due to the ingression of asthenospheric melts during decompression which may have prompted melting and related compositional modification of depleted mantle remnants.

On the basis of this tectonic model and, in particular, the inference of large-scale lithospheric removal, one might expect seismic velocities beneath the Hearne Province to be diminished with little internal variation. However, a comparison of Figure 3.3 with Figure 17 of Boerner et al. [2000] indicates a strong positive correlation between conductivity and velocity. Large-scale lithospheric removal is difficult to reconcile with this observation and it seems likely, therefore, that the agent responsible for enhanced conductivity has not affected the bulk elastic properties (and hence bulk composition or thickness) of the lithospheric mantle, or at least not in the manner generally expected.

The teleseismic results are therefore interpreted to indicate that mantle below the central Hearne Province was thickened but not delaminated. In particular, it is suggested that inwarddirected subduction from the northwest along the Snowbird Tectonic Zone, the east along the boundary with the Trans-Hudson Orogen, and possibly the south along the Vulcan Structure [Gorman, 2000; but see also Eaton et al., 1999] created the thickest lithosphere beneath the Loverna Block through successive imbrication of subducted lithospheres. Geophysical observations (e.g., Hyndman, [1988]) at modern subduction zones are often interpreted to indicate that dehydration above the slab plays an important role in modifying the overlying lithosphere. Accordingly, it is suggested that a volumetrically minor species (e.g., hydrous phases, Boerner et al., [1999]; dissolved hydrogen, Karato, [1990]; carbon, Roberts et al., [1999]) introduced through successive episodes of subduction has created a highly conductive Hearne lithosphere which, nonetheless,
is characterized by thick and high velocity lithosphere.

Another, possibly surprising, feature in the $P$-wave velocity model is the change from higher velocity beneath the Loverna Block to lower velocities beneath the northernmost MHB. This substantial variation ( $\sim 2 \%$ ) supports the conjecture of Eaton et al. [1999] that the two domains are genetically distinct and that the Vulcan Structure marks a structural suture (see also Gorman et al., [2001]). However, the observation that the MHB represents a long-standing stratigraphic high is more difficult to reconcile with the velocity model. In particular, it suggests that the MHB may be underlain by mantle that is on average less dense than the surroundings, resulting in a net buoyancy. The difficulty arises in identifying plausible mechanisms. A thermal contribution would agree with the observed velocity trend (i.e., higher temperature producing lower density and lower velocity) but is deemed unlikely given typical diffusion time scales and the timing of the last episode of tectonic activity. A second possibility, increased iron depletion which characterizes ancient Archean lithospheres [Jordan, 1988], cannot be appealed to because, although it results in decreased density of peridotite, it is also accompanied by increased elastic wave velocities. Rather, it may be that the differential motions are related to dynamic forcing of a more distant origin involving subduction [Pysklywec and Mitrovica, 2000]; however, in that case there must still be regional controls on short wavelength structure. Finally, it is worth noting that continentalscale tomographic investigations [van der Lee and Nolet, 1997; Frederiksen et al., 2001] also note a general decrease in mantle velocity southward from the Hearne Province to Wyoming Province.

The nature of lithospheric anisotropy, as inferred from the analysis of $S K S$, is thought to be closely related to an imprint of the most recent episode of tectonism to have affected a region [Silver and Chan, 1991]. Depth localization of anisotropy is poorly constrained since the near vertical $S K S$ raypath provides only a vertical averaging of mantle fabric. One must therefore also consider the possible influence of plate motion and asthenospheric flow on the observables $\phi$ and
$\delta t$ [Vinnik et al., 1995; Fouch et al., 2000]. Analysis of SKS waves in the present experiment has revealed a fairly uniform set of anisotropy splitting parameters. The average polarization direction corresponds approximately to both the normal to an average northwest-directed shortening in the Hearne Province [Ross et al., 2000], and to the direction of North American absolute plate motion ( $\phi_{a p m} \sim 54^{\circ}$, [Minster and Jordan, 1978]; the dashed line in Figure 3.5). This coincidence, as in many parts of Precambrian North America (see, e.g., Silver and Chan, [1991]), renders it difficult to separate relative contributions from fossil anisotropy in the lithosphere and flow in the underlying asthenosphere. However, a lack of systematic variation in splitting parameters that might correspond to the clockwise rotation in geological strike of the basement from north to south (see Figure 3.5) suggests that at least some component of the splitting signal originates in the asthenosphere.

The results from receiver function analysis in the present experiment reveal north-to-south variations in crustal thickness. Stations at the northern end of the array exhibit relatively uniform depths to the Moho of $\sim 38 \mathrm{~km}$, whereas stations within the MHB are located over significantly thicker crust ( $>40 \mathrm{~km}$ ). Although teleseismic stations and the aforementioned reflection (SALT, CAT) and refraction (SAREX) surveys are offset at points by distances in excess of 100 km (see Figure 2.1), comparisons drawn between interpreted profiles are still informative. Generally, Moho depths retrieved in the present study vary consistently with those obtained from active-source surveys. For example, station AB01 may be compared to SALT line 31 ( 43 km vs $\sim 47 \mathrm{~km}$ ) and SAREX ( 43 km vs $\sim 48 \mathrm{~km}$ ); station WALA to SALT line $30(50 \mathrm{~km}$ vs $\sim 51 \mathrm{~km}$ ); station AB05 to SAREX ( 38 km vs $\sim 45 \mathrm{~km}$ ); and station AB06 to CAT line 7 ( 38 km vs $\sim 38 \mathrm{~km}$ ). Discrepancies between Moho depth estimates may result from some combination of two factors. First, as illustrated by Clowes et al. [2001] in their comparison of SAREX refraction and SALT
(line 21) reflection profiles, significant variations in depth to the Moho exist along strike. Second, Moho depth estimates retrieved by the aforementioned techniques may be biased differently according to frequency bandwidth and nature of discontinuity (e.g., Zandt and Owens, [1986]).

Part II

INVERSION/MIGRATION OF SCATTERED TELESEISMIC BODY-WAVES: NUMERICAL MODELLING

## Chapter 5

## Motivation for Part II

Part II overview: Advances in our understanding of the continental lithosphere and underlying upper mantle have relied heavily on studies involving multichannel processing of teleseismic body wave data, most notably $P$-wave travel time tomography (e.g., VanDecar, [1991]). Recent investigations (e.g., Revenaugh, [1995]; Bostock, [1998]; Ryberg and Weber, [2000]), however, indicate that further insight into detailed lithospheric structure is to be obtained through multichannel analysis of secondary scattered waves in the $P$-coda. The research presented in Part II of the thesis investigates a theoretically rigorous yet practically implementable solution to the inverse scattering/migration problem in earthquake seismology.

### 5.1 Overview of Method

The theoretical framework for inverse scattering/migration of teleseismic waves exploited in Part II of this thesis was developed in Bostock et al. [2000; hereafter referred to as Paper $\Pi$ ], which is an extension of the earlier approach of Bostock and Rondenay [1999; hereafter referred to as $B R$ ]. The issue addressed in both of these works is formal, yet practical, inversion of scattered teleseismic body wave coda recorded on dense surface arrays for underlying 2-D lithospheric structure using inverse-scattering theory (see Figure 5.1). In particular, the approach is amenable to: i) irregular receiver sampling; ii) independent appraisal of contributions from individual scattering mode interactions including free-surface multiples; and iii) simultaneous treatment of multiple


Figure 5.1: Plan view of a quasi-linear array of receivers (inverted triangles) above a 2-D heterogeneous lithosphere illuminated by an incident teleseismic wavefield. The incident wavefront is assumed to be planar at the scale of the array, with horizontal slowness $\mathbf{p}_{\perp}^{0}=\left(p_{1}^{0}, p_{2}\right)$, and conservation of the $x_{2}$ slowness component $p_{2}$ for all derived waves as required by Snell's law. (source: Paper I)
events from arbitrary back azimuths. Below I provide a necessarily brief outline of essential steps involved in the derivation of the inversion formula. Readers interested in more complete theoretical development are directed to Paper I.

The inversion method is based on a high-frequency, single-scattering formulation of the forward-scattering problem in which the frequency domain scattered displacement field, $\Delta u_{n}^{q}\left(\mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}, \omega\right)$, from a planar (in horizontal aspect) incident wave for a particular scattering mode interaction, $q$ (see Figure 5.2), and direction component, $n$, can be expressed (c.f. equation (26)
a)


b)

$\begin{aligned} & P \\ & \\ & \rightarrow q=4\end{aligned}$

c)



Figure 5.2: Schematic diagrams illustrating the scattering modes employed in Part II: a) forward-scattering of incident $P$-wave into $P(q=1)$ and $S(q=2)$ waves; b) back-scattering of free-surface-reflected $P$-wave into $P(q=3)$ and $S(q=4)$ waves; and c) back-scattering of free-surface-reflected $S$-wave into $P(q=5), S(V)(q=6)$, and $S(H)(q=7)$ waves. (source: Rondenay, [2000])
in Paper I) as,

$$
\begin{equation*}
\Delta u_{n}^{q}\left(\mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}, \omega\right)=K(\omega) \int \mathrm{d} \mathbf{x} f^{q}(\mathbf{x}, \theta) \mathcal{A}_{n}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right) \mathrm{e}^{\mathrm{i} \omega \mathcal{T}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)} \tag{5.1}
\end{equation*}
$$

where 2-D spatial variables $\mathbf{x}, \mathbf{x}^{\prime}$ represent scatterer and receiver locations, respectively, and lie within the $x_{1}, x_{3}$ plane (i.e., 2-D strike parallels $x_{2}$-coordinate, with $x_{1}$ orthogonal in horizontal plane and $x_{3}$ vertical, positive downward; see Figure 5.1); and $\mathbf{p}_{\perp}^{0}=\left(p_{1}^{0}, p_{2}\right)$ is the horizontal slowness of the incident wave. A frequency factor $K(\omega)=\omega^{2} / \sqrt{-\mathrm{i} \omega}$ arises from the planar nature of the incident wavefield and the 2-D experimental geometry. The scattering potential,

$$
\begin{equation*}
f^{q}(\mathbf{x}, \theta)=\sum_{l=1}^{3} W_{l}^{q}(\theta) \Delta m_{l}(\mathbf{x}) \tag{5.2}
\end{equation*}
$$

includes the scattering angle $\theta$ dependent radiation pattern coefficients $W_{l}^{q}(\theta)$, and material property perturbations $\Delta m_{l}=[\Delta \alpha / \alpha, \Delta \beta / \beta, \Delta \rho / \rho]$, where $\alpha, \beta$, and $\rho$ are the $P$-wave and $S$-wave
velocities, and density of the reference medium, respectively. Full expressions for $f^{q}(\mathbf{x}, \theta)$ are given in Appendix B. Quantity $\mathcal{T}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)$ is the arrival time at receiver $\mathbf{x}^{\prime}$ of energy scattered from model point $\mathbf{x}$, and $\mathcal{A}_{n}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)$ represents the product of geometrical amplitudes of the incident and scattered waves. That is, for $q=1$, for example,

$$
\begin{equation*}
\mathcal{A}_{n}^{1}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)=A^{0}\left(x_{3}\right) A^{P}\left(x_{3} ; x_{3}^{\prime}, \mathbf{p}_{\perp}^{0}\right) \hat{x}_{n}^{P}\left(x_{3}^{\prime}, \mathbf{p}_{\perp}^{0}\right) \tag{5.3}
\end{equation*}
$$

where $A^{0}\left(x_{3}\right)$ is the incident wavefield amplitude, and $A^{P}\left(x_{3} ; x_{3}^{\prime} ; \mathbf{p}_{\perp}^{0}\right)$ and $\hat{x}_{n}^{P}\left(x_{3}^{\prime}, \mathbf{p}_{\perp}^{0}\right)$ are the amplitude and source polarization of the scattered $P$-wave Green's function, respectively.

At this point in the derivation, it proves expedient to introduce a new, filtered time-series, $v_{n}^{q}\left(\mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}, t\right)$, through multiplication of both sides of equation (5.1) by factor $-\mathrm{i} \operatorname{sgn}(\omega) / \sqrt{-\mathrm{i} \omega}$ and a subsequent Fourier transform,

$$
\begin{align*}
& v_{n}^{q}\left(\mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}, t\right)=-\frac{1}{2 \pi} \int \mathrm{~d} \omega \mathrm{e}^{\mathrm{i} \omega t} \Delta u_{n}^{q}\left(\mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}, \omega\right) \mathrm{i} \operatorname{sgn}(\omega) / \sqrt{-\mathrm{i} \omega} \\
& \quad=-\int \mathrm{dx} f^{q}(\mathbf{x}, \theta) \mathcal{A}_{n}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right) \frac{1}{2 \pi} \int \mathrm{~d} \omega \mathrm{i} \omega \mathrm{e}^{\mathrm{i} \omega\left[t-\mathcal{T}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)\right]_{\mathrm{i} \operatorname{sgn}}(\omega)}  \tag{5.4}\\
& \quad=-\int \mathrm{dx} f^{q}(\mathbf{x}, \theta) \mathcal{A}_{n}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right) \mathcal{H}\left\{\delta^{\prime}\left[t-\mathcal{T}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)\right]\right\},
\end{align*}
$$

where $\mathcal{H}\{\cdot\}$ denotes a Hilbert transform. Note that $\delta^{\prime}\left[t-\mathcal{T}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)\right]$ is the derivative of the singular function for the isochronal curve which defines the locus of all model points $\mathbf{x}$ from which scattered energy arrives simultaneously at receiver $\mathbf{x}^{\prime}$ (See Figure 5.3).

Assuming that the scattering body is localized to the vicinity of a point $\mathbf{x}_{0}$, and that the amplitude function $\mathcal{A}_{n}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)$ may be approximated by $\mathcal{A}_{n}^{q}\left(\mathbf{x}_{0}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)$ (i.e., that $\mathcal{A}_{n}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)$ is a slowly varying function for $\mathbf{x}_{0}$ near $\mathbf{x}$; see Miller et al., [1987]), equation (5.4) may be rewritten as,

$$
\begin{equation*}
v_{n}^{q}\left(\mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}, t\right) \approx \frac{\mathcal{A}_{n}^{q}\left(\mathbf{x}_{0}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)}{\left|\nabla \mathcal{T}_{0}^{q}\right|^{2}} \int \mathrm{dx} f^{q}(\mathbf{x}, \theta) \mathcal{H}\left\{\delta^{\prime}\left[\mathbf{n} \cdot\left(\mathbf{x}-\mathbf{x}_{0}\right)\right]\right\} \tag{5.5}
\end{equation*}
$$



Figure 5.3: Geometrical quantities considered for scattered teleseismic waveform inversion. All quantities are projected onto the $x_{1}-x_{3}$ plane, and represented with solid lines when strictly confined to this plane, or with dashed lines when they have a non-zero component in the $x_{2}$ direction. (source: Paper I)
where $\mathcal{T}_{0}^{q}=\mathcal{T}_{0}^{q}\left(\mathbf{x}_{0}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)$ is the travel time for a scattered wave from point $\mathbf{x}_{0}$ to receiver $\mathbf{x}^{\prime}, \nabla \mathcal{T}_{0}^{q}$ is the spatial gradient of the total travel time function evaluated at $\mathbf{x}_{0}$, and $\mathbf{n}=\mathcal{T}_{0}^{q} /\left|\nabla \mathcal{T}_{0}^{q}\right|^{2}$ is a unit vector defining the direction of the gradient of the total travel time function (see Figure 5.3)

Contracting both sides of equation (5.5) with $\mathcal{A}_{n}^{q}\left(\mathbf{x}_{0}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)$, rearranging terms, and integrating over the full range of $\psi$ (the angle defined by the dot product of unit vector $\mathbf{n}$ with the $\mathbf{x}_{3}$ axis)
yields,

$$
-\frac{1}{4 \pi} \int_{0}^{2 \pi} \mathrm{~d} \psi \int \mathrm{~d} \mathbf{x} f^{q}(\mathbf{x}, \theta) \mathcal{H}\left\{\delta^{\prime}\left[\mathbf{n} \cdot\left(\mathbf{x}_{0}-\mathbf{x}\right)\right]\right\}
$$

$$
\begin{equation*}
=\frac{1}{4 \pi} \int \mathrm{~d} \psi \frac{\left|\nabla \mathcal{T}_{0}^{q}\right|^{2}}{\left|\mathcal{A}^{q}\right|^{2}} \sum_{n} \mathcal{A}_{n}^{q}\left(\mathbf{x}_{o}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right) v_{n}^{q}\left(\mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}, t=\mathcal{T}_{0}^{q}\right) \tag{5.6}
\end{equation*}
$$

where $\left|\mathcal{A}^{q}\right|^{2}=\sum_{n} \mathcal{A}_{n}^{q}\left(\mathbf{x}_{0}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right) \mathcal{A}_{n}^{q}\left(\mathbf{x}_{0}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right)$.
Equation (5.6) is reminiscent of the definition of the inverse Radon transform [Beylkin, 1985] in two dimensions. The 2-D Radon transform pair is given by,

$$
\begin{align*}
& F(\mathbf{n}, s)=\int \mathrm{d} \mathbf{x} f(\mathbf{x}) \delta(s-\mathbf{n} \cdot \mathbf{x})  \tag{5.7}\\
& \begin{aligned}
f\left(\mathbf{x}_{0}\right) & =-\frac{1}{4 \pi} \int \mathrm{~d} \mathbf{n} \mathcal{H}\left[\left.\frac{\partial}{\partial s} F(\mathbf{n}, s)\right|_{s=\mathbf{n} \cdot \mathbf{x}_{0}}\right] \\
= & -\frac{1}{4 \pi} \int \mathrm{~d} \mathbf{n} \int \mathrm{~d} \mathbf{x} f(\mathbf{x}) \mathcal{H}\left\{\delta^{\prime}\left[\mathbf{n} \cdot\left(\mathbf{x}_{0}-\mathbf{x}\right)\right]\right\} \\
\quad= & -\frac{1}{4 \pi} \int_{0}^{2 \pi} \mathrm{~d} \psi \int \mathrm{~d} \mathbf{x} f(\mathbf{x}) \mathcal{H}\left\{\delta^{\prime}\left[\mathbf{n} \cdot\left(\mathbf{x}-\mathbf{x}_{0}\right)\right]\right\}
\end{aligned}
\end{align*}
$$

The definition of the inverse Radon transform in equation (5.8) indicates that function $f(\mathbf{x})$ (in our case the scattering potential at constant $\theta$ ) may be reconstructed at any point $\mathbf{x}_{0}$ by summing function $F(\mathbf{n}, s)$ over all planes passing through point $\mathbf{x}_{0}$. Unlike the Radon transform, however, the surfaces represented in the present derivation are curvilinear (e.g., the isochrons illustrated in Figure 5.3) but these may treated as locally planar in the vicinity of point $\mathbf{x}_{0}$ (see Miller et al., [1987]). This identification allows for the definition of a back projection operator that reconstructs the $\theta$-dependent scattering potential, $f^{q}(\mathbf{x}, \theta)$, as,

$$
\begin{equation*}
f^{q}\left(\mathbf{x}_{0}, \dot{\theta}\right)=\frac{1}{4 \pi} \int \mathrm{~d} \psi \frac{\left|\nabla \mathcal{T}_{0}^{q}\right|^{2}}{\left|\mathcal{A}^{q}\right|^{2}} \sum_{n} \mathcal{A}_{n}^{q}\left(\mathbf{x}_{0}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right) v_{n}^{q}\left(\mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}, t=\mathcal{T}_{0}^{q}\right) \tag{5.9}
\end{equation*}
$$

Using the definition of $f^{q}\left(\mathbf{x}_{0}, \theta\right)$ in equation (5.2), and recognizing that the number of measurements over variable $\theta$ yield an overdetermined linear system at model point $\mathbf{x}_{0}$, the angular dependence $\theta$ may be treated in a least-squares sense thus allowing for the recovery of material property perturbations, $\Delta \mathrm{m}$, as

$$
\begin{equation*}
\Delta \mathrm{m}=\mathrm{H}^{-1} \mathrm{~g} \tag{5.10}
\end{equation*}
$$

The elements of $g$ are given by,

$$
\begin{align*}
& g^{q}\left(\mathbf{x}_{0}\right)=\frac{1}{4 \pi} \int \mathrm{~d}\left|\mathbf{p}_{\perp}^{0}\right| \int \mathrm{d} \theta \int \mathrm{~d} \psi \sum_{k} W_{k}^{q}(\theta) \frac{\left|\nabla \mathcal{T}_{0}^{q}\right|^{2}}{\left|\mathcal{A}^{q}\right|^{2}} \sum_{n} \mathcal{A}_{n}^{q}\left(\mathbf{x}_{0}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right) v_{n}^{q}\left(\mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}, t=\mathcal{T}_{0}^{q}\right)  \tag{5.11}\\
& \quad=\frac{1}{4 \pi} \int \mathrm{~d}\left|\mathbf{p}_{\perp}^{0}\right| \int \mathrm{d} \gamma \int \mathrm{~d} x_{1}^{\prime}\left|\frac{\partial(\psi, \theta)}{\partial\left(x_{1}^{\prime}, \gamma\right)}\right| \sum_{k} W_{k}^{q}(\theta) \frac{\left|\nabla \mathcal{T}_{0}^{q}\right|^{2}}{\left|\mathcal{A}^{q}\right|^{2}} \sum_{n} \mathcal{A}_{n}^{q}\left(\mathbf{x}_{0}, \mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}\right) v_{n}^{q}\left(\mathbf{x}^{\prime}, \mathbf{p}_{\perp}^{0}, t=\mathcal{T}_{0}^{q}\right)
\end{align*}
$$

where the Jacobian transformation has been introduced in equation (5.11) to transform the integration over geometrical quantities $\psi, \theta$ (see Figure 5.3) into source, $\gamma$, and receiver, $x_{1}^{\prime}$, variables. The elements of the Hessian matrix H are given by,

$$
\begin{equation*}
H_{l q}\left(\mathbf{x}_{0}\right)=\int \mathrm{d}\left|\mathbf{p}_{\perp}^{0}\right| \int \mathrm{d} \theta \sum_{k} W_{l}^{k}\left(\theta,\left|\mathbf{p}_{\perp}^{0}\right|, \mathbf{x}_{0}\right) W_{k}^{q}\left(\theta,\left|\mathbf{p}_{\perp}^{0}\right|, \mathbf{x}_{0}\right) . \tag{5.12}
\end{equation*}
$$

The inverse problem can, accordingly, be viewed simply as a weighted diffraction stack over all sources and receivers with weights determined through analogy with the Radon transform. That is, the scattering potential, $g^{q}\left(\mathbf{x}_{0}\right)$, defined in equation (5.11) is constructed at any given model point $\mathbf{x}_{0}$ by an event-by-event weighted summation of energy along travel time curves corresponding to the arrival time of the scattered at the surface from point $\mathbf{x}_{\mathbf{0}}$. This function is then employed in equation (5.10) to recover the perturbations in material properties, $\Delta \mathrm{m}$, through a trivial $3 \times 3$ matrix multiplication.

### 5.2 Research Goal

The objective of Part II of this thesis is to investigate the potential of this inversion methodology through a variety of numerical simulations. Specifically, its main contribution is to evaluate the performance of the algorithm in a controlled, idealized environment and to provide insight into limitations and potential problems that may arise during inversion of field data (i.e., Rondenay et al., [2000]). To this end, synthetic 2-D simulations have been designed which comprise interaction of upward-propagating plane waves with a model lithospheric suture zone. This class of
structure is a typical target in lithospheric-scale geophysical surveys (e.g., Calvert et al., [1995]) and one which produces a response sufficiently complex that conventional teleseismic processing (i.e., radial component receiver functions) approaches may yield results which are not straightforward to interpret. Details of the subduction zone model used in numerical simulations, including the generation of synthetic data from this model and its subsequent preprocessing, are discussed in Chapter 6. In the following chapter, the synthetic data set is utilized in a series of inversions to identify the importance of the roles that various parameters play in the recovery of structure.

### 5.3 Author's Contribution

The research contained in Part II of this thesis has been accepted pending minor revisions in the Journal of Geophysical Research as the second installment of a tri-partite series (i.e., Paper I; Shragge et al., [2000], hereafter referred to as Paper II; and Rondenay et al., [2000], hereafter referred to as Paper III) that develops, tests, and applies this new teleseismic imaging technique. This material has also been presented in abbreviated form in a series of posters and oral communications at scientific meetings. Chapter 5 comprises a summation of theory outlined in Paper I, and is similar to the overview presented in the thesis of my co-author Stéphane Rondenay [Rondenay, 2000]. Chapters 6 and 7 are based in major part on Paper II. The research in all of Papers I, II, and III was conducted in close collaboration between myself, Michael Bostock, and Stéphane Rondenay. While the principal author of each paper was responsible for the main contribution, co-authors were actively involved in most aspects of the research. In particular, I assisted in the writing of the computer code dealing with inversion of multi-event in-plane sources, generated and developed preprocessing procedures for synthetic data, and devised, conducted, and analy-: sized results from numerical simulations.

## Chapter 6

## Generation of synthetic data

Chapter overview: This chapter begins with a brief summary of documented lithospheric sutures as imaged in seismic reflection surveys and which have motivated the design of our idealized numerical model. The forward modelling of teleseismic $P$-wave propagation through this lithospheric model using a pseudo-spectral numerical method is then described, and followed by an outline of the preprocessing required to prepare raw displacement seismograms for inversion.

### 6.1 Lithospheric Suture Model

Lithospheric suture zones are the signatures of collisions between continents which generally culminate extended periods of ocean-continent subduction. As such, they represent one of the more complex structures likely to be encountered in studies of the continental lithosphere. The evolution of lithospheric suture zones arising from smaller-scale, convergent boundaries (i.e., less than $\sim 150 \mathrm{~km}$ in horizontal extent) has been investigated using 2-D, finite-element modelling [Beaumont and Quinlan, 1994; Fullsack, 1995] for a range of crustal models differing in geothermal gradient, composition and structural geometry. A key factor in the development of lithospheric sutures is the location of the brittle-ductile transition within the crustal column which divides the crust rheologically into an overlying viscoplastic layer and an underlying, more competent lithospheric mantle (e.g., Fullsack, [1995]). The lithospheric detachment point associated with
this boundary governs the location where lithospheric materials partition into obducting and subducting segments. In hotter thermal regimes where the brittle-ductile transition occurs nearer to surface, there is a corresponding increase in the volume of lithospheric material subducted into the mantle. Thus, the thickness of a subducting segment is thought to be strongly dependent on the vertical extent of crustal column below the brittle-ductile transition. The situations modelled by Beaumont and Quinlan [1994] for which transitions occurred in the mid-crust resulted in subducted segment thicknesses between $10-20 \mathrm{~km}$.

Suture zones have been documented in a number of high-resolution seismic reflection profiles traversing both Phanerozoic and Precambrian orogens (e.g., Pfiffner et al., [1990]; Calvert et al., [1995]; Cook et al., [1998]). These studies provide the principal constraints on the deepseated geometry of suture zones through: i) reflections from the Moho which often indicate crustal thickening followed by abrupt thinning; ii) the presence of dipping mantle reflections that document penetration of subducted crustal segments, in some cases, to depths of $\sim 100 \mathrm{~km}$ or more (e.g., Cook et al., [1998]); and iii) reflection amplitudes which suggest that subducted material has undergone dehydration reactions (e.g., Calvert et al., [1995]). Although seismic reflection profiles provide the most detailed views of lithospheric sutures, they are limited in at least two respects. First, the use of anthropogenic sources limits depth penetration to uppermost mantle levels. Second, the narrow and relatively high signal bandwidth of conventional seismic sources often precludes accurate determination of signal polarity and, hence, sign of impedance contrasts. Densely sampled teleseismic profiles would not suffer from these limitations and, accordingly, could provide important complementary information on complex lithospheric structures.

Our idealized lithospheric suture model, shown in Figure 6.1 and loosely based on the aforementioned studies, is defined over a $360 \times 120 \mathrm{~km}^{2}$ section and consists of three materials with differing elastic properties. A low-velocity crustal layer overlies a faster upper mantle (see Table


Figure 6.1: Idealized lithospheric subduction-suture model. Velocities and densities of each material are given in Table 6.1.
6.1 for model velocities and densities). At the location of the suture, crustal material from the lithospheric block to the left bifurcates, with the lower segment descending into the mantle. At a depth of $\sim 40 \mathrm{~km}$, it converts to velocities and density higher then the surrounding mantle (note the proportionally greater increase in $S$-velocity) and thereafter folds and thins to the right of the model. Heterogeneity comprises both dipping, locally planar interfaces (i.e., the slowly varying Moho, subducting segments) from which specular conversions and reflections are expected, and sharper, angular discontinuities (i.e., mantle wedge, slab fold) which will produce more complex scattering, loosely referred to here as diffraction. Thus, the model should afford a reasonable test of the algorithm's ability to image a range of structural complexity.

### 6.2 Generation of Synthetics

Several data sets of two-component seismograms were computed through the lithospheric model described above using a 2-D, elastic pseudo-spectral code [Kosloff et al., 1990]. The model was

Table 6.1: Lithospheric model parameters.

| Parameter | P-wave velocity $\left(\mathrm{km} \mathrm{s}^{-1}\right)$ | S-wave velocity $\left(\mathrm{km} \mathrm{s}^{-1}\right)$ | Density $\left(\mathrm{g} \mathrm{cm}^{-3}\right)$ |
| :---: | :---: | :---: | :---: |
| Crust | 6.2 | 3.6 | 2.8 |
| Mantle | 8.0 | 4.5 | 3.2 |
| Relict Crust | 8.1 | 4.9 | 3.3 |

discretized at 0.6 km intervals in both directions which, given the minimum velocity in our model ( $3.6 \mathrm{~km} \mathrm{~s}^{-1}$ ), allows for frequencies up to 4.0 Hz to be accurately modelled. This value corresponds approximately to the upper frequency limit of $P$-wave seismograms from teleseismic events.

The data sets comprise a suite of plane $P$-waves interacting with the suture model at a range of incident horizontal slownesses, $p_{1}^{0}=[0.05,-0.05,0.06,-0.06,0.07,-0.07] \mathrm{s} \mathrm{km}^{-1}$. A Gaussian waveform, $\mathrm{e}^{-18 t^{2}}$ ( $t$ measured in seconds), was chosen to approximate the incident delta-function pulse (i.e., Green's function) required in the inverse formulation. The planar wavefront is a close approximation to that of a teleseismic $P$-wavefront predicted from spherical Earth models where the change in horizontal slowness over a 300 km interval at the Earth's surface will not generally exceed a few percent. The output seismogram sections consist of 600 traces recorded at the free surface. These sections are subsequently desampled by a factor of five to yield data sections of 120 traces sampled at 3 km intervals as input for inversion. To allow incorporation of the freesurface back-scattered response, the duration spanned by individual records is approximately 50 s after the direct arrival.

### 6.3 Preprocessing and Synthetic Results

Implementation of the method in Paper I requires that the so-called scattered wavefield $\Delta \mathbf{u}$ is effectively isolated (see equations (24-25,43-45), Paper $I$ ). The technique employed here is that of $B R$ who detail a procedure through which the incident wavefield $\mathbf{u}^{0}$ (i.e., the $P$-wavefield which would propagate in a smoothly varying 1-D reference medium) is approximately removed from a multichannel seismogram section to yield an estimate of the scattered wavefield. The method may be summarized as follows: i) the raw data sections, $\mathbf{u}=\left[u_{1}, u_{2}\right]$, are transformed into upgoing $P$ - and $S$-wavefield sections, $\mathbf{w}=[P, S V]$, via the free-surface transfer matrix [Kennett, 1991]; ii) multi-channel cross-correlation [VanDecar and Crosson, 1990] is applied to a window about the direct (high-pass filtered) $P$-arrival to allow optimal alignment of the wavefield sections; iii) the aligned $P$-section is decomposed into its principal components through diagonalization of the zero-lag, cross-correlation matrix; iv) the first (or first few) principal component(s) are identified with the source time function of the incident wavefield while the remaining principal components (or some selection thereof) are associated with the scattered wavefield; v) the scattered displacement sections are reconstituted from the $P$ - and $S$ - wavefield sections using the inverse free-surface transfer matrix; vi) in practice, individual source time function estimates (as obtained by the above procedure) are then deconvolved from the scattered displacement to yield the scattered wavefield $\Delta \mathbf{u}$; however, this step is not employed here as the Gaussian source-time function is essentially a low-pass filtered approximation of the delta-function-like impulse-response required by theory. The final stage of preprocessing entails a convolution of $\Delta u$ with a filter whose frequency domain representation is $F(\omega)=-\mathrm{i} \operatorname{sgn}(\omega) / \sqrt{-\mathrm{i} \omega}$, to produce a new time series v (equation (5.4)) as required by the 2-D plane-wave migration scheme to directly image jumps in elastic properties across discontinuities. Each trace is then subjected to a bandpass Butterworth
filter between 0.1 Hz and 4.0 Hz to approximate the typical frequency bandwidth from a deep teleseismic event.

Two-component, processed (but prior to application of $F(\omega)$ ) scattered displacement waveforms (i.e., $\Delta \mathbf{u}$ ) for incident horizontal slownesses, $p_{1}^{0}$, of -0.07 and $0.07 \mathrm{~s} \mathrm{~km}^{-1}$ are shown in Figures 6.2a-d. Individual scattered phases apparent in Figures $6.2 \mathrm{a}-\mathrm{d}$ include both forwardscattered waves ( $q=1,2$; equations $(24,25$ ) in Paper $I$ ), and back-scattered modes ( $q=3,4,5,6,7$; equations (24,25,43-45) in Paper I) afforded through free-surface reflection of the incident upgoing $P$-wavefield into downgoing $P$ - and $S$-waves. To summarize (see also Figure 5.2) $q=1,3$ represent $P$ - $P$ scattering; $q=2,4$ represent $P-S$ scattering; $q=5$ represents $S$ - $P$ scattering; and $q=6,7$ represent different modes of $S$-S scattering (only one $S-S$ scattering mode, $q=6$, is required for the in-plane, 2-D geometry of the present problem). The horizontal component response for right incidence, Figure 6.2a, is dominated by a $P-S$ conversion ( $q=2$ ) from the Moho arriving at $\sim 5 \mathrm{~s}$ and a $P-P$ diffraction $(q=1)$ centered at 100 km with apex at $\sim 0 \mathrm{~s}$. Subsequent arrivals are most obvious on seismograms to the right of the suture: i) a combination of $P-S(q=2)$ scattering originating from either side of the subducted crustal segment at $\sim 10-11 \mathrm{~s} ;$ ii) a back-scattered $P-S$ conversion ( $q=4$ ) from the Moho at $\sim 17 \mathrm{~s}$; and iii) a $S$-S reflection ( $q=6$ ) from the Moho at $\sim 22 \mathrm{~s}$. The vertical component, Figure 6.2b, emphasizes two further arrivals: i) a $P-P$ phase reflection ( $q=4$ ) from the Moho at $\sim 11 \mathrm{~s}$; and ii) a weaker $S-P$ conversion $(q=5)$ from the Moho at $\sim 17 \mathrm{~s}$. These phases represent the dominant contributions to scattering; several other "kinematic analogues" involving interaction with the free surface may also be identified but these represent effectively second order contributions.

Figures $6.2 \mathrm{c}, \mathrm{d}$ present the corresponding displacement components for a plane wave incident from the left. The amplitudes of converted phases $(q=2,4,5)$ from the near-planar horizons are diminished with respect to Figures $6.2 \mathrm{a}, \mathrm{b}$ because the specular angles are closer to perpendicular
(and hence the conversion coefficients closer to zero) while diffracted signals from the suture are more pronounced. A series of diffractions, all centered near $\sim 150 \mathrm{~km}$ in offset, are apparent: i) a $P-P$ diffraction $(q=1)$ with apex at $\sim 0 \mathrm{~s}$ (Figure 6.2d); ii) two $P-S$ diffractions $(q=2)$ with apices at $\sim 5 \mathrm{~s}$ and $\sim 7 \mathrm{~s}$ (Figure 6.2c); iii) a $P-P$ diffraction ( $q=3$ ) with apex at $\sim 10 \mathrm{~s}$ (Figure 6.2d); and iv) a $S-P$ diffraction ( $q=5$ ) with apex at $\sim 16 \mathrm{~s}$ (Figure 6.2 d ).


Figure 6.2: Preprocessed (but prior to filtering by $F(\omega)$ ) synthetic data sections generated for the idealized lithospheric model in Figure 6.1. a) $\Delta u_{1}$ and b) $\Delta u_{3}$ components for an incident plane wave with horizontal slowness $p_{1}^{0}=-0.07 \mathrm{~s} \mathrm{~km}^{-1}$; c) $\Delta u_{1}$ and d) $\Delta u_{3}$ components for an incident plane wave with horizontal slowness $p_{1}^{0}=0.07 \mathrm{~s} \mathrm{~km}^{-1}$. Examples of travel time curves, $\mathcal{T}^{q}(q=1,2,3,4,5,6$ in order of increasing time) are superposed for model point $\mathbf{x}=[240 \mathrm{~km}, 40 \mathrm{~km}]$.

## Chapter 7

## Migration Simulations

Chapter overview: The inversion methodology described in Chapter 5 is applied here to the synthetic data sets in a series of investigations to assess its potential and limitations. Simplifications to the migration/inversion theory permitted by the 2-D, in-plane propagation of the present experiment are discussed, followed by a brief summary of travel time calculations, the 1-D earth model chosen as reference, and the parameterization of inversion results. An examination of the numerical simulations begins with the imaging potential of forward-scattered modes from single events. Subsequently, all forward-scattered $P-S$ responses are combined in a single inversion to gauge the improvements resulting from enhanced source coverage. The remaining free surface reflected modes are then investigated to compare differences in resolution between forward- and back-scattered interactions. Following this, a majority of scattering interactions from all events are combined in a single inversion to further improve the reconstruction of subsurface discontinuous structure. This chapter concludes with a brief analysis of the degradation accruing from the introduction of noise and reduced station sampling.

### 7.1 Simplifications to Migration Theory

The intent of this section is to summarize several key results from the overview presented in Chapter 5 and amend them slightly to forms appropriate for the in-plane geometry of the present problem. The most obvious simplification involves the restriction of all vector quantities (e.g., filtered,
scattered displacement $\mathbf{v}$, incident slowness $\mathbf{p}^{0}$, scattered slowness $\mathbf{p}^{S}$ ) to the $x_{1}, x_{3}$ plane. Accordingly, equation (5.11) is rewritten as,
$g^{q}(\mathbf{x})=\int \mathrm{d} p_{1}^{0} \int \mathrm{~d} x_{1}^{\prime}\left|\frac{\partial(\psi, \theta)}{\partial\left(x_{1}^{\prime}, p_{1}^{0}\right)}\right| \sum_{l} W_{l}^{q}\left(x_{1}^{\prime}, p_{1}^{0}, \mathbf{x}\right) \frac{\left|\nabla \mathcal{T}^{q}\right|^{2}}{\left|\mathcal{A}^{q}\right|^{2}} \sum_{n} \mathcal{A}_{n}^{q}\left(\mathbf{x}, \mathbf{x}^{\prime}\right) v_{n}^{q}\left(\mathbf{x}^{\prime}, p_{1}^{0}, t=\mathcal{T}^{q}\right),(7$,
where quantities are defined as in Chapter 5. Note that the integration over sources is now cast in terms of the Cartesian horizontal slowness, $p_{1}^{0}$, rather then its absolute value, $\left|\mathbf{p}_{\perp}^{0}\right|=$ $\sqrt{\left(p_{1}^{0}\right)^{2}+\left(p_{2}\right)^{2}}$, cf. Chapter 5. The Jacobian change of variables in equation (7.1) is altered by the loss of one degree of freedom in the data variables, and, accordingly equation (73) in Paper $I$ reduces to,

$$
\begin{equation*}
\left|\frac{\partial(\psi, \theta)}{\partial\left(x_{1}^{\prime}, p_{1}^{0}\right)}\right|=\left|\frac{\cos \phi}{\left(J^{S}\right)^{2}} \frac{\alpha}{\sqrt{1-\left(p_{1}^{0}\right)^{2} \alpha^{2}}}\right| \tag{7.2}
\end{equation*}
$$

for a $P$ - $S$ interaction where $\phi$ is the surface take-off angle of the scattered ray and $J^{S}$ is the $S$-wave geometrical spreading function between scatter point, $\mathbf{x}$, and receiver location, $\mathbf{x}^{\prime}$. The Jacobian is modified from equation (7.2) when considering other scattering interactions by: i) a replacement of $\cdot J^{P}$ for $J^{S}$ for modes $q=1,3,5$; and ii) a substitution of $\beta$ for $\alpha$ for modes $q=5,6$. With these modifications, the material property perturbations at any particular scatter point are similarly retrieved as,

$$
\begin{equation*}
\Delta m=H^{-1} g \tag{7.3}
\end{equation*}
$$

where the three elements of gradient, g , are obtained from equation (7.1), and the Hessian, H , following equation (5.12), is given by,

$$
\begin{equation*}
H_{l q}(\mathbf{x})=\int \mathrm{d} p_{1}^{0} \int \mathrm{~d} \theta \sum_{k} W_{l}^{k}\left(\theta, p_{1}^{0}, \mathbf{x}\right) W_{k}^{q}\left(\theta, p_{1}^{0}, \mathbf{x}\right) \tag{7.4}
\end{equation*}
$$

where the dependence of $W_{l}^{q}$ on $\theta, p_{1}^{0}$, and $\mathbf{x}$ is explicitly identified.

### 7.2 Calculation of Travel Times

Equation (7.1) is akin to the diffraction stack of classical migration and defines the potential function $g^{q}(\mathbf{x})$, at any model point $\mathbf{x}$, as a weighted summation over all receivers and events of the data along predicted travel time curves for the particular mode interaction, $q$, as determined for the smoothly varying, 1-D reference medium (see section 7.3). Examples of travel time curves for the different scattering modes are superposed on Figures 6.2a-d for a model point $\mathbf{x}=[240 \mathrm{~km}, 40$ $\mathrm{km}]$ at Moho depth. The scattered wave travel times, $\mathcal{T}^{q}$, normalized with respect to the direct $P$ arrival, are easily calculated for a 1-D reference model. For a direct $P-S$ conversion ( $q=2$ ) travel time $\mathcal{T}^{(2)}$ is given by,

$$
\begin{gather*}
\mathcal{T}^{(2)}\left(\mathbf{x}, x_{1}^{\prime}, p_{1}^{0}\right)=\int_{0}^{x_{3}} \frac{\mathrm{~d} y_{3}}{\beta\left(y_{3}\right) \sqrt{1-\beta^{2}\left(y_{3}\right)\left(p_{1}^{S}\right)^{2}}}-p_{1}^{0}\left(x_{1}^{\prime}-x_{1}\right)  \tag{7.5}\\
-\int_{0}^{x_{3}} \frac{\mathrm{~d} y_{3}}{\alpha\left(y_{3}\right)} \sqrt{1-\alpha^{2}\left(y_{3}\right)\left(p_{1}^{0}\right)^{2}} \tag{7.6}
\end{gather*}
$$

where $p_{1}^{S}$ is the horizontal slowness of the ray travelling between scatter point $\mathbf{x}$ and receiver $\mathbf{x}^{\prime}$. For a direct $P-P$ interaction ( $q=1$ ), $p_{1}^{P}$ (horizontal slowness of the scattered $P$-wave) and $\alpha$ are substituted for $p_{1}^{S}$ and $\beta$, respectively.

When calculating the expected arrival times for back-scattered waves, both the extra time incurred by a wave traveling to surface and returning to the scatterer point, and for the possibility of $P-S$ conversion at the free surface must be taken into account. Thus, the expected travel time for, e.g., a back-scattered $P-P(q=3)$ interaction is

$$
\begin{gather*}
\mathcal{T}^{(3)}\left(\mathbf{x}, x_{1}^{\prime}, p_{1}^{0}\right)=\int_{0}^{x_{3}} \frac{\mathrm{~d} y_{3}}{\alpha\left(y_{3}\right) \sqrt{1-\alpha^{2}\left(y_{3}\right)\left(p_{1}^{P}\right)^{2}}}-p_{1}^{0}\left(x_{1}^{\prime}-x_{1}\right)  \tag{7.7}\\
+\int_{0}^{x_{3}} \frac{\mathrm{~d} y_{3}}{\alpha\left(y_{3}\right)} \sqrt{1-\alpha^{2}\left(y_{3}\right)\left(p_{1}^{0}\right)^{2}} \tag{7.8}
\end{gather*}
$$

Travel times for other phases are calculated with slight modification: i) $P$-S scattering ( $q=4$ ) requires substitutions of $p_{1}^{S}$ and $\beta$ for $p_{1}^{P}$ and $\alpha$ in the first term; ii) for $S-P$ scattering ( $q=5$ ), $\alpha$ is
replaced by $\beta$ in the third term; and iii) $S$ - $S$ scattering ( $q=6$ ) requires all three of the substitutions in i) and ii).

Two further points are worth noting. First, the travel times of scattered modes which share similar polarizations at the receiver differ only by static time shifts for the 1-D reference model. Second, asymmetries about the scatter point exist in the travel time moveout with distance which result from the $p_{1}^{0}$ dependence in equations (7.5) and (7.6) as manifest by tilt in the hyperbolae in Figures 6.2a-d. In particular, the expected travel time moveout for, e.g., a forward-scattered $P-S$ interaction in a 1-D medium is noted to be,

$$
\begin{equation*}
\frac{\partial \mathcal{T}^{(2)}}{\partial x_{1}^{\prime}}\left(\mathbf{x}, x_{1}^{\prime}, p_{1}^{0}\right)=-p_{1}^{0}+\frac{\partial p_{1}^{S}}{\partial x_{1}^{\prime}} \int_{0}^{x_{3}} \mathrm{~d} y_{3} \frac{\beta\left(y_{3}\right) p_{1}^{S}}{\left[1-\beta^{2}\left(y_{3}\right)\left(p_{1}^{S}\right)^{2}\right]^{\frac{3}{2}}} \tag{7.9}
\end{equation*}
$$

### 7.3 Reference Model and Parameterization

In the sections that follow, a 1-D reference model consisting of a 40 km thick crust overlying an upper mantle half-space is utilized. Velocities and densities of these two reference materials are identical to the crustal and mantle properties specified in the first two rows of Table 6.1. In addition, inversion for only $\Delta \alpha / \alpha$ and $\Delta \beta / \beta$, the two most linearly independent parameters in the case of forward-scattering (e.g., $B R$ ), is considered since the third parameter (i.e., $\Delta \rho / \rho$ ) remains effectively indeterminate for surface receiver arrays. In situations where back-scattering is represented, this is recognized as not the best choice of parameter combination (see, e.g., Forgues and Lambaré, [1997]). However, Forgues and Lambaré [1997] demonstrate that any choice of parameterization may be simply recovered through matrix multiplications after inversion. Thus, for simplicity, all inversion results are presented in terms of the physically meaningful, though not always optimal, velocity perturbations.

### 7.4 Receiver Function Image of Synthetic Data

Before proceeding to detail individual experiments, it is desirable to present a basis for comparison between our inversion methodology and conventional teleseismic processing. A receiverfunction image (e.g., Langston, [1979]) constructed via least-squares simultaneous deconvolution (e.g., Gurrola et al., [1995]) of individual $S$-wave traces by corresponding $P$-wave traces (as determined in step 1 of the pre-processing sequence) from all six events is shown in Figure 7.1. Well defined $P$-to- $S$ conversions from the Moho and subducted crust are apparent to the left and right of the suture. However, traces to the left side of the section and near the suture at offsets of $\sim 120 \mathrm{~km}$ are corrupted by strong forward-scattered $P-P$ and $P-S$ diffraction energy (cf. data


Figure 7.1: Least squares receiver function image constructed through simultaneous deconvolution of individual $S$-wave traces by corresponding $P$-wave traces for all six suites of seismograms.
sections in Figures 6.2). Free-surface multiples corresponding to interactions $q=3,4,6$ are evident at times later than 10 s , and may be potentially misinterpreted as deeper structure. A formal, multiple mode inversion is expected to significantly reduce these contaminating artifacts, and allow for an improved imaging of more complex structure.

### 7.5 Experiment I - Single Event Inversion of Forward-Scattering Modes

To examine the bias resulting from source direction, forward-scattering modes $q=1,2$ are inverted separately for two events with positive ( $p_{1}^{0}=0.07 \mathrm{~s} \mathrm{~km}^{-1}$ ) and negative ( $p_{1}^{0}=-0.07 \mathrm{~s} \mathrm{~km}^{-1}$ ) horizontal slowness (Figures 7.2 and 7.3, respectively). Because of the limited sampling afforded by a single event, recovery is restricted to a single material parameter for both modes using the Jacobian, $\left|\partial \psi / \partial x_{1}^{\prime}\right|$, in equation (74) of Paper $I$. In particular, perturbations $\Delta \alpha / \alpha$ for $q=1$ and $\Delta \beta / \beta$ for $q=2$ are considered as optimal choices for forward scattering. Furthermore, due to the strong amplitudes evident in Figure 7.1 between zero and four seconds, a cosine taper has been applied to the beginning of all traces employed in this and the following experiments. The $P$-velocity perturbations reconstructed from the forward-scattered $P-P(q=1)$ interaction are presented in Figure 7.2a. Vertical streaking about the suture zone with the correct general polarity is observed though little focussing of energy is achieved. $S$-velocity perturbations recovered by the direct $P-S(q=2)$ interaction (Figure 7.2b) afford a much better delineation of the normal-relict crust boundary. Note that apart from this boundary, little of the remaining, more planar components of structure is reconstructed.

Figures 7.3a,b present the same parameter and mode interaction combinations as Figures $7.2 \mathrm{a}, \mathrm{b}$, but now for an incident wave with slowness $-0.07 \mathrm{~s} \mathrm{~km}^{-1}$. Figure 7.3 a shows increased streaking about the suture zone, relative to Figure 7.2 a , and a correlation with actual structure


Figure 7.2: Reconstructed velocity perturbations from inversion of forward-scattered modes for a left-incident plane wavefield $\left(p_{1}^{0}=0.07 \mathrm{~s} \mathrm{~km}^{-1}\right)$. a) $\Delta \alpha / \alpha$ from $q=1$; b) $\Delta \beta / \beta$ from $q=2$.
that is poorer still. Unlike Figure 7.2b, the $S$-velocity image in Figure 7.3 b reconstructs the relict crust-mantle boundaries (though displaced upwards from their true locations), but identifies little contrast with adjacent normal crust. The Moho is now more prominent, though its location is


Figure 7.3: Reconstructed velocity perturbations from inversion of forward-scattered modes for a right-incident plane wavefield $\left(p_{1}^{0}=-0.07 \mathrm{~s} \mathrm{~km}^{-1}\right)$. a) $\Delta \alpha / \alpha$ from $q=1$; b) $\Delta \beta / \beta$ from $q=2$.
slightly displaced from the known lithospheric model. The mislocation of structure is related to the inadequacy of the reference velocity model in correctly predicting the travel times of scattered phases. Where Moho depths are less then 40 km , there is a corresponding "pull up" of underlying
structure, while the converse is true in areas where the Moho exceeds 40 km .

The present experiment serves to illustrate several of the shortcomings inherent in singlemode inversions in general, and forward-scattering inversions in particular. Of specific note is the poor recovery of structure in the forward-scattered $P$ - $P$ images (Figures 7.2a and 7.3a). The origin of this lack of sensitivity can be traced directly to the factor $\left|\nabla \mathcal{T}^{q}\right|^{2}$ that appears in equation (7.1) and weights the diffraction stack according to the sensitivity of total travel time to scatterer location. For $q=1$ in the strictly forward-scattering direction $\left(\theta=180^{\circ}\right)$, the sensitivity is zero, as discussed extensively by Marquering et al. [1999], and remains relatively small for a range of angles about the forward direction. A consequence of this behaviour is that considerable tradeoff exists between the lateral extent of the scattering body and its depth in near-forward directions. This effect is manifest in the diffuse appearance of reconstructed anomalies in Figures 7.2a and 7.3a.

A second issue worth noting is the contrasting recovery of structure for events with differing horizontal slowness. In general, the successful reconstruction of a material property perturbation is dependent upon the strength of the associated scattered response. In the case of extended, quasiplanar discontinuities separating gently varying media, this scattered response will be most pronounced in directions corresponding to the specular angle of interaction. For $P$-to- $S$ conversions ( $q=4$ ), however, the effect of conversion coefficients, which tend to zero as the specular angle approaches $0^{\circ}$, must also be considered (i.e., normal incidence). In this case, the incident wavefield has little sensitivity to structure since no $P-S$ scattering results. Consequently, the structure can be regarded as falling within the null space of the problem and will not be recovered within the inversion.

These effects are well illustrated in Figures 7.2b and 7.3b. In the case of left incidence (Figure $6.2 \mathrm{c}, \mathrm{d}), P-S(q=2)$ scattering from the subhorizontal boundaries is weak since the incoming wavefield is incident at near-normal angles. The sole exception is the short boundary segment between normal and relict crust which is sufficiently limited and oblique to generate strong diffractions. Consequently, only the latter feature is well resolved in the inversion Figure 7.2b. In contrast, illumination of the lithospheric section by a plane wave from the right (Figure 6.2a,b) produces stronger $P$ - $S$ conversions from the near-planar structures as they are oriented at more oblique angles to the incident wave. Accordingly, Figure 7.3b documents an improved reconstruction of the Moho and the lateral extent of subducted relict crust.

### 7.6 Experiment II - Multiple Event Inversion of Forward-Scattering Modes

The present section investigates the effect of simultaneous inversion of multiple events on the reconstruction of material properties. In this experiment and those that follow, the Jacobian change of variables in equation (7.2) is employed as a range of incident slowness $p_{1}^{0}$ is now sampled. The reconstructed $P$ - and $S$-velocity images from the inversion of all six forward-scattered $P$ - $P$ and $P-S$ interactions are presented in Figures 7.4a,b, respectively. As might be expected, the image in Figure 7.4 b combines the structural features imaged separately in Figures 7.2 b and 7.3 b . In particular, all of the boundaries which define the suture are now well resolved, and Moho velocity perturbations are retrieved at laterally coherent levels along the entire breadth of the array. However, the amplitude of the reconstructed Moho discontinuity is underestimated relative to the short wavelength suture. This is, in part, a consequence of the volume-scattering formulation as is discussed further in the following section. It is also worth noting that some contamination from scattered $P-P(q=1)$ enters the image at shallow crustal levels.


Figure 7.4: Reconstructed velocity perturbations from inversion of forward-scattered modes of all six incident plane wavefields. a) $\Delta \alpha / \alpha$ from $q=1$; b) $\Delta \beta / \beta$ from $q=2$.

Figure 7.4a documents little improvement over Figures 7.2a and 7.3a in the recovery of
discontinuous structure. In contrast with Figure 7.4b, recovered perturbations show poor correlation with the true structure. This result is thought to be due to a combination of relative insensitivity of the forward-scattered $P-P$ mode to structural location and a distortion of the actual scattered $P$-wavefield incurred during preprocessing. Accordingly, this mode will not be employed in further inversions.

### 7.7 Experiment III - Multiple Event, Back-Scattered Mode Inversion

In this section, the inversion procedure is applied to back-scattered modes $q=3,4,5,6$ (Figures $7.5 \mathrm{a}-$ d) from all six incident wavefields. Figure 7.5a presents the reconstructed $P$-velocity perturbations for the $P-P(q=3)$ reflection. In this image, the Moho and normal-relict crustal boundary


Figure 7.5: Reconstructed velocity perturbations from inversion of individual back-scattered modes for all six incident plane wavefields. a) $\Delta \alpha / \alpha$ from $q=3$; b) $\Delta \beta / \beta$ from $q=4$; c) $\Delta \beta / \beta$ from $q=5 ;$ d) $\Delta \beta / \beta$ from $q=6$.
are accurately delineated with locations that correspond well with the actual model. In particular, the spatial localization and placement of these discontinuities is enhanced relative to that of the forward-scattering inverisons. This improvement is directly attributable to the increased sensitivity of back-scattered travel time variations to scatterer location as represented by $\left|\nabla \mathcal{T}^{q}\right|^{2}$ in equation (7.1) and discussed in Paper I. Figure 7.5b presents the $S$-velocity perturbations recovered from the $P-S(q=4)$ reflection. The Moho is once more well defined and is less smeared near the suture zone than in Figure 7.5a. The two discontinuous interfaces of the relict crust are also apparent and indicate the presence of a high velocity anomaly. Later arriving energy from the $S-S(q=6)$ reflection contaminates the image at depths of $\sim 60 \mathrm{~km} . S$-velocity perturbations retrieved from the $S-P(q=5)$ reflection are shown in Figure 7.5c. The strongest anomaly is located at mid-lower crustal depths, and indicates that the image is strongly contaminated by the earlier and stronger $P-P(q=3)$ reflections. $S$-velocity perturbations retrieved by the $S-S(q=6)$ interaction (Figure 7.5d) document a good recovery of the Moho and suture. Earlier $P-S$ energy is erroneously mapped to depths near $\sim 30 \mathrm{~km}$. In general, back-scattered modes are observed to effectively resolve laterally discontinuous structure, though artifacts caused by the mis-imaging of interfering phases are more prevalent than in Figure 7.4.

An assumption inherent in the methodology is that individual scattering modes have been isolated prior to data input. In practice, this assumption can only be realized insofar as scattering modes ending as $P$ - or $S$-waves can be separated on the basis of approximately orthogonal polarizations. Therefore, an obvious complication is the degree to which single mode reconstructions are contaminated by energy from other scattering interactions of similar wavetype. For example, over the range of horizontal slowness utilized in these investigations, the free-surface corrected amplitudes of $S-S$ reflections from the Moho are only slightly larger ( $\sim 30 \%$ ) than those for $P-S$ reflections. Since these phases possess similar amplitudes, both yield accurate estimates, relative
to each other, of the jumps in material properties across the Moho. However, each phase also contributes to phantom structure in the other image. This situation also exists for $P$-waves, though, in this case the amplitudes of $P-P$ reflections outweigh those of $S-P$ by a factor of $\sim 4$. As a result, the back-scattered $S$ - $P$ interaction is far less effective in the imaging process. As shown in the following section, the problem of mode inseparability is partially offset through simultaneous inversion of all modes; however, it is important for confident interpretation that individual mode inversions are examined at this stage for features common to all images.

Although the appearance of forward and back-scattered phases on seismograms (e.g., $P-S$ interactions $q=2,4$ in Figure 6.2) are quite similar (aside from the static time shifts), they produce rather different images. In particular, the jump across the Moho in the forward-scattering reconstruction is noted to be rather more diffuse relative to shorter wavelength structure than for back scattered modes. An understanding of this behaviour requires an examination of the contributions from different factors $\left(W_{l}^{q}(\theta),\left|\nabla \mathcal{T}^{q}\right|^{2},\left|\partial \psi / \partial x_{1}^{\prime}\right|\right)$ that weight the diffraction stack. These quantities and their product are presented in Figures 7.6 and 7.7 for modes $q=2,4$, respectively. The depth of the scatterer point under consideration is 40 km and the slowness of the incident wave is $0.05 \mathrm{~s} \mathrm{~km}^{-1}$. For reference, the location of the travel time minimum is shown in Figures 7.6 and 7.7 as a dotted line. Figures 7.6 b ,d show the travel time sensitivity $\left|\nabla \mathcal{T}^{(2)}\right|^{2}$ and radiation pattern $W_{2}^{2}(\theta)$ factors. Both factors contribute to significant downweighting near the travel time apex, whereas the Jacobian of the transformation of variables, $\left|\partial \psi / \partial x_{1}^{\prime}\right|$ (Figure 7.6c) is sharply peaked and maximally weights the diffraction stack close to the minimum traveltime. The cumulative product which represents the total weight along the diffraction stack is bimodal with the travel time apex falling away from the global minimum such that at the apex, the diffraction stack receives only $\sim 25 \%$ of the maximum weight. In the case of $P-S$ back-scattering ( $q=4$ ) (Figure 7.7), travel time sensitivity and radiation pattern are more favorably positioned in Figures 7.7b,d


Figure 7.6: Diffraction stack weights for forward-scattered $P-S$ interaction ( $q=2$ ). a) Total travel time $\mathcal{T}^{(2)} ;$ b) travel time sensitivity $\left|\nabla \mathcal{T}^{(2)}\right|^{2}$; c) Jacobian of the transformation of variables $\mid \partial \psi\left(\partial x_{1}^{\prime} \mid\right.$ d d radiation pattern $W_{2}^{2}$; and e) the product of b), c) and d).
such that the maximum of the total weighting function is in close proximity to the apex of $\mathcal{T}^{(4)}$. The Jacobian (Figure 7.7c) is slightly broadened relative to that in Figure 7.6c. Consequently, the cumulative weighting function, Figure 7.7 e, is again bimodal, but the $\mathcal{T}^{(4)}$ apex is now positioned close to the location of maximum weight.

While these weights are appropriate and effective for single point scatterers, they have a rather different efficiency for quasi-planar structures. Since the diffraction hyperbolae are tangent to the travel time locii for planar discontinuities at the hyperbola apex, reconstruction using forward scattered energy extracts a lower proportion of signal from the true arrival. It is therefore more


Figure 7.7: Diffraction stack weights for back-scattered $P$-S interaction ( $q=4$ ). a) Total travel time $\mathcal{T}^{(4)} ;$ b) travel time sensitivity $\left|\nabla \mathcal{T}^{(4)}\right|^{2} ;$ c) Jacobian of the transformation of variables $\left|\partial \psi / \partial x_{1}^{\prime}\right|$; d) radiation pattern $W_{2}^{4}$, and e) the product of $b$ ), c) and d).
prone to contamination from noise and may be more susceptible to incomplete coverage in $\psi$.

### 7.8 Experiment IV - Multiple Mode, Multiple Event Inversion

In this experiment, the improvements in model reconstruction incurred through simultaneous inversion of both forward- and back-scattered modes $(q=2,3,4,6)$ are examined. As mentioned above, the inability to fully separate individual scattering modes results in images that suffer from cross-mode contamination (cf. Experiment III). Therefore, it is justified to weight individual mode contributions in the simultaneous inversion on some a priori basis which accounts for this
contamination. In particular, relative signal amplitudes as governed by free surface reflection coefficients and the anticipated response of underlying heterogenity must be considered. In the present experiment, these factors have been considered in a somewhat subjective fashion and have been chosen so as to i) downweight the $P-S(q=4)$ interaction by $30 \%$ to bring the contaminating $S-S(q=6)$ interaction structure to approximately comparable levels with the $P-S(q=4)$ phantom structure observed in the $S-S(q=6)$ interaction image, and ii) upweight the forward-scattered $P-S$ ( $q=2$ ) interaction by a factor of 6 to allow for a more balanced representation of forward-scattering in the simultaneous inversion results.


Figure 7.8: Reconstructed velocity pertubations from simultaneous inversion of individually weighted modes from all six events. a) $\Delta \alpha / \alpha$; b) $\Delta \beta / \beta$.

Figures 7.8a,b present the reconstructed $P$ - and $S$-velocity perturbations from the inversion of the four individually weighted modes, respectively. In both figures, the Moho and the relict crust are well delineated in reference to model structure. In accordance with the true velocity model (see Table 6.1 ), the $S$-velocity image identifies the relict crust as a relatively larger perturbation than its $P$-velocity counterpart. Furthermore, in contrast with the images recovered from individual back-scattered waves (i.e., Experiment III), the resulting $S$-velocity perturbation image is now less contaminated by phantom structure leaving the suture zone more clearly evident. In general, the simultaneous inversion of individually weighted scattering interactions realizes a wellconstrained, and improved reconstruction of the model.

### 7.9 Experiment V-Noise and Spatial Aliassing Issues

The previous experiments have employed noise free data sets to investigate the inversion/migration methodology developed in Paper I. The ability of the approach to perform in a situation where data quality and recording geometry are more representative of typical field experiments is now examined. The idealized suture model has been altered to include a surface layer of 6 km average thickness (Figure 7.9a). The layer is parameterized by $P$-wave and $S$-wave velocity and density values of $5.2 \mathrm{~km} \mathrm{~s}^{-1}, 2.8 \mathrm{~km} \mathrm{~s}^{-1}$, and $2.5 \mathrm{~g} \mathrm{~cm}^{-3}$, respectively, and exhibits random undulations of maximum amplitude 1.2 km which have give rise to strong near-surface body-wave and surface-wave scattering (see Figures $7.9 \mathrm{~b}, \mathrm{c}$ ). In addition, a random selection of 60 stations has been taken from the previous 120 stations resulting in receiver spacings between 3 km and 15 km , and data sections have been contaminated with Gaussian noise of zero mean and a standard deviation equivalent to $5 \%$ of the amplitude of the incident wave. Given the frequency content of the incident wavefield, it is expected that the resulting images will suffer to some degree from the effects of spatial aliassing.

Figures $7.9 \mathrm{~d}, \mathrm{e}$ present the $P$ - and $S$-velocity images using the same weighted mode contribution and event selection as Experiment IV. In both images, the Moho and normal-relict crustal discontinuities remain well delineated. The $S$-velocity image also accurately retrieves the remaining discontinuities of the suture zone. Scattering from the near-surface layer has introduced to the images an additional source of noise that has been mapped back to the near surface. Additive white noise has had a relatively minor effect on the reconstruction since it is largely cancelled through the integration in equation (7.1). Figure 7.9d, e do suffer from spatial aliasing manifest


Figure 7.9: Reconstructed velocity pertubations from simultaneous inversion of individually weighted modes from all events including a model with an additional surface layer, and data with additive noise and reduced irregular station distribution. a) Revised subduction-suture model; b) Preprocessed (but prior to filtering by $F(\omega)) \Delta u_{3}$ synthetic data section ( $p_{1}^{0}=-0.07 \mathrm{~s} \mathrm{~km}^{-1}$ ) generated for model shown in a); c) As in b) but for $\Delta u_{1}$; d) $\Delta \alpha / \alpha$; e) $\Delta \beta / \beta$.
through the high wavenumber "speckle" which is, however, most prominent at shallow levels and diminishes with depth so as not to pose significant impediment in the reconstruction of lower crustal and mantle structures.

## Chapter 8

## Concluding Remarks

Part I of this thesis presents the results from a number of analyses involving teleseismic data collected in southern Alberta. Results from travel-time tomography document a high velocity anomaly underlying a substantial portion of the southern Hearne Province which is interpreted to represent deep-seated lithospheric structure. These results suggest that the bulk of the lithosphere has remained intact, and that the anomously high mantle conductivity values noted in magnetotelluric surveys are probably the result of connected hydrous minerals or some other conductive species introduced during subduction processes responsible for a thickened lithospheric root. Shear wave splitting results reveal a relatively uniform set of splitting parameters with an average polarization direction that broadly corresponds to both the (presumably) preserved fossil strain fields arising from the ca. 1.8 Ga , NW-SE shortening of the Hearne Province, and to the absolute motion of the North American Plate. Processing of radial receiver functions yield Moho depth estimates that are fairly uniform in the northern sections of the array, but reveal crustal thickening to the south within Medicine Hat Block.

Part II of this thesis presents results from a number of simulations that test the inverse scattering/migration methodology for teleseismic waves developed in Paper $I$. The flexibility of the method has been exploited to investigate the roles that different scattering mode and event combinations play in the recovery of structure corresponding to an idealized lithospheric suture. In
particular, it demonstrates: i) the contrasting and complementary sensitivity of forward- and backscattered modes to structure; ii) the reduction in problem null-space that accompanies improved source coverage; iii) improvements in structural recovery incurred through simultaneous treatment of different scattering modes, and iv) the robustness of the approach in the presence of typical signal-to-noise levels and irregular station geometry.

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## Appendix A

## Data set employed in Chapter 3 analyses

Table A.1: Events used in the analyses of Chapter 3. Individual techniques are denoted by: TTI- $P$, $P$-wave travel-time inversion; RF, receiver functions; SWS, shear wave splitting.

| Date YY:MM:DD | Time hh:mm:ss.s | Latitude ${ }^{\circ} \mathrm{N}$ | Longitude ${ }^{\circ} \mathrm{E}$ | $\begin{gathered} \text { Depth } \\ \text { km } \end{gathered}$ | $\mathrm{m}_{6}$ | TTI-P | RF | SWS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 96:05:02 | 13:34:28 | -4.548 | 154.833 | 500 | 5.6 |  | X | X |
| 96:06:17 | 11:22:18 | -7.137 | 122.589 | 587 | 6.6 |  | X |  |
| 96:07:20 | 07:41:15 | -19.820 | -177.643 | 357 | 5.7 |  | X |  |
| 96:08:05 | 22:38:22 | -20.690 | -178.310 | 550 | 6.4 |  | X |  |
| 96:10:19 | 14:53:48 | -20.410 | -178.510 | 425 | 5.8 |  | X | X |
| 96:11:05 | 09:41:34 | -31.160 | 179.998 | 369 | 6.4 |  | X | X |
| 97:09:04 | 04:23:37 | -26.569 | 178.336 | 625 | 6.3 |  | X | X |
| 98:03:29 | 19:48:16 | -17.552 | -179.092 | 537 | 6.5 |  | X | X |
| 98:05:16 | 02:22:03 | -22.227 | -179.519 | 586 | 6.1 |  | X | X |
| 98:07:29 | 07:14:24 | -32.312 | -71.286 | 51 | 6.3 |  | X |  |
| 98:08:02 | 04:40:46 | 39.573 | 76.999 | 69 | 5.6 |  | X |  |
| 98:08:04 | 18:59:20 | -0.593 | -80.393 | 33 | 6.2 |  | X |  |
| 98:08:05 | 10:42:21 | 56.164 | 163.360 | 33 | 5.2 | X |  |  |
| 98:08:09 | 05:27:43 | 53.008 | 171.139 | 25 | 5.3 | X |  |  |
| 98:08:11 | 02:07:49 | 25.014 | 123.017 | 141 | 5.0 | X |  |  |
| 98:08:13 | 06:16:51 | 53.068 | 171.106 | 33 | 5.0 | X |  |  |
| 98:08:18 | 18:00:12 | 45.852 | 149.116 | 116 | 5.4 | X |  |  |
| 98:08:20 | 06:40:55 | 28.932 | 139.329 | 441 | 6.1 | X | X |  |
| 98:08:20 | 09:36:34 | 45.561 | 136.926 | 351 | 5.2 | X |  |  |
| 98:08:20 | 14:56:40 | 51.531 | 175.389 | 33 | 5.4 | X |  |  |
| 98:08:20 | 15:00:08 | 51.618 | 175.248 | 33 | 5.6 | X | X |  |
| 98:08:23 | 05:36:12 | 14.697 | 120.046 | 70 | 6.1 | X |  |  |
| 98:08:23 | 13:57:15 | 11.663 | -88.038 | 55 | 5.7 | X | X |  |
| 98:08:25 | 07:41:40 | 30.079 | 88.109 | 33 | 5.3 | X |  |  |
| 98:08:27 | 16:51:45 | 19.274 | -108.446 | 10 | 5.0 | X |  |  |
| 98:08:28 | 15:31:38 | 51.437 | 175.526 | 33 | 5.5 | X |  |  |
| 98:08:28 | 23:46:43 | 35.522 | 139.879 | 76 | 5.2 | X |  |  |

Table A.1: (Continued)

| $\begin{gathered} \text { Date } \\ \text { YY:MM:DD } \end{gathered}$ | Time hh:mm:ss.s | Latitude ${ }^{\circ} \mathrm{N}$ | Longitude ${ }^{\circ} \mathrm{E}$ | $\begin{array}{r} \text { Depth } \\ \mathrm{km} \end{array}$ | $\mathrm{m}_{b}$ | TTI-P | RF | SWS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 98:08:30 | 01:48:08 | 17.092 | 148.133 | 33 | 6.0 | X | X |  |
| 98:08:30 | 14:34:43 | 53.669 | 161.867 | 33 | 5.5 | X |  |  |
| 98:09:01 | 01:19:37 | -17.555 | -174.771 | 220 | 5.3 | X |  |  |
| 98:09:01 | 08:29:12 | -28.049 | -70.439 | 87 | 4.6 | X |  |  |
| 98:09:01 | 14:33:41 | -16.758 | -173.597 | 33 | 5.4 | X |  |  |
| 98:09:02 | 08:37:29 | 5.410 | 126.764 | 50 | 6.6 | X | X | X |
| 98:09:03 | 06:43:03 | 39.539 | 77.260 | 33 | 5.1 | X |  |  |
| 98:09:03 | 07:58:21 | 39.716 | 140.760 | 38 | 5.7 | X | X |  |
| 98:09:03 | 17:37:58 | -29.450 | -71.715 | 27 | 6.2 | X | X |  |
| 98:09:03 | 18:15:56 | 27.850 | 86.941 | 33 | 5.6 | X |  |  |
| 98:09:05 | 05:16:17 | -6.646 | 155.209 | 37 | 5.5 | X |  |  |
| 98:09:07 | 00:23:02 | -32.347 | -111.950 | 10 | 5.3 | X |  |  |
| 98:09:07 | 00:39:30 | -36.240 | -97.711 | 10 | 5.2 | X |  |  |
| 98:09:07 | 02:35:03 | -5.766 | 152.061 | 41 | 5.2 | X |  |  |
| 98:09:08 | 09:10:03 | 13.257 | 144.007 | 141 | 5.8 | X | X | X |
| 98:09:09 | 11:27:59 | 40.035 | 15.980 | 10 | 5.2 | X |  |  |
| 98:09:10 | 09:51:24 | -20.028 | -70.378 | 33 | 4.9 | X |  |  |
| 98:09:12 | 09:03:48 | -24.512 | -67.119 | 187 | 5.1 | X |  |  |
| 98:09:12 | 10:58:04 | -14.233 | -72.615 | 91 | 5.3 | X |  |  |
| 98:09:14 | 23:16:46 | 51.618 | -173.150 | 33 | 5.7 | X | X |  |
| 98:09:15 | 07:24:06 | 38.346 | 140.508 | 50 | 5.1 | X |  |  |
| 98:09:15 | 08:35:51 | -5.624 | 151.637 | 83 | 5.6 | X | X |  |
| 98:09:16 | 02:12:02 | -6.580 | 154.868 | 87 | 5.4 | X |  |  |
| 98:09:16 | 11:09:14 | -24.047 | -66.744 | 175 | 4.9 | X |  |  |
| 98:09:16 | 11:28:56 | -3.242 | -79.348 | 104 | 4.9 | X |  |  |
| 98:09:17 | 16:41:20 | -38.294 | -93.623 | 10 | 5.3 | X |  |  |
| 98:09:17 | 18:14:25 | 30.593 | 137.557 | 487 | 4.5 | X |  |  |
| 98:09:18 | 03:51:13 | 43.212 | 148.097 | 10 | 5.1 | X |  |  |
| 98:09:20 | 00:36:46 | -21.392 | -174.709 | 100 | 5.1 | X |  |  |
| 98:09:21 | 13:00:04 | 48.360 | 148.771 | 396 | 4.3 | X |  |  |
| 98:09:22 | 01:16:55 | 11.822 | 143.154 | 9 | 5.8 | X | X |  |
| 98:09:24 | 18:53:40 | 46.313 | 106.288 | 33 | 5.5 | X |  |  |
| 98:09:27 | 11:07:16 | -20.270 | -175.876 | 207 | 5.2 | X |  |  |
| 98:09:28 | 13:34:30 | -8.194 | 112.413 | 152 | 6.4 |  |  | X |
| 98:09:28 | 19:23:23 | 3.840 | 126.407 | 30 | 5.8 | X |  |  |
| 98:09:29 | 22:14:49 | 44.209 | 20.080 | 10 | 5.2 | X |  |  |
| 98:10:01 | 03:41:13 | 13.738 | -45.565 | 10 | 5.4 | X |  |  |

Table A.1: (Continued)

| $\begin{gathered} \text { Date } \\ \text { YY:MM:DD } \end{gathered}$ | Time hh:mm:ss.s | Latitude <br> ${ }^{\circ} \mathrm{N}$ | Longitude ${ }^{\circ}$ E | $\begin{array}{r} \text { Depth } \\ \mathrm{km} \end{array}$ | $\mathrm{m}_{b}$ | TTI-P | RF | SWS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 98:10:01 | 20:44:06 | 9.675 | -82.443 | 31 | 5.3 | X |  |  |
| 98:10:02 | 12:49:35 | 27.268 | 101.016 | 48 | 5.2 | X |  |  |
| 98:10:03 | 11:15:42 | 28.505 | 127.615 | 227 | 5.6 | X | X |  |
| 98:10:06 | 12:27:41 | 37.247 | 21.107 | 10 | 5.0 | X |  |  |
| 98:10:07 | 06:53:14 | 21.536 | 143.057 | 300 | 5.2 | X |  |  |
| 98:10:07 | 13:41:01 | 12.902 | 146.719 | 34 | 5.0 | X |  |  |
| 98:10:08 | 04:51:42 | -16.119 | -71.404 | 136 | 6.1 | X | X |  |
| 98:10:09 | 05:37:52 | -15.393 | -173.392 | 33 | 5.0 | X |  |  |
| 98:10:09 | 11:54:36 | 11.321 | -86.451 | 69 | 5.5 | X |  |  |
| 98:10:10 | 04:12:08 | -33.518 | -72.078 | 33 | 5.3 | X |  |  |
| 98:10:10 | 20:18:39 | 23.728 | 142.936 | 33 | 4.9 | X |  |  |
| 98:10:11 | 12:04:54 | -21.040 | -179.110 | 624 | 5.4 | X |  |  |
| 98:10:11 | 21:44:16 | -27.329 | -63.335 | 583 | 5.3 | X |  |  |
| 98:10:12 | 07:36:39 | -15.278 | -173.579 | 77 | 4.9 | X |  |  |
| 98:10:13 | 18:13:04 | -20.861 | -178.770 | 600 | 4.5 | X |  |  |
| 98:10:13 | 20:41:13 | 40.027 | 143.297 | 33 | 5.4 | X |  |  |
| 98:10:14 | 01:36:19 | 60.711 | -44.050 | 10 | 5.1 | X |  |  |
| 98:10:14 | 02:54:04 | -5.915 | 151.036 | 33 | 5.5 | X | X |  |
| 98:10:18 | 01:39:01 | 24.718 | 141.237 | 110 | 5.5 | X | X |  |
| 98:10:18 | 08:33:54 | 19.285 | 145.341 | 152 | 5.4 | X |  |  |
| 98:10:18 | 22:09:19 | 86.283 | 75.609 | 10 | 5.2 | X |  |  |
| 98:10:19 | 23:57:33 | -21.254 | -68.969 | 124 | 4.7 | X |  |  |
| 98:10:20 | 11:52:40 | -20.594 | -174.299 | 33 | 5.1 | X |  |  |
| 98:10:23 | 01:48:51 | -2.423 | -76.359 | 147 | 5.3 | X |  |  |
| 98:10:24 | 08:27:12 | -17.719 | -175.222 | 261 | 5.1 | X |  |  |
| 98:10:25 | 03:54:39 | -17.974 | -69.592 | 33 | 5.0 | X |  |  |
| 98:10:25 | 20:06:05 | 36.438 | 68.585 | 61 | 5.1 | X |  |  |
| 98:10:26 | 02:34:57 | -21.224 | -178.905 | 574 | 4.9 | X |  |  |
| 98:10:27 | 11:33:37 | 33.489 | 141.398 | 33 | 5.3 | X |  |  |
| 98:10:28 | 21:39:52 | 11.985 | 143.528 | 33 | 5.5 | X |  |  |
| 98:10:31 | 12:45:50 | -17.881 | -178.318 | 577 | 4.8 | X |  |  |
| 98:10:31 | 13:38:50 | 29.328 | 142.042 | 19 | 5.1 | X |  |  |
| 98:10:31 | 14:03:32 | 53.049 | 157.859 | 53 | 5.2 | X |  |  |
| 98:11:02 | 23:10:59 | 43.665 | 147.620 | 58 | 5.5 | X | X |  |
| 98:11:04 | 02:54:13 | 52.094 | -176.143 | 35 | 5.1 | X |  |  |
| 98:11:05 | 03:45:17 | -10.322 | -78.364 | 51 | 5.4 | X |  |  |
| 98:11:05 | 20:09:55 | -15.613 | -177.990 | 33 | 5.0 | X |  |  |

Table A.1: (Continued)

| $\begin{gathered} \text { Date } \\ \text { YY:MM:DD } \end{gathered}$ | $\begin{gathered} \text { Time } \\ \text { hh:mm:ss.s } \end{gathered}$ | Latitude ${ }^{\circ} \mathrm{N}$ | Longitude ${ }^{\circ} \mathrm{E}$ | $\begin{array}{r} \text { Depth } \\ \mathrm{km} \end{array}$ | $\mathrm{m}_{b}$ | TTI-P | RF | SWS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 98:11:06 | 12:24:07 | -21.806 | -67.162 | 210 | 4.9 | X |  |  |
| 98:11:07 | 05:35:41 | 41.592 | 142.109 | 79 | 4.7 | X |  |  |
| 98:11:09 | 04:57:41 | -21.148 | -174.452 | 53 | 5.2 | X |  |  |
| 98:11:10 | 00:12:03 | 27.176 | 142.730 | 33 | 5.0 | X |  |  |
| 98:11:10 | 17:45:12 | -22.027 | -68.239 | 118 | 4.7 | X |  |  |
| 98:11:11 | 23:35:46 | 53.511 | -164.530 | 33 | 4.9 | X |  |  |
| 98:11:11 | 23:36:33 | 1.079 | -85.275 | 33 | 5.5 | X |  |  |
| 98:11:13 | 17:38:58 | -21.572 | -68.222 | 123 | 5.4 | X |  |  |
| 98:11:14 | 14:23:15 | 11.709 | 143.248 | 33 | 5.3 | X |  |  |
| 98:11:15 | 02:44:12 | -21.589 | -176.504 | 149 | 5.9 | X | X | X |
| 98:11:15 | 04:51:42 | -4.075 | -104.181 | 10 | 4.8 | X |  |  |
| 98:11:15 | 07:58:14 | 13.001 | 143.607 | 142 | 4.9 | X |  |  |
| 98:11:15 | 08:23:08 | -9.344 | -71.292 | 596 | 4.7 | X |  |  |
| 98:11:15 | 23:08:33 | 37.658 | 137.320 | 21 | 4.9 | X |  |  |
| 98:11:17 | 03:16:08 | -26.830 | -113.290 | 10 | 5.4 | X |  |  |
| 98:11:17 | 03:57:58 | 7.666 | -82.780 | 17 | 5.2 | X |  |  |
| 98:11:17 | 17:08:02 | -21.695 | -179.131 | 600 | 5.1 | X |  |  |
| 98:11:17 | 22:27:32 | 22.675 | 120.959 | 33 | 5.2 | X |  |  |
| 98:11:18 | 07:32:53 | -10.551 | 165.116 | 62 | 4.8 | X |  |  |
| 98:11:19 | 07:28:07 | -9.045 | -78.618 | 73 | 4.7 | X |  |  |
| 98:11:19 | 15:39:19 | 22.605 | 125.783 | 10 | 5.8 | X | X |  |
| 98:11:20 | 18:14:32 | -28.384 | -112.814 | 10 | 4.8 | X |  |  |
| 98:11:20 | 21:23:38 | -16.291 | -178.088 | 438 | 5.2 | X |  |  |
| 98:11:21 | 16:59:47 | 49.233 | 89.186 | 10 | 5.2 | X |  |  |
| 98:11:22 | 12:25:37 | 52.200 | 178.888 | 144 | 4.8 | X |  |  |
| 98:11:22 | 19:01:10 | 27.242 | 125.715 | 253 | 4.4 | X |  |  |
| 98:11:23 | 09:30:19 | -23.716 | -70.512 | 33 | 5.5 | X | X |  |
| 98:11:23 | 12:15:50 | -18.184 | -174.974 | 209 | 5.2 | X |  |  |
| 98:11:23 | 19:48:10 | 37.979 | 141.479 | 79 | 5.4 | X |  |  |
| 98:11:23 | 19:58:25 | 45.080 | 147.162 | 140 | 4.8 | X |  |  |
| 98:11:24 | 01:06:11 | -22.625 | -69.243 | 80 | 5.3 | X |  |  |
| 98:11:24 | 21:24:31 | -16.000 | -172.703 | 33 | 5.2 | X |  |  |
| 98:11:24 | 23:54:46 | -16.515 | -174.751 | 223 | 5.4 | X |  |  |
| 98:11:25 | 18:05:25 | -7.859 | 158.622 | 48 | 5.9 | X | X |  |
| 98:11:27 | 10:27:02 | -32.140 | -69.328 | 127 | 5.2 | X |  |  |
| 98:11:28 | 09:58:09 | -7.589 | -74.416 | 149 | 5.3 | X |  |  |
| 98:11:28 | 15:21:05 | -15.363 | 172.964 | 33 | 5.4 | X |  |  |

Table A.1: (Continued)

| $\begin{gathered} \text { Date } \\ \text { YY:MM:DD } \end{gathered}$ | Time hh:mm:ss.s | Latitude ${ }^{\circ} \mathrm{N}$ | Longitude ${ }^{\circ} \mathrm{E}$ | $\begin{array}{r} \text { Depth } \\ \mathrm{km} \end{array}$ | $\mathrm{m}_{6}$ | TTI-P | RF | SWS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 98:11:29 | 17:14:00 | 48.132 | 148.441 | 396 | 4.9 | X |  |  |
| 98:11:30 | 20:18:28 | 14.096 | 145.451 | 115 | 4.6 | X |  |  |
| 98:12:01 | 10:38:45 | 53.099 | -164.338 | 22 | 5.6 | X | X |  |
| 98:12:02 | 07:11:03 | -33.457 | -109.346 | 10 | 5.4 | X |  |  |
| 98:12:04 | 18:50:07 | 43.620 | -28.760 | 10 | 4.8 | X |  |  |
| 98:12:05 | 01:12:47 | 52.121 | -169.412 | 33 | 5.4 | X |  |  |
| 98:12:05 | 08:06:51 | 10.985 | -86.163 | 96 | 4.9 | X |  |  |
| 98:12:06 | 23:31:00 | -21.059 | -179.133 | 600 | 4.9 | X |  |  |
| 98:12:08 | 02:32:57 | 18.819 | -64.046 | 30 | 5.6 | X |  |  |
| 98:12:09 | 07:35:49 | 19.243 | 145.437 | 151 | 4.9 | X |  |  |
| 98:12:10 | 08:21:14 | -7.952 | -71.416 | 649 | 5.1 | X |  |  |
| 98:12:11 | 08:37:50 | -31.266 | -68.918 | 118 | 5.5 | X | X |  |
| 98:12:11 | 21:42:46 | -16.597 | -172.758 | 33 | 5.2 | X |  |  |
| 98:12:13 | 17:31:58 | 13.345 | -44.845 | 10 | 5.3 | X |  |  |
| 98:12:13 | 20:07:52 | 13.338 | -44.949 | 10 | 4.9 | X |  |  |
| 98:12:14 | 04:30:56 | 30.923 | 137.654 | 464 | 4.9 | X |  |  |
| 98:12:14 | 16:25:24 | -38.214 | -71.033 | 138 | 5.4 | X |  |  |
| 98:12:16 | 00:18:45 | 31.287 | 131.286 | 42 | 5.5 | X | X |  |
| 98:12:23 | 15:14:07 | 17.509 | -94.660 | 145 | 4.5 | X |  |  |
| 98:12:26 | 11:29:09 | 10.604 | -63.544 | 33 | 5.4 | X |  |  |
| 98:12:27 | 00:38:26 | -21.632 | -176.376 | 144 | 6.1 | X | X | X |
| 98:12:28 | 07:23:31 | 20.780 | -74.673 | 10 | 5.6 | X |  |  |
| 98:12:30 | 03:32:37 | -1.646 | -77.881 | 169 | 5.0 | X |  |  |
| 99:01:01 | 16:20:30 | 36.139 | 141.636 | 33 | 5.1 | X |  |  |
| 99:01:07 | 11:38:07 | 43.827 | 148.233 | 56 | 5.0 | X |  |  |
| 99:01:07 | 18:13:41 | 67.768 | 141.306 | 33 | 5.4 | X |  |  |
| 99:01:07 | 22:16:07 | -20.480 | -173.988 | 33 | 4.9 | X |  |  |
| 99:01:09 | 03:05:37 | 44.390 | 147.313 | 119 | 5.8 | X |  |  |
| 99:01:11 | 10:48:50 | 52.166 | 159.626 | 33 | 5.3 | X |  |  |
| 99:01:12 | 02:32:25 | 26.741 | 140.170 | 441 | 5.9 |  | X |  |
| 99:01:12 | 08:49:20 | -5.421 | 151.681 | 43 | 5.5 |  | X |  |
| 99:01:16 | 03:01:01 | 50.749 | -169.920 | 33 | 5.4 | X |  |  |
| 99:01:16 | 09:06:36 | 12.308 | 144.105 | 33 | 5.2 | X |  |  |
| 99:01:17 | 05:29:12 | 43.185 | 147.525 | 52 | 5.2 | X |  |  |
| 99:01:17 | 15:28:35 | 13.232 | 145.528 | 33 | 5.2 | X |  |  |
| 99:01:19 | 03:35:33 | -4.596 | 153.235 | 114 | 5.8 | X |  |  |
| 99:01:20 | 07:38:35 | 52.291 | 179.801 | 170 | 4.7 | X |  |  |

Table A.1: (Continued)

| $\begin{gathered} \text { Date } \\ \text { YY:MM:DD } \end{gathered}$ | $\begin{gathered} \text { Time } \\ \text { hh:mm:ss.s } \end{gathered}$ | Latitude ${ }^{\circ} \mathrm{N}$ | Longitude ${ }^{\circ} \mathrm{E}$ | $\begin{array}{r} \text { Depth } \\ \mathrm{km} \end{array}$ | $\mathrm{m}_{6}$ | TTI-P | RF | SWS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 99:01:20 | 13:43:48 | -17.752 | -178.930 | 600 | 4.6 | X |  |  |
| 99:01:20 | 23:10:04 | -14.801 | -75.865 | 33 | 5.2 | X |  |  |
| 99:01:21 | 22:02:16 | 38.649 | 142.904 | 33 | 5.3 | X |  |  |
| 99:01:24 | 00:37:04 | 30.618 | 131.086 | 33 | 6.1 | X | X |  |
| 99:01:24 | 07:01:58 | -21.132 | -174.659 | 33 | 5.7 | X | X |  |
| 99:01:24 | 13:15:52 | 54.476 | 161.458 | 33 | 5.3 | X |  |  |
| 99:01:25 | 10:37:13 | -18.026 | -178.445 | 640 | 5.0 | X |  |  |
| 99:01:25 | 18:19:16 | 4.461 | -75.724 | 17 | 5.9 | X | X |  |
| 99:01:25 | 22:40:16 | 4.370 | -75.682 | 10 | 5.5 | X |  |  |
| 99:01:26 | 12:30:49 | -20.515 | -174.207 | 41 | 5.5 | X | X |  |
| 99:01:26 | 16:18:30 | -17.762 | -178.809 | 596 | 4.7 | X |  |  |
| 99:01:27 | 08:20:31 | 48.376 | 156.213 | 71 | 5.2 | X |  |  |
| 99:01:27 | 10:13:53 | 6.711 | -82.678 | 10 | 5.4 | X |  |  |
| 99:01:28 | 08:10:05 | 52.886 | -169.123 | 67 | 6.3 | X | X |  |
| 99:01:30 | 03:51:05 | 41.674 | 88.463 . | 23 | 5.9 | X | X |  |
| 99:01:31 | 05:07:13 | 43.157 | 46.841 | 33 | 5.3 | X |  |  |
| 99:01:31 | 16:51:51 | 37.148 | 141.339 | 44 | 5.3 | X |  |  |
| 99:01:31 | 19:29:11 | 43.455 | 146.960 | 33 | 5.6 | X | X |  |
| 99:02:04 | 19:43:14 | 1.083 | -30.545 | 10 | 4.9 | X |  |  |
| 99:02:05 | 11:39:45 | -12.616 | 166.966 | 213 | 5.7 | X | X |  |
| 99:02:05 | 14:37:53 | 47.505 | 147.158 | 407 | 5.4 | X |  |  |
| 99:02:06 | 13:36:12 | 53.561 | 160.402 | 38 | 5.2 | X |  |  |
| 99:02:06 | 17:45:24 | 19.200 | 121.265 | 33 | 5.4 | X |  |  |
| 99:02:06 | 21:47:59 | -12.853 | 166.697 | 90 | 6.3 | X | X | X |
| 99:02:09 | 10:38:47 | -15.543 | -173.343 | 33 | 5.0 | X |  |  |
| 99:02:10 | 09:22:35 | -21.706 | -178.842 | 549 | 5.1 | X |  |  |
| 99:02:12 | 17:44:48 | 44.474 | 149.678 | 33 | 5.6 | X |  |  |
| 99:02:14 | 11:22:37 | 44.506 | 149.710 | 33 | 5.7 | X | X |  |
| 99:02:14 | 21:12:24 | -15.507 | 167.996 | 10 | 5.9 | X | X |  |
| 99:02:15 | 09:41:52 | -18.065 | -178.530 | 624 | 4.7 | X |  |  |
| 99:02:16 | 14:56:40 | 15.647 | -87.130 | 10 | 5.2 | X |  |  |
| 99:02:17 | 21:58:54 | -21.143 | -70.040 | 33 | 5.6 | X | X |  |
| 99:02:19 | 19:10:00 | 85.573 | 87.037 | 10 | 5.1 | X |  |  |
| 99:02:21 | 18:14:37 | 43.214 | 46.825 | 65 | 5.1 | X |  |  |
| 99:02:22 | 08:02:11 | 86.278 | 73.394 | 10 | 5.2 | X |  |  |
| 99:02:23 | 12:23:43 | 53.785 | 171.134 | 24 | 4.9 | X |  |  |
| 99:02:23 | 18:56:56 | -20.804 | -174.071 | 36 | 5.3 | X |  |  |

Table A.1: (Continued)


Table A.1: (Continued)

| $\begin{gathered} \text { Date } \\ \text { YY:MM:DD } \end{gathered}$ | Time hh:mm:ss.s | Latitude ${ }^{\circ} \mathrm{N}$ | Longitude ${ }^{\circ} \mathrm{E}$ | $\begin{array}{r} \text { Depth } \\ \mathrm{km} \end{array}$ | $\mathrm{m}_{6}$ | TTI-P | RF | SWS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 99:04:04 | 10:37:00 | 16.097 | -97.341 | 33 | 5.1 | X |  |  |
| 99:04:06 | 00:08:22 | 39.400 | 38.307 | 10 | 5.1 | X |  |  |
| 99:04:06 | 04:51:05 | 24.451 | -46.374 | 10 | 5.3 | X |  |  |
| 99:04:08 | 13:10:34 | 43.607 | 130.350 | 566 | 6.4 | X | X |  |
| 99:04:09 | 12:16:01 | -26.354 | 178.221 | 621 | 5.5 | X | X |  |
| 99:04:11 | 19:15:18 | 18.452 | 145.154 | 513 | 4.5 | X |  |  |
| 99:04:12 | 01:54:16 | -22.864 | -70.466 | 40 | 4.9 | X |  |  |
| 99:04:13 | 10:38:48 | -21.422 | -176.460 | 164 | 6.4 | X | X | X |
| 99:04:14 | 07:25:04 | 6.772 | -72.946 | 161 | 4.8 | X |  |  |
| 99:04:17 | 00:56:25 | 19.248 | -155.489 | 11 | 5.6 | X | X |  |
| 99:04:17 | 07:20:55 | 38.213 | 75.518 | 137 | 4.9 | X |  |  |
| 99:04:17 | 08:17:58 | 36.038 | 21.683 | 40 | 4.7 | X |  |  |
| 99:04:18 | 23:14:19 | 34.192 | 139.460 | 129 | 4.8 | X |  |  |
| 99:04:19 | 09:12:47 | 50.888 | 156.423 | 118 | 5.4 | X |  |  |
| 99:04:20 | 01:46:17 | -18.791 | -177.957 | 586 | 5.1 | X |  |  |
| 99:04:20 | 19:04:08 | -31.888 | -179.040 | 96 | 6.2 | X |  | X |
| 99:04:23 | 18:56:26 | 13.123 | 145.142 | 53 | 5.5 | X | X |  |
| 99:04:24 | 08:45:16 | -18.043 | -178.449 | 568 | 5.2 | X |  |  |
| 99:04:25 | 09:10:44 | -31.798 | -69.255 | 116 | 5.3 | X |  |  |
| 99:04:25 | 12:27:05 | 36.441 | 140.473 | 82 | 5.3 | X |  |  |
| 99:04:26 | 07:07:02 | 53.965 | 159.190 | 127 | 4.8 | X |  |  |
| 99:04:26 | 13:20:07 | 85.672 | 84.830 | 10 | 5.2 | X |  |  |
| 99:04:26 | 18:17:26 | -1.648 | -77.783 | 173 | 5.6 | X | X |  |
| 99:04:28 | 08:47:55 | 45.464 | 26.183 | 156 | 5.1 | X |  |  |
| 99:04:28 | 08:47:55 | 45.464 | 26.183 | 156 | 5.1 | X |  |  |
| 99:04:29 | 07:46:08 | 28.867 | 131.148 | 33 | 5.8 | X | X |  |
| 99:04:29 | 15:01:18 | -26.611 | -114.396 | 10 | 5.1 | X |  |  |
| 99:04:30 | 03:30:38 | 44.181 | 20.071 | 14 | 5.0 | X |  |  |
| 99:05:02 | 16:13:50 | 56.794 | -34.268 | 10 | 5.3 | X |  |  |
| 99:05:02 | 20:39:26 | 29.083 | 131.212 | 33 | 5.1 | X |  |  |
| 99:05:05 | 22:41:30 | 14.364 | -94.673 | 33 | 5.8 | X | X |  |
| 99:05:08 | 05:11:54 | 11.509 | -86.793 | 62 | 4.9 | X |  |  |
| 99:05:08 | 19:44:35 | 45.449 | 151.630 | 63 | 6.2 | X | X |  |
| 99:05:08 | 22:12:45 | 14.214 | -91.945 | 39 | 5.3 | X |  |  |
| 99:05:10 | 08:47:25 | -30.397 | -69.163 | 65 | 5.2 | X |  |  |
| 99:05:10 | 20:33:02 | -5.159 | 150.880 | 138 | 6.5 | X | X | X |
| 99:05:12 | 17:59:22 | 43.032 | 143.835 | 103 | 5.9 | X | X |  |

Table A.1: (Continued)

| $\begin{gathered} \text { Date } \\ \text { YY:MM:DD } \end{gathered}$ | Time hh:mm:ss.s | Latitude ${ }^{\circ} \mathrm{N}$ | Longitude ${ }^{\circ} \mathrm{E}$ | $\begin{array}{r} \text { Depth } \\ \mathrm{km} \end{array}$ | $\mathrm{m}_{6}$ | TTI-P | RF | SWS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 99:05:16 | 00:51:20 | -4.751 | 152.486 | 74 | 6.0 | X | X |  |
| 99:05:16 | 15:25:53 | -2.642 | 138.217 | 59 | 6.1 | X |  |  |
| 99:05:17 | 10:07:56 | -5.165 | 152.877 | 27 | 5.5 |  | X |  |
| 99:05:18 | 20:20:16 | 85.632 | 86.146 | 10 | 5.1 | X |  |  |
| 99:05:19 | 18:40:00 | 15.896 | -92.929 | 95 | 4.7 | X |  |  |
| 99:05:21 | 04:31:25 | -22.886 | -68.520 | 130 | 4.7 | X |  |  |
| 99:05:25 | 16:42:05 | -27.931 | -66.934 | 169 | 5.3 | X |  |  |
| 99:05:26 | 23:56:32 | 85.605 | 86.526 | 10 | 5.1 | X |  |  |
| 99:05:28 | 04:52:55 | 12.568 | -87.377 | 86 | 5.3 | X |  |  |
| 99:05:30 | 15:56:45 | 55.796 | 110.030 | 10 | 5.3 | X |  |  |
| 99:06:02 | 07:34:41 | -20.858 | -179.002 | 644 | 5.0 | X |  |  |
| 99:06:07 | 16:10:33 | 73.017 | 5.187 | 10 | 5.3 | X |  |  |
| 99:06:07 | 16:35:46 | 73.077 | 5.453 | 10 | 5.2 | X |  |  |
| 99:06:08 | 12:04:00 | 15.040 | -60.421 | 55 | 5.3 | X |  |  |
| 99:06:09 | 00:02:04 | -19.281 | -173.557 | 33 | 5.0 | X |  |  |
| 99:06:09 | 07:07:31 | 49.316 | 158.396 | 33 | 5.2 | X |  |  |
| 99:06:10 | 09:08:13 | 56.137 | -161.609 | 172 | 4.7 | X |  |  |
| 99:06:10 | 15:07:21 | 36.237 | 71.212 | 112 | 5.2 | X |  |  |
| 99:06:11 | 07:50:15 | 37.560 | 21.110 | 58 | 4.8 | X |  |  |
| 99:06:13 | 22:48:21 | 13.872 | -50.152 | 10 | 5.0 | X |  |  |
| 99:06:15 | 20:42:05 | 18.386 | -97.436 | 70 | 6.4 | X | X |  |
| 99:06:16 | 18:35:59 | -17.037 | -173.362 | 75 | 5.6 |  | X |  |
| 99:06:21 | 17:43:04 | 18.324 | -101.539 | 69 | 6.0 | X | X |  |
| 99:06:26 | 22:05:28 | -17.956 | -178.187 | 590 | 5.3 | X |  |  |
| 99:06:29 | 05:50:89 | -9.468 | 147.854 | 33 | 5.8 |  | X |  |
| 99:06:29 | 10:55:11 | -15.710 | -72.496 | 118 | 5.1 | X |  |  |
| 99:06:29 | 23:18:05 | 36.622 | 71.353 | 189 | 5.9 | X | X |  |
| 00:04:23 | 09:27:22 | -28.250 | -62.890 | 603 | 6.9 |  |  | X |
| 00:04:23 | 17:01:17 | -28.280 | -62.820 | 610 | 6.1 |  |  | X |

## Appendix B

## Scattering potentials

This appendix gives the full expressions for the scattering potentials,

$$
\begin{equation*}
f^{q}(\mathbf{x}, \theta)=\sum_{l=1}^{3} W_{l}^{q}(\theta) \Delta m_{l}(\mathbf{x}) \tag{B.1}
\end{equation*}
$$

which include the $\theta$-dependent radiation patterns $W_{l}^{q}(\theta)$ and material property perturbations $\Delta m_{l}(\mathbf{x})=(\Delta \alpha / \alpha, \Delta \beta / \beta, \Delta \rho / \rho)$. These expressions are derived in Paper I for different scattering modes:

$$
\begin{align*}
& f^{1,3}(\mathbf{x}, \theta)= \rho^{0}\left(2 \frac{\Delta \alpha}{\alpha^{0}}+\frac{\Delta \beta}{\beta^{0}}\left(2\left(\frac{\beta^{0}}{\alpha^{0}}\right)^{2}(\cos 2 \theta-1)\right)+\right. \\
&\left.\frac{\Delta \rho}{\rho^{0}}\left(1+\cos \theta+\left(\frac{\beta^{0}}{\alpha^{0}}\right)^{2}(\cos 2 \theta-1)\right)\right)  \tag{B.2}\\
& f^{2,4}(\mathbf{x}, \theta)=\rho^{0}\left(\frac{\Delta \beta}{\beta^{0}}\left(2 \frac{\beta^{0}}{\alpha^{0}} \sin 2 \theta\right)+\frac{\Delta \rho}{\rho^{0}}\left(\sin \theta+\frac{\beta^{0}}{\alpha^{0}} \sin 2 \theta\right)\right)  \tag{B.3}\\
& f^{5}(\mathbf{x}, \theta)=\rho^{0}\left(\frac{\Delta \beta}{\beta^{0}}\left(2 \frac{\beta^{0}}{\alpha^{0}} \sin 2 \theta\right)+\frac{\Delta \rho}{\rho^{0}}\left(\sin \theta+\frac{\beta^{0}}{\alpha^{0}} \sin 2 \theta\right)\right)  \tag{B.4}\\
& f^{6}(\mathbf{x}, \theta)=\rho^{0}\left(\frac{\Delta \beta}{\beta^{0}}(2 \cos 2 \theta)+\frac{\Delta \rho}{\rho^{0}}(\cos \theta+\cos 2 \theta)\right)  \tag{B.5}\\
& f^{7}(\mathbf{x}, \theta)=\rho^{0}\left(\frac{\Delta \beta}{\beta^{0}}(2 \cos \theta)+\frac{\Delta \rho}{\rho^{0}}(1+\cos \theta)\right) \tag{B.6}
\end{align*}
$$

