# TRANSIENT EFFECTS IN OILFIELD CEMENTING FLOWS 

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## Abstract

In this thesis we study instabilities and mud channel removal during the primary cementing of an oil well. This process involves displacement of a sequence of non-Newtonian fluids along a narrow eccentric annulus. Previously, this has been modelled as a pseudo-steady process using a Hele-Shaw approximation. In such models, the stream function is governed by a steady elliptic problem and the fluids, (modelled via a concentration equation), simply advect along the annulus. It has been shown that for certain rheological and physical parameters, the displacement front will advances much faster on the wide side of the annulus than on the narrow side. In extreme cases the displacement front does not advance at all on the narrow side of the annulus and a static mud channel results as the finger advances up the wide side. In this thesis we consider whether the interface of a progressively advancing finger will remain stable to small perturbations. There is in fact experimental evidence that interfacial instabilities can occur in this situation. We find that the interface is in fact stable whenever there is a static mud channel on the narrow side of the annulus. Consequently, we also investigate how a mud channel might be removed by pulsation of the flow rate. Study of these two phenomena cannot be undertaken with the pseudo-steady framework. Therefore, we extend this model to flows that are fully transient. The transient model consists of a nonlinear evolution equation for the stream function.

In chapter 3 we show that this transient model is in fact well-posed. In chapter 4 we study stability of multi-layer parallel flows, i.e. long fingers. If both fluids are yielded at the interface, instabilities may arise for different combinations of the 14 dimensionless parameters. These instabilities are related to a jump in tangential velocity at the interface and do not appear to have been identified before. In chapter 5 we investigate the case where a static mud channel develops on the narrow side of the annulus. Our stability theory predicts only linear stability. We therefore study the effects of a finite pulsation of the flow rates via numerical simulation. It seems that if we perturb the flow from the beginning of the displacement, the transient model fully captures the effects of the perturbation and the width of the mud channel is reduced. The pseudo-steady velocity model does not report any significant changes with respect to the results using a constant flow rate. If however, pulsation is applied after the mud channel has already formed, the removal of the mud channel will be unsuccessful.

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To my parents and my wife

## Chapter 1

## Introduction

The aim of this thesis it to study instabilities and mud channel removal during the primary cementing of an oil well. This process involves displacement of a sequence of non-Newtonian fluids along a narrow eccentric annulus. It has been shown that for certain rheological and physical parameters of the fluids, the displacement front will advance much faster on the wide side of the annulus than on the narrow side. In extreme cases the displacement front does not advance at all on the narrow side of the annulus and a static mud channel results as the finger advances up the wide side. We consider whether the interface of a progressively advancing finger will remain stable to small perturbations and also investigate how a mud channel might be removed by pulsation of the flow rate, which is sometimes used in the cementing industry.

### 1.1 Primary cementing process description

The primary cementing process proceeds as follows: After drilling the well to the desired depth, Figure 1.1a, the drillpipe is removed and a large steel casing is placed down to the bottom of the well, leaving a gap of $\approx 2 \mathrm{~cm}$ between the outside of the tube and the inside of the wellbore, i.e. an annulus. At this time, the drilling mud is still in the wellbore, Figure 1.1b. To remove the drilling mud, the cement is first pumped down the inside of the steel casing, Figure 1.1c, and emerges at the end where it then flows into the narrow annular gap between the casing and the well bore, displacing the drilling mud, Figure 1.1d. The cement later hardens into an hydraulic seal, Figure 1.1e. Without the hydraulic seal, well productivity can be reduced significantly and fluids can also leak up to the surface, where they may become an environmental or safety hazard. The fluids used in cementing, (drilling muds, spacer fluids and cement slurries), are


Figure 1.1: Schematic of the primary cementing process: a) Drilling to desired depth. b) Placement of the steel casing, well bore still full of drilling mud. c) Cement is pumped down inside of the steel casing. d) Cement displaces drilling mud through the annular region. e) Cement hardens into hydraulic seal for zone isolation .
generally characterised as viscoplastic shear-thinning fluids. In practice a sequence of fluids is pumped along an annulus (spacer fluids and cement slurries), each fluid displacing the one in front, and the annular geometry changes slowly in the axial direction. However, the fluid volumes pumped are relatively large and the essential dynamics of the displacement may be studied by considering what happens between any 2 fluids on an annular section of constant geometry. Note that Figure 1.1 presents an ideal case. Ușually the angle of inclination of the well is different from zero. Centralizers are attached to the steel casing to prevent the casing from slumping to the lower side of the annulus. Even though they prevent the slumping, there are not always enough centralizers fitted and/or the stiffness of the centralizers is not designed well. Thus, the annulus nearly always is eccentric. Another problem is that a full removal of the drilling mud is not always achieved. A mud channel is frequently left along the narrow part of the annulus. This dehydrates during setting of the cement and the dried mud forms a porous conduit along the length of the well. Therefore, complete zonal isolation is not achieved.

### 1.2 Non-Newtonian Fluids

As stated in the previous section, the fluids used in cementing are generally characterized as visco-plastic shear thinning fluids. In this section we give a short summary of the constitutive equations which characterize these fluids. For a more complete review see [7, 15, 37].

Let us assume shear flow of a liquid between two parallel plates which its velocity vector can be expressed as

$$
\boldsymbol{u}=(\dot{\gamma} y, 0,0)
$$

where $\dot{\gamma}$ is the shear rate. In the case of a Newtonian fluid the stress distribution for shear flow is given in the following form:

$$
\begin{equation*}
\sigma_{y x}=\eta \dot{\gamma} \quad \sigma_{x z}=\sigma_{y z}=0 \quad \sigma_{x x}-\sigma_{y y}=0 \quad \sigma_{y y}-\sigma_{z z}=0 . \tag{1.1}
\end{equation*}
$$

A fluid is called Newtonian if satisfies (1.1) and if $\eta$ does not vary with shear rate and is constant with respect to the time of shearing and the stress in the liquid falls to zero after shearing is stopped. Any liquid that does not satisfies any of these requirements is called non-Newtonian. We will concentrate in fluids for which the difference in the normal stresses is zero, i.e. inelastic fluids, and for which the shear viscosity depends on the shear rate but not explicitly on time.

### 1.2.1 Time independent non-Newtonian fluids

We can subdivide this type of fluids into 3 groups,

- Shear thinning (or pseudo-plastic) fluids, for which the viscosity decreases with increasing shear rate.
- Shear thickening (or dilatant) fluids, for which the viscosity increases with increasing shear rate.
- Yield stress (or visco-plastic) fluids, for which a finite shear stress is required to initiate the flow.


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These fluids are illustrated in Figure 1.2, and we now give some examples of the common fluid models.


Figure 1.2: Schematic behaviour of shear stress vs. shear rate for time independent nonNewtonian fluids.

For many fluids, as the shear rate approaches zero or infinity the value of $\eta$ approaches constant values, say $\eta_{0}$ and $\eta_{\infty}$, i.e. these are 2 Newtonian limits. Two models that contain this behaviour are:

$$
\begin{equation*}
\frac{\eta-\eta_{\infty}}{\eta_{0}-\eta_{\infty}}=\frac{1}{\left(1+(K \dot{\gamma})^{m}\right)} \tag{1.2}
\end{equation*}
$$

which is called the Cross model, and

$$
\begin{equation*}
\frac{\eta-\eta_{\infty}}{\eta_{0}-\eta_{\infty}}=\frac{1}{\left(1+(K \dot{\gamma})^{2}\right)^{m / 2}} \tag{1.3}
\end{equation*}
$$

which is called the Carreau model. Here $K$ and $m$ are empirically fitted constant rheological parameters.

If we assume $\eta \ll \eta_{0}$ and $\eta \gg \eta_{\infty}$, at leading order and by a simple redefinition of parameters, the Cross model reduces to

$$
\begin{equation*}
\eta=K \dot{\gamma}^{n-1} \tag{1.4}
\end{equation*}
$$

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which is the well known power law model. If instead we assume that $\eta \ll \eta_{\infty}$ then again, by a redefinition of parameters, the Cross model reduces to

$$
\begin{equation*}
\eta=\eta_{\infty}+K \dot{\gamma}^{n-1} \tag{1.5}
\end{equation*}
$$

which is called Sisko model. If $n=0$ in Sisko's model and we redefine the parameters we obtain the Bingham model,

$$
\begin{equation*}
\eta=\mu+\frac{\tau_{Y}}{\dot{\gamma}} \tag{1.6}
\end{equation*}
$$

where $\mu$ is called the plastic viscosity and $\tau_{Y}$ is the yield stress. The Bingham model is the simplest model that describes visco-plastic behaviour. Other widely used visco-plastic models are:

$$
\begin{equation*}
\sqrt{\tau}=\sqrt{\tau_{Y}}+\sqrt{K \dot{\gamma}} \tag{1.7}
\end{equation*}
$$

which is called the Casson model, and is given in terms of the shear stress, and finally

$$
\begin{equation*}
\eta=K \dot{\gamma}^{n-1}+\frac{\tau_{Y}}{\dot{\gamma}} \tag{1.8}
\end{equation*}
$$

which is called the Herschel-Bulkley model. Note that if $n=1$ and $\tau_{Y}=0$ we recover the Newtonian model, if $n=1$ and $\tau_{Y} \neq 0$ we recover the Bingham model, and if $n<1$ and $\tau_{Y}=0$ we recover the power law model. The Herschel-Bulkley model will be the one used for the rest of the thesis.

### 1.3 Literature Review

### 1.3.1 Industrial background

There is an extensive technical literature on primary cementing, see [21] for a review, as this is a large industrial business. Only a part of this literature will be reviewed in this section. Early development of cementing techniques dates back to 1911, we refer to the reader to [65], which gives a good review of the techniques used up to the 1980's. Next we will give a small overview of these techniques.

In the 1940 's a good amount of attention was given to the cementing process. Investigations on the effectiveness of water pre-flushes ahead of cement, pipe movement, well centered casing,
mud thinning prior to cementing and turbulent flows were carried out. In the 1960's mud displacements studies with cement slurries in turbulent flow proved to be very effective, this gave to rise to the development of the "scavenger slurry" which is a low density, highly dispersed cement slurry which is pumped before the tail slurry, i.e. the cement that is expected to provide casing support and a hydraulic seal. In the 1970's researchers started looking at how to get a good cementing job in deviated holes. Studies on the effects of centralizers and pipe movement were carried out. In the 1980's attention was given to the design of a mud-cement compatible spacer systems that were capable of transitioning to turbulence at reasonable pump rates.

Since the late 60's analytical and experimental studies have been carried out in order to understand better when a good displacement takes place. For example, McLean et.al., [52], considered an analytical model which was later tested against experimental data. They considered the division of the annulus into sectors each one treated as an equivalent sector of a concentric annulus and assumed that the pressure gradient was identical in each sector, (effectively a crude early finite element method!). From this approach they found a critical value for the yield stress of the displacing fluid, in this case cement, for which the pressure gradient in each sector was large enough to initiate flow in all the annular region. This prediction seems to over estimate the actual experimental value of the yield stress of the cement. They claimed that this is due to the fact of neglecting the effect of interfacial shear stresses at the cement mud interface. Later, Clark \& Carter, [13] studied the effect of interfacial stresses. They found that these effects can be very important in initiating mud flow and claimed that adverse buoyancy effects could be compensated by displacing the cement at higher rates.

Another example of an early analytical model can be found in [9], where the authors approximated the displacement zone as a narrow gap annulus between concentric cylinders, or equivalently the region between two parallel plates. Their model is 1-D pseudo-steady, i.e. they calculate the 1-D steady state solution for the velocity at fixed time. Through integration of the velocity along the interface and determining its position, they are able to find the volume of fluid displaced as a function of time. On the other hand, Courtier et.al., [14] actually considered the effect of eccentricity using the basic slot model, which consists in considering independent
rectangular slots of varying heights at different azimuthal positions. All the analytical models were tested against available experimental data. For the experimental results we refer to the reader to $[11,50,69,70,93]$. There have also been a large number of field-based case studies, which we do not review here.

For a good displacement in the annular region to happen, almost all laboratory and field studies agree that:

- The displacing fluid should be heavier than the displaced one for near vertical annulus.
- The use of centralizers to reduce the eccentricity of the annulus because one of the main causes of mud channeling is the casing eccentricity.
- Fluids with higher viscosity and yield stress are better able to displace fluids with lower viscosity and yield stress.

Most industrial literature agrees that if turbulence is achieved in the annulus it's the best tool to remove the mud from the well. Unfortunately, this is not always attainable in the field. For example the flow rates required may exceed capacity of the pumps at the rig site.

### 1.3.2 Displacement flow in different geometries

We present a brief review of some of the work done in simulating, either analytically, numerically or in a lab setting, the displacement of one miscible fluid by another in different geometries.

In 1993, Tehrani et.al., [78], presented experimental results on the displacement of one fluid by another in an eccentric annulus, using both Newtonian and non-Newtonian fluids. Their study shows that the interface between the the two fluids with similar rheologies but different densities is stable as long as the denser fluid lies below the interface. If the interface becomes close to vertical, a bad displacement and a long finger develops. In this case, small perturbations in the flow field may trigger gravity driven instabilities. Azimuthal instabilities also can occur. The intensity of these instabilities is related to density difference and flow rate. These interfacial instabilities are one of the main subjects of the thesis and will be investigated in chapter 4. These results lead to the work reported by Szabo \& Hassager in [73]. The fluids they consider
are Newtonian, their simulations support earlier experimental findings. They solved the full 3-D problem using a mixed Lagrangian-Eulerian method. They solve the equations of motion plus an equation for the moving grid to track the interface between the fluids.

Some experimental work concerning the downward or upward vertical displacements of two miscible Newtonian fluids in a vertical pipe can be found in [5, 67]. Scoffoni et.al., [67] considered only the downward displacement of a heavier and more viscous fluid by a lighter and less viscous fluid. A long finger develops along the length of the pipe and at high rates the finger destabilizes showing either an axisymmetric pattern or an asymetric corkscrew mode. Baasubramaniam et.al., [5] considered the opposite case, i.e. the displacement of lighter and less viscous fluids by a heavier and more viscous fluids in downward or upward displacements. For the case of downward displacements they also observed interfacial instabilities. They found that a long finger develops along the length of the duct. It has been shown, [40] that the interface between the fluids is destabilized and thus fingering occurs when a more viscous fluid is displaced by a less viscous one. This is due to the advection of the initial flat interface by the nonuniform velocity profile within the pipe. For downward displacements they also observed interfacial instabilities. For the upward displacements the interface was stable although a "spike" formed at the tip of the finger.

## Eccentric Hele-Shaw model for annular displacements

Bittleston et.al., [10], developed an eccentric Hele-Shaw model for the displacement of one fluid by another in an annular geometry, an oil well. The fluids considered satisfy the HerschelBulkley model. Because the annulus geometry is long and thin they derive a 2-D model of the bulk fluid motions by averaging across the gap and scaling the Navier-Stokes equations with the annular gap width, keeping only leading order terms. The model consists of an elliptic partial differential equation for the stream function and an advection equation for the concentration field. In fact the concentration equation field will determine the position of the interface between the fluids. In this thesis, we will extend this model to its transient version. In this way, we will be able to study interfacial instabilities when a long finger develops along the annulus.

A longer description of the model will be given in chapter 2. This model was further studied by Pelipenko \& Frigaard in [57]. They showed that in fact the model is well-posed. Also, for certain fluid properties, a steady state solution exists, i.e. an interface that advances steadily along the annulus at the mean pumping speed. These solutions can be found analytically for concentric annuli and annuli with small eccentricities via a perturbation method. Their approach is also numerical, see [58]. They solve the system of equations using an augmented Lagrangian Algorithm, see [26, 30, 31]. This method fully describes unyielded regions of the flow if any. In [59], by the means of a lubrication approximation, Pelipenko \& Frigaard deduce conditions, based on certain combinations of rheological and physical parameters, for when a good displacement takes place or not. They start from the assumption that a long finger has already developed along the annulus and assume pseudo-parallel flow away from the displacement front. By doing this they are able to predict the interface speed on both the wide and narrow side of the annulus and thus determine when the wide side interface moves faster or slower than that on the narrow side. This criterion defines whether or not a good displacement takes place, or whether the displacement front fingers along the well.

### 1.3.3 Displacements in a Hele-Shaw cell

It is well known that the motion of a Newtonian fluid between two fixed parallel plates which are sufficiently close together (Hele-Shaw cell) is

$$
\begin{equation*}
u=-\frac{1}{2 \mu} \frac{\partial p}{\partial x} z(d-z) \text { and } v=-\frac{1}{2 \mu} \frac{\partial p}{\partial y} z(d-z) \tag{1.9}
\end{equation*}
$$

where $z$ is the coordinate perpendicular to the plates and $d$ is the thickness of the gap between the plates, see [8]. Averaging across the gap we get an expression for the gap-averaged velocity , given by:

$$
\begin{equation*}
\bar{u}=-\frac{d^{2}}{12 \mu} \nabla p \tag{1.10}
\end{equation*}
$$

which is in fact the velocity field of a Newtonian fluid flowing through a porous medium of permeability $k=\frac{d^{2}}{12}$. Thus, a direct analogy between the displacement flows in a Hele-Shaw cell and in a porous medium can be demonstrated. Note that (1.10) is Darcy's law for a Newtonian fluid, we are interested in non-Newtonian flows, thus a nonlinear version of Darcy's
law has to be derived. In this section we present various models used in the literature.

## Hele-Shaw displacements for Newtonian Fluids

We commence by describing the classical study of Saffman \& Taylor, [64], regarding these type of flows. They studied the displacement of a viscous fluid by another in a Hele-Shaw cell. They found that the stability of the interface depends on the direction of the motion of the fluids, i.e. if the less viscous fluid is displacing the more viscous one, the thinner fluid penetrates the thicker fluid in the form of a long fingers, regularly spaced along the planar front. These fingers themselves may destabilize and form sub-fingers, and the whole process may repeat leading to fractal patterns of fingers. The displacement is ineffective and the interface is said to be unstable. This phenomenon has been termed "viscous fingering". If the more viscous fluid displaces the less viscous one, the interface will be stable. They also found that this is true whatever the density of the fluids is, provided that the velocity is large. Since then, several studies can be found in the literature concerning Hele-Shaw displacements and we review only some of them, for an extensive review of the subject up to the mid 80 's we refer the reader to [40]. Although this field of work involves both immsicible and miscible displacements, we concentrate in the latter ones which are the type of flows considered in the thesis.

In 1986 Tan \& Homsy, [74] studied the linear stability of a miscible displacement in a porous medium for rectilinear flow. They assume that the fluids have the same densities and that the viscosity varies exponentially across the interface depending on the concentration of the fluid on top. Quasi-steady state approximation is considered because the dispersive concentration profile is time dependent. They showed that the linear displacement process is unstable for an unfavorable mobility ratio, i.e. viscosity of the displacing fluid is less than the viscosity of the displaced one. Thus the results are in agreement with the theory for immiscible fluids. Later, they performed numerical simulations, [75], to investigate the nonlinear effects of these instabilities. They found that once a finger becomes large enough the tip of the finger becomes unstable and splits. Lajeunesse et.al., $[46,47]$, considered the downward displacement of a fluid by a lighter and less viscous one on a vertical Hele-Shaw cell. It seems that for a mobility ration
greater than a critical value of 1.5 , there exists a critical velocity for which the interface of the finger becomes unstable leading to a 3-D pattern involving regularly spaced fingers.

Density driven unstable flows on Hele-Shaw cells have been also considered. Fernandez et.al., [25], studied density driven instabilities between miscible fluids in an horizontal Hele-Shaw cell, the fluids considered had equal viscosities and clearly, the heavier fluid was on top of the lighter fluid. They found that when the Rayleigh number, Ra, is small, the Rayleigh number being the ratio of convective and diffusive transport in the concentration (note that in the presence of net flow this number is just the Péclet number), diffusive effects are important and fingers with a wavelength $R a^{-1}$ may appear. For the opposite case, $R a$ large, diffusive effects are negligible and a finger with a wavelength of the order of the gap width develops. Later, Goyal et.al., [36] conducted similar experiments with the difference that the fluids had different viscosities. They found that the instability is reduced if the viscosity is increased in either fluid. Another interesting finding was that the instability also depends on the thickness of the interface (diffuse layer).

## Hele-Shaw displacements for non-Newtonian fluids

As stated before, in order to study these type of flows, we require a nonlinear version of Darcy's law, due to the fact that now the viscosity of the fluid is shear rate dependent. Barenblatt et.al. derived this relation for porous media flows, see [6]. They described the flow of a visco-plastic fluid through a porous medium by the following nonlinear Darcy's law with a limiting pressure. gradient,

$$
\begin{array}{lll}
\bar{u}=-\frac{k}{\mu}\left(1-\frac{G}{|\nabla p|}\right) \nabla p & & |\nabla p|>G, \\
\bar{u}=0 & & |\nabla p| \leq G \tag{1.12}
\end{array}
$$

where G is the limiting pressure gradient defined as,

$$
G=C \tau_{y} k^{-1 / 2}
$$

with $C$ being a dimensionless constant. In this 2-D system can be reduced to a single equation for the stream function if desired. Similar approaches to this type of nonlinear Darcy's law
can be found in the literature, see for example [4, 16, 34, 44, 53]. Actually, the model derived in [10], which is the model we follow and extend in this thesis, uses a nonlinear Darcy's law with a limiting pressure gradient as well. Having stated this, we review some of the work that has been done concerning non-Newtonian fluid displacements in a Hele-Shaw cell or a porous medium.

Wilson, [84], considered the Saffman-Taylor problem for a power law fluid displaced by air. By means of linear stability analysis he found that the growth rate of the instability is $O\left(n^{-1 / 2}\right)$, where $n$ is the power law index. Later Azaiez \& Singh, [4] considered miscible displacements of shear thinning fluids by a Newtonian fluid or vice versa. Their approach is the same as the one reported by Tan \& Homsy, [74] for Newtonian fluids, the quasi-steady-state approach. They found that for the $\mathrm{nN}-\mathrm{N}$ case, (meaning non-Newtonian-Newtonian displacement), the displacement can be unstable even for favorable mobility ratio, this was expected due to the shear thinning behaviour of the fluid. For the N-nN case they found similar results as for the $\mathrm{N}-\mathrm{N}$ case. Studies in which both fluids are non-Newtonian are presented below.

Pascal studied the classical planar interfacial instabilities in porous media for immiscible viscoplastic fluids, $[53,54,55]$. He used the Muskat approach, in which one assumes a long finger that extends ahead of the moving interface, if the velocity of the fluid within the finger is faster than the velocity of the fluid outside, the interface is unstable. By doing this he showed that in contrast with what the theory of displacement flows in a porous medium for Newtonian fluids says, even with the presence of an unfavorable mobility ratio, i.e. less viscous fluid displacing more viscous fluid, a positive difference of yield stresses will aid the displacement. By positive we mean that the displacing fluid should have a higher yield stress than the displaced one for this to happen. He also considered the effects of inclination of the media, and showed that gravity forces tend to stabilize the interface as long as the heavier fluid is under the lighter one. In 1999 Coussot, [16], considered the Saffman-Taylor problem involving two yield stress fluids. He found the same characteristics of instabilities as for Newtonian fluids except that the wavelength of maximum growth can be small even at vanishing velocities, i.e. instability may occur even in the absence of gravity effects, as soon as the yield stress of the displaced fluid is

## Chapter 1. Introduction

sufficiently large. Linder et al., [48], validated these results with their experimental work.

### 1.3.4 Interfacial instabilities

There have been a number of studies of stability of immiscible iso-density multi-layer Poiseuille flows, in the two classical geometries of a plane channel and pipe. Whereas we are interested in the multi-layer Hele-Shaw cell, the interfacial instability is similar in both flows, so we will review of some of the most relevant work done for this subject. Both analytical and numerical studies of these instabilities have taken place. Explanations of physical mechanisms that govern this type of instability have been given in [38].

The earliest theoretical investigations of instability due to viscosity stratification is probably the classical study of Yih [88]. He studied the stability of two superposed fluids of different viscosities in plane Couette and Poiseuille flows, using an asymptotic method based on long wavelengths. He found that for equal densities of the fluids the flow can be unstable at any Reynolds number, however small. Hooper and Boyd, [41], produced the first studies of such flows in the short wavelength limit, for a plane Couette flow in which each fluid occupies a halfplane. Yiantsios and Higgins [87] consider a plane Poiseuille flow, in which 2 fluids flow together down a channel. They extend Yih's results [88] for small wave numbers (long wavelength) to large wave numbers, accounting for differences in density and fluid layer thickness ratios, as well as the effects of interfacial tension and gravity. They show that the flow is unstable for large wave-numbers and is linearly unstable to a shear mode instability. In 2003 Ern et.al., [23], reported that instability exists in a shear flow of two superposed miscible fluids with the presence of a continuous but steep variation of viscosity. They showed that Yih's results are valid even in the case of continuous changes in viscosity. Instabilities arise provided that the thickness of the interface is small and the Péclet number to be sufficiently large, i.e. close to the immisicible limit. Govindarajan, [35] found similar results using a different profile for the viscosity variation.

For the case of multi-layer flows in porous media, Raghavan \& Marsden, [62] showed, in the presence of density difference, if the heavier fluid was in top of the lighter one, the flow was
unstable for all viscosity ratios, if the case is reversed, the lighter on top of the heavier, they claim that the flow will be unstable only for large viscosity ratios (the more viscous on top of the less viscous). The isodensity case resulted stable for all wavelengths. Similar results were found by Gondret et al. in [33]. Although their problem is not the same as the one we consider in chapter 4 , their results are fairly similar with the ones that we found in §4.5.1. Zeybek \& Yortsos, [91], considered the parallel flow of two immiscible fluids in a Hele-Shaw cell with equal densities. Through a weakly nonlinear analysis they found that the frequencies of the disturbances of the flow are strictly real and when the mobility ratio is different from one this theory predicts the presence of dispersive waves. They showed analytically and later experimentally the presence of solitary waves at the interface of the fluids. After the waves or solitons leave the channel the interface remained flat. These results reaffirm that without buoyant forces, the interface will be linearly stable. Gondret et al., [32] first considered the parallel flow of a gas on top of a viscous fluid in a horizontal Hele-Shaw cell. They presented a Kelvin-Helmholtz-Darcy theory which consists of Euler equations for inviscid flow coupled to Darcy's law for the viscous flow. They found that linear stability is governed by inertia only. Later, Hinch \& Plouraboué, [60, 39], compared Gondret's analysis by considering the full Navier-Stokes equations and got similar results.

### 1.4 Overview

As the brief account above indicates, unsteady displacements along an eccentric annular geometry often occur. Under this situation a long finger develops due to the fact that the displacement front will advance much faster on the wide side of the annulus than on the narrow side. When this happens, the interface between the fluids is likely to become unstable. Despite the large number of theoretical, numerical and experimental studies, as shown in the previous section, the stability of multi-layer parallel flows in a vertical or near vertical eccentric Hele-Shaw cell, to the knowledge of the author, has not been studied. The work presented in this thesis focuses on extending to a transient version the Hele-Shaw model for displacement flows presented in [10] and to study the the stability problem just mentioned above. In addition, we investigate
how a mud channel might be removed by pulsation of the flow rate.

The thesis is in 4 main parts, which we outline below.

1. Extension of an existing Hele-Shaw type model to the transient regime, i.e. in which the velocity field satisfies a time dependent evolution problem. Fluid accelerations are considered, but not the nonlinear inertial terms. This is presented in chapter 2.
2. In chapter 3 we address theoretical questions for our model, i.e. existence and uniqueness of a solution, and various results on the continuity of the solution with respect to changes in the model parameters. Thus, we show that our model is well-posed. We also prove that the transient model decays to the solution of a steady state problem, provided that the concentration doesn't change.
3. Stability of multi-layer annular flows is addressed in chapter 4. Unsteady displacement fronts tend to elongate along the well on the wide side. Eventually, they become effectively multi-layer flows. We investigate if such flows are in fact stable. A wide range of analytical and numerical results on this problem are presented in chapter 4.
4. One particular multi layer flow is when the mud is not removed by the displacement and remains stuck in channel along the narrow side of the annulus. We investigate removal of this layer by techniques such as pulsation of the flow. This work involves mostly numerical simulation. The results are to be found in chapter 5 .

## Chapter 2

## Model derivation

In this chapter we give a physical derivation of transient model and justification for its usage. A brief outline of the chapter is as follows.

In $\S 2.1$ we outline the pseudo-steady model of [10] for the bulk displacement flow and explain a number of phenomena for which that model is inadequate. This motivates derivation of the fully transient displacement model.

In $\S 2.2$ we derive a reduced shear flow model from the Navier-Stokes equations and discuss a. number of the modelling assumptions.

In $\S 2.3$ we approximate the reduced shear flow model from $\S 2.2$ by a gap-averaged model. We show that this approximation is valid when the concentration field changes slowly.

In $\S 2.4$ we finally derive the transient Hele-Shaw model from the gap-averaged reduced shear flow model in $\S 2.3$. This is the basic model that we study for the remainder of the thesis.

In $\S 2.5$ we introduce a variant of the model in $\S 2.4$, in which the fluids are separated by an interface, as opposed to being represented by a concentration field. This variant is more convenient for studying interfacial stability. Since the concentration field is only advected, (as is the interface), the two models are very similar. Their equivalence is discussed.

### 2.1 The pseudo-steady model and some open questions

In primary cementing the annular gap is comparatively narrow, compared to the other geometric length-scales. This is a large-scale, imprecise industrial process in which resolution on the scale of the annular gap is not always of interest. Therefore, the modelling approach that has been taken in the past is to average across the annular gap, leading to a two-dimensional Hele-Shaw type model. This basic approach to modelling the displacement is described in detail in [10] and analysed further in $[57,58,59]$.

The pseudo-steady model of [10] relates to an entire annular wellbore, along which the geometry may vary slowly. Here we will consider instead a single section of the well on which a constant geometry may be assumed. This is also the simplification considered in [57, 58, 59]. Assume therefore that fluid 1 (cement) displaces fluid 2 (drilling mud), along a uniform inclined annulus that is initially full of fluid 2 . This situation has been modelled by the following elliptic problem:

$$
\begin{align*}
\nabla \cdot\left[S_{s}+f\right] & =0  \tag{2.1}\\
\boldsymbol{S}_{s}=\left[\frac{\chi\left(\left|\nabla \Psi_{s}\right|\right)}{\left|\nabla \Psi_{s}\right|}+\frac{\tau_{Y}}{H\left|\nabla \Psi_{s}\right|}\right] \nabla \Psi_{s} & \Longleftrightarrow\left|S_{s}\right|>\frac{\tau_{Y}}{H}  \tag{2.2}\\
\left|\nabla \Psi_{s}\right|=0 & \Longleftrightarrow\left|S_{s}\right| \leq \frac{\tau_{Y}}{H} \tag{2.3}
\end{align*}
$$

where $\Psi_{s}$ is the stream function. The unwrapped narrow annular space is $(\phi, \xi) \in(0,1) \times(0, Z)$, the annular gap half-width is $H(\phi)=1+e \cos \pi \phi ; e \in[0,1)$ is the eccentricity, defined as $\Delta r /\left(r_{o}-r_{i}\right)$, see Figure 2.1. The vectorfield $f$ and the rheological properties of the fluids may depend on time and space via the mean concentration of fluid $1, \bar{c}(\phi, \xi, t)$. The function $\chi$ is a positive increasing function of $\left|\nabla \Psi_{s}\right|$, arising from the viscous shear-thinining behaviour of the fluids, and $\tau_{Y}$ is the fluid yield stress. The concentration is simply advected:

$$
\begin{align*}
\frac{\partial}{\partial t}[H \bar{c}]+\frac{\partial}{\partial \phi}\left[H \bar{v}_{s} \bar{c}\right]+\frac{\partial}{\partial z}\left[H \bar{w}_{s} \bar{c}\right] & =0  \tag{2.4}\\
\frac{\partial \Psi_{s}}{\partial \phi}=H \bar{w}_{s}, \quad \frac{\partial \Psi_{s}}{\partial z} & =-H \bar{v}_{s} \tag{2.5}
\end{align*}
$$

## Chapter 2. Model derivation



Figure 2.1: Geometry of the well

The derivation of (2.1), at fixed $t$, follows straightforwardly from the reduced shear flow model:

$$
\begin{align*}
& 0=-\frac{\partial p}{\partial \phi}+\frac{\partial}{\partial y} \tau_{s, \phi y}+g_{\phi}  \tag{2.6}\\
& 0=-\frac{\partial p}{\partial \xi}+\frac{\partial}{\partial y} \tau_{s, z y}+g_{z} \tag{2.7}
\end{align*}
$$

where $\boldsymbol{\tau}_{s}=\left(\tau_{s, \phi y}, \tau_{s, \xi y}\right)$ depends on the velocity gradients $\left(v_{s, y}, w_{s, y}\right)$ at leading order, and $\left(g_{\phi}, g_{\xi}\right)$ represents the gravitational acceleration. These equations, plus the incompressibility condition, are essentially averaged with respect to $y$, that is, the radial coordinate scaled relative to the distance from the centreline of the annulus, and cross-differentiated, to give (2.1).

The pseudo-steady model above has been studied in some detail. In a uniform inclined eccentric annulus, it appears that the displacement is characterised by one of 3 different behaviours.
a) If the interface between fluids advances steadily along the well at the same speed all around the annulus, this leads to a good displacement. These steady states have been observed in the simulations of [10]. For small eccentricities, analytical expressions can be found that describe their shape, see [57]. Their stability has been characterised in [58].
b) If a steady state displacement is not found, then invariably the interface remains unsteady and elongates into a progressively long finger.

## Chapter 2. Model derivation

c) An extreme version of the unsteady displacement occurs when the fluids on the narrow side do not move, which represents a static mud channel. This can also be observed in the simulations of [10].

The lubrication model in [59] can be used to predict which of the above states occurs for any given set of process parameters and its fair to say that the mechanisms by which the finger and mud channel form are quite well understood.

### 2.1.1 Open problems and limitations

The pseudo-steady model has a number of physical limitations and there are also aspects of the primary cementing process that are not accounted for in this model.

1. Since the velocity field is governed by a steady Hele-Shaw model for which the inertial and acceleration terms in the Navier-Stokes equations have been neglected, it seems unlikely that any type of interfacial instability can manifest, apart from the type of "viscous fingering" that results from gradients in mobility ratio. The unsteady displacement finger that advances on the wide side of the annulus can be interpreted as this type of instability. As the finger grows progressively longer, we find essentially a multi layer flow along the length of the annulus. Since these multi-layer flows are stationary solutions of the full pseudo-steady model, it is unlikely that there can be any fingering type of frontal instability of these interfaces.

In reality however, it is well known that many viscous-viscous interfaces in shear flows suffer from interfacial instabilities, see e.g. [42]. In the context of annular displacements, Tehrani et al. report observing interfacial instabilities in their experiments as the interface becomes progressively parallel to the annulus axis, [78].

We study the onset of interfacial instabilities in chapter 4 . When the interface is unstable this is likely to lead to azimuthal mixing of the fluids. In the context of primary cementing, this effect could conceivably be either good or bad for the process. If the displacing fluid considered is a cement slurry, then the risk inherent mixing is that the slurry becomes contaminated and does not later harden when the displacement stops. On the other hand,
if the displacing fluid is a spacer fluid, then azimuthal mixing may lead to a better actual displacement of the mud than if the two fluids had remained separated by a distinct interface.
2. Although static mud channel formation is well understood, methods for removal of the mud channel are not well understood. One result of our study in chapter 4 will be to show that flows with an established static mud channel are linearly stable. Thus, interfacial instabilities do not appear to be a viable method for removal. We study this further in chapter $\S 5$.
3. Pulsation techniques have been advocated at different times for primary cementing, e.g. [19], with different claimed benefits. The effect of pulsation on displacement flows is not at all understood. To pulse the flow rate is perfectly possible with the pseudo-steady model (2.1), via the boundary conditions. However, we shall see in chapter 5 that there is little effect with the pseudo-steady model. The effects are quite different when a fully transient model is considered.
4. The effects of casing rotation and reciprocation on displacement are unknown. These are common practices in horizontal well cementing. We do not however study this problem, as it requires significant changes to the type of model that we present.

Problems (1)-(3) require a transient model for the velocity field, since different timescales occur in the displacement. If the displacement flows are still relatively slow, so that Hele-Shaw type averaging retains validity, then retaining the acceleration terms in the Navier-Stokes equations leads to the simplest transient models.

### 2.2 Modelling transient bulk flow cementing displacements

As we can see in Figure 2.1, the geometry of the well can be modelled by an eccentric annulus, thus a cylindrical coordinate system $(\hat{r}, \theta, \hat{\xi})$ is used to describe the well. From now on we will focus on only a uniform annulus. This will suffice to model a section of the well, for example, a 10 m long casing stand or a laboratory experiment. Thus, $\hat{\xi}$ is the measured depth, but measured
upwards from the bottom of the annulus. The inclination angle of the annulus is denoted by $\beta$. The inner and outer radius of the eccentric annulus are denoted by, $\hat{r}_{i}$ and $\hat{r}_{o}$, respectively. $\hat{r}_{i}$ is just the radius of the inner cylinder and $\hat{r}_{o}$ is the radius of the outer cylinder, see Figure 2.1. We define the mean radius as $\hat{r}_{a}^{*}=\frac{1}{2}\left(\hat{r}_{o}+\hat{r}_{i}\right)$ and the mean half-gap width $\hat{d}=\frac{1}{2}\left(\hat{r}_{o}-\hat{r}_{i}\right)$.

To model annular cementing flows we adopt the same approach as in [10]. We will concentrate on the displacement of drilling mud by a cement slurry, thus we have only one concentration field. In outline, we start from the Navier-Stokes equations, written locally on each section of the annulus. We observe that the annular gap scale is typically much smaller than either the circumferential or axial length-scales. This leads us to differential scaling of velocities and lengths, as is characteristic of Hele-Shaw flow modelling. The Navier-Stokes equations and concentration equation for the model are:

$$
\begin{align*}
\hat{\rho}\left(\frac{\partial \hat{u}}{\partial \hat{t}}+\hat{\boldsymbol{u}} \cdot \nabla \hat{u}\right) & =\frac{1}{\hat{r}} \frac{\partial}{\partial \hat{r}}\left(\hat{r} \hat{\tau}_{\hat{r} \hat{r}}\right)+\frac{1}{\hat{r}} \frac{\partial}{\partial \theta} \hat{\tau}_{\hat{r} \theta}+\frac{\partial}{\partial \hat{\xi}} \hat{\tau}_{\hat{r} \hat{\xi}}-\frac{\hat{\tau}_{\theta \theta}}{\hat{r}}-\frac{\partial \hat{p}}{\partial \hat{r}}+\hat{\rho} \hat{g}_{\hat{r}}  \tag{2.8}\\
\hat{\rho}\left(\frac{\partial \hat{v}}{\partial \hat{t}}+\hat{\boldsymbol{u}} \cdot \nabla \hat{v}\right) & =\frac{1}{\hat{r}^{2}} \frac{\partial}{\partial \hat{r}}\left(\hat{r}^{2} \hat{\tau}_{\theta \hat{r}}\right)+\frac{1}{\hat{r}} \frac{\partial}{\partial \theta} \hat{\tau}_{\theta \theta}+\frac{\partial}{\partial \hat{\xi}} \hat{\tau}_{\theta \hat{\xi}}-\frac{1}{\hat{r}} \frac{\partial \hat{p}}{\partial \theta}+\hat{\rho} \hat{g}_{\theta}  \tag{2.9}\\
\hat{\rho}\left(\frac{\partial \hat{w}}{\partial \hat{t}}+\hat{\boldsymbol{u}} \cdot \nabla \hat{w}\right) & =\frac{1}{\hat{r}} \frac{\partial}{\partial \hat{r}}\left(\hat{r} \hat{\tau}_{\hat{\xi} \hat{r}}\right)+\frac{1}{\hat{r}} \frac{\partial}{\partial \theta} \hat{\tau}_{\hat{\xi} \theta}+\frac{\partial}{\partial \hat{\xi}} \hat{\tau}_{\hat{\xi} \hat{\xi}}-\frac{\partial \hat{p}}{\partial \hat{\xi}}+\hat{\rho} \hat{g}_{\hat{\xi}}  \tag{2.10}\\
0 & =\frac{1}{\hat{r}} \frac{\partial}{\partial \hat{r}}(\hat{r} \hat{u})+\frac{1}{\hat{r}} \frac{\partial \hat{v}}{\partial \theta}+\frac{\partial \hat{w}}{\partial \hat{\xi}} \tag{2.11}
\end{align*}
$$

where the gravitational vector is given by

$$
\begin{gather*}
\hat{g}_{\hat{r}}=\hat{g} \sin \beta(\hat{\xi}) \cos \theta, \quad \hat{g}_{\theta}=\hat{g} \sin \beta(\hat{\xi}) \sin \theta, \quad \hat{g}_{\hat{\xi}}=\hat{g} \cos \beta(\hat{\xi}) . \\
\frac{\partial c}{\partial \hat{t}}+\frac{1}{\hat{r}} \frac{\partial}{\partial \hat{r}}(\hat{r} \hat{u} c)+\frac{1}{\hat{r}} \frac{\partial}{\partial \theta}(\hat{v} c)+\frac{\partial}{\partial \hat{\xi}}(\hat{w} c)=\nabla \cdot(\hat{D}(c, \hat{\boldsymbol{u}}) \nabla c) \tag{2.12}
\end{gather*}
$$

where $c \in[0,1]$.

### 2.2.1 Non-Dimensionalisation

The main aim of this section is to reduce the system (2.8)-(2.10) and (2.12) to the transient version of (2.6)-(2.7). We will average accross the gap thickness using almost the same scaling arguments as in [10]. This is motivated by the fact that in a typical oil well we have the following situation,
annular gap $\ll$ annular circumference $\ll$ annular length-scale.

Our goal is to derive a two-dimensional evolution model of the fluid motion, in azimuthal and axial directions. In contrast to $[10]$ we will scale time with a viscous time scale and thus retain time derivatives terms at leading order but not the inertial terms. In the following sections we present geometry and velocity scales and the reduced systems of equations.

## Geometry and velocity scales

Here we consider a single section of the well, of uniform geometry and length $Z$. We scale axial and azimuthal coordinates as:

$$
z=\frac{\hat{\xi}-\hat{\xi}_{b h}}{\pi \hat{r}_{a}^{*}}, \quad \phi=\frac{\theta}{\pi}
$$

A measure of the narrowness of the annulus, $\delta^{*}$, is given by:

$$
\delta^{*}=\frac{\hat{d}}{\hat{r}_{a}^{*}}
$$

The dimensionless eccentricity is given by:

$$
e=\frac{\hat{e}}{2 \hat{d}}
$$

## Chapter 2. Model derivation

Therefore the radial coordinate is scaled relative to the distance from the centreline of the annulus,

$$
y=\frac{\hat{r}-\hat{r}_{a}^{*}}{\hat{r}_{a}^{*} \delta^{*}} .
$$

The axial and azimuthal velocities are scaled with:

$$
U^{*}=\frac{\hat{Q}^{*}}{A^{*}},
$$

where $\hat{Q}^{*}$ is the maximum flow rate and $A^{*}$ is the cross-sectional area of the annulus. The radial velocity is scaled with $\delta^{*} U^{*}$.

We assume that the fluids can be described by the Herschel-Bulkley model, characterized by density, yield stress, consistency and power-law index. A characteristic scale for the rate of strain is:

$$
\hat{\dot{\gamma}}^{*}=\frac{\hat{U}^{*}}{\hat{r}_{a}^{*} \delta^{*} \pi}
$$

using this we can define scales for the shear stress, viscosity and pressure:

$$
\hat{\tau}^{*}=\max _{k}\left[\hat{\tau}_{k, Y}+\hat{\kappa}_{k}\left(\hat{\dot{\gamma}}^{*}\right)^{n_{k}}\right], \quad \hat{\mu}^{*}=\frac{\hat{\tau}^{*}}{\hat{\dot{\gamma}}^{*}}, \quad \hat{p}^{*}=\frac{\hat{\tau}^{*} \pi}{\delta^{*}} .
$$

The dimensionless rheological parameters are defined by:

$$
\hat{\tau}_{k, Y}^{*}=\hat{\tau}^{*} \tau_{k, Y}, \quad \hat{\kappa}_{k}\left(\hat{\gamma}^{*}\right)^{n_{k}}=\hat{\mu}^{*} \kappa_{k}, \quad n_{k}=n_{k},
$$

and the fluid densities are scaled with

Chapter 2. Model derivation

$$
\hat{\rho}^{*}=\max _{k}\left[\hat{\rho}_{k}\right] .
$$

Up to here the scales used to reduce our system of equations are the same as in [10], the only difference between our scaling and theirs is the timescale we use:

$$
t_{v}^{*}=\frac{\hat{\rho}^{*}\left(\delta^{*} \pi \hat{r}_{a}^{*}\right)^{2}}{\mu^{*}}
$$

clearly this is a viscous timescale.

### 2.2.2 Reduced shear flow equations

After the non-dimensionalisation we have the following reduced system of equations, at leading order in $\delta^{*}$ :

$$
\begin{align*}
0 & =-\frac{\partial p}{\partial y}  \tag{2.13}\\
\rho \frac{\partial v}{\partial t} & =-\frac{1}{r_{a}} \frac{\partial p}{\partial \phi}+\frac{\partial}{\partial y} \tau_{\phi y}+\frac{g_{\phi}}{S t^{*}}  \tag{2.14}\\
\rho \frac{\partial w}{\partial t} & =-\frac{\partial p}{\partial z}+\frac{\partial}{\partial y} \tau_{\xi y}+\frac{g_{\xi}}{S t^{*}}  \tag{2.15}\\
0 & =\frac{\partial u}{\partial y}+\frac{1}{r_{a}} \frac{\partial v}{\partial \phi}+\frac{\partial w}{\partial \xi} \tag{2.16}
\end{align*}
$$

where

$$
S t^{*}=\frac{\hat{\mu}^{*} \hat{U}^{*}}{\hat{\rho}^{*} \hat{g}^{*}\left[\hat{r}_{a}^{*} \delta^{*}\right]^{2}}
$$

is the Stokes number which is usually $\ll 1$.

The boundary conditions at $y=0$ and $y=H$ are:

$$
\begin{array}{cc}
u=v=w=0, & \text { at } y=H \\
u=\tau_{\phi y}=\tau_{\xi y}=0, & \text { at } y=0, \tag{2.18}
\end{array}
$$

## Chapter 2. Model derivation

The leading order rate of strain is characterized by:

$$
\begin{equation*}
\dot{\gamma} \sim\left[\left(\frac{\partial v}{\partial y}\right)^{2}+\left(\frac{\partial w}{\partial y}\right)^{2}\right]^{1 / 2}+O\left(\delta^{*} / \pi\right) \tag{2.19}
\end{equation*}
$$

In regions where the fluids are yielded we have:

$$
\begin{equation*}
\tau_{\phi y} \sim \eta(\dot{\gamma})\left(\frac{\partial v}{\partial y}\right)+O\left(\delta^{*} / \pi\right) \quad \tau_{\xi y} \sim \eta(\dot{\gamma})\left(\frac{\partial w}{\partial y}\right)+O\left(\delta^{*} / \pi\right) \tag{2.20}
\end{equation*}
$$

where $\eta$ is an effective viscosity depending on the rate of strain and rheological parameters. We return to this reduced system of equations in $\S 2.3$, but first continue with the gap-averaged velocity and concentration fields.

### 2.2.3 2D averaged velocity field

The aim of our modelling exercise is to reduce to a 2D model system of equations, from the shear flow model (2.13)-(2.16). To eliminate $u$ from (2.16), we average across the gap and use the no-slip conditions at the walls of the annulus. A simplifying assumption made is that the annulus is assumed sufficiently narrow that the velocity components are approximated to $O\left(\delta^{*}\right)$ by a velocity field that can be assumed symmetric about the annular gap centreline, $y=0$, i.e. curvature effects only manifest at $O\left(\delta^{*}\right)$. This allows us to average over only half the annular gap. The result of the averaging is:

$$
\begin{equation*}
\frac{\partial}{\partial \phi}(H \bar{v})+\frac{\partial}{\partial \dot{\xi}}(H \bar{w})=0 \tag{2.21}
\end{equation*}
$$

where

$$
\bar{v}(\phi, \xi, t)=\frac{1}{H} \int_{0}^{H} v \mathrm{~d} y \quad \bar{w}(\phi, \xi, t)=\frac{1}{H} \int_{0}^{H} w \mathrm{~d} y .
$$

Equation (2.21) is satisfied using a stream function formulation:

$$
\begin{equation*}
H \bar{w}=\frac{\partial \Psi}{\partial \phi}, \quad H \bar{v}=-\frac{\partial \Psi}{\partial \xi} \tag{2.22}
\end{equation*}
$$

### 2.2.4 Fluid concentration field

The concentration of the displaced fluid is governed by (2.12). As our aim is to derive a 2D model for the large-scale motions of the fluids, we will want to work with quantities that are averaged across the annular gap. Thus, we proceed as in [10], by assuming that the concentration is uniform across the annular gap. This allows us to interchange the average of $c u$ with the product of the the averages of $c$ and $\mathbf{u}$, in the convective terms. Such approximations are commonly made in hydraulic models, e.g. for the inertial terms, often including a prefactor to represent an assumed degree of correlation between the two averaged quantities. Here we have no a priori knowledge of whether the concentration and velocity fields are correlated and have taken the prefactor to be 1 .

With this assumption, we average across the gap, use the non-slip conditions at the walls and assume no flux of any fluid through the walls. The concentration satisfies:

$$
\begin{equation*}
\frac{1}{\epsilon} \frac{\partial}{\partial t}[H \bar{c}]+\frac{\partial}{\partial \phi}[H \bar{v} \bar{c}]+\frac{\partial}{\partial \xi}[H \bar{w} \bar{c}]=\frac{1}{P e}\left(\frac{\partial}{\partial \phi}\left[H \hat{D} \frac{\partial \bar{c}}{\partial \phi}\right]+\frac{\partial}{\partial \xi}\left[H \hat{D} \frac{\partial \bar{c}}{\partial \xi}\right]\right) \tag{2.23}
\end{equation*}
$$

where the diffusivity has been scaled with a global diffusivity scale, $\hat{D}^{*}$, and the Péclet number is defined as:

$$
P e=\frac{\pi \hat{r}_{a}^{*} \hat{w}^{*}}{\hat{D}^{*}}
$$

Typically, $\hat{D}^{*} \sim 10^{-10} m^{2} s^{-1}$, thus $P e \gg 1$ and typically $P e \sim 10^{10}-10^{12}$. Denoting the scaled annular length by $Z$, we may expect a diffusive layer of intermediate concentrations, in the azimuthal and axial directions, of thickness $\sim \sqrt{Z / P e} \ll 1$. Therefore we neglect molecular diffusion in our leading order model. Thus, the concentration field will satisfy:

$$
\begin{equation*}
\frac{1}{\epsilon} \frac{\partial}{\partial t}[H \bar{c}]+\frac{\partial}{\partial \phi}[H \bar{v} \bar{c}]+\frac{\partial}{\partial \xi}[H \bar{w} \bar{c}]=0 \tag{2.24}
\end{equation*}
$$

We shall assume that during the displacement, at the bottom of the well the concentration will
always be zero and at the top one, i.e. cement at the bottom and mud at the top. As well, natural boundary conditions at $\phi=0,1$ will be assumed.

Note the only difference between (2.24) and (2.4) results from the different timescale used. We shall denote the ratio $t_{v}^{*} / t_{a v}^{*}$ where $t_{a v}^{*}=\frac{\pi \hat{r}_{a}^{*}}{\hat{U}^{*}}$ by $\epsilon$, i.e.

$$
\epsilon=\frac{\hat{\rho}^{*} \delta^{* 2} \pi \hat{r}_{a}^{*} \hat{U}^{*}}{\hat{\mu}^{*}}
$$

Formally $\epsilon \rightarrow 0$ as $\delta^{*} \rightarrow 0$. However, although $t_{v}^{*}=O\left(\delta^{* 2}\right)$, the ratio $t_{v}^{*} / t_{a v}^{*}$ is often $O(1)$. This is because $\delta^{*}$ is small but not zero and $\frac{\hat{\mu}^{*}}{\hat{\rho}^{*} r_{a}^{*}}$ is relatively large. In this thesis we shall consider both $\epsilon \rightarrow 0$, for which the concentration is fixed in time as the velocity evolves on the viscous timescale, and also $\epsilon \sim 1$. Since much of the complexity of the model is in the evolution of the velocity field, it can be helpful to isolate this evolution. Theoretical properties of the velocity problem (in the limit $\epsilon \rightarrow 0$ ) are studied in chapter 3. Later, in chapters 4 and 5 we consider the distinguished limit of $\epsilon>0$ and $\delta^{*} \rightarrow 0$. For this model, both velocity and concentration fields are fully transient and coupled. This allows us to study the transient phenomena cited before in $\S 2.1$.

Apart from viscous and advective timescales, there may be timescales associated with the growth of any temporal instability or with pulsation of the flow. These may excite a response in either the velocity field or the concentration field. This is another motivation for retaining the time derivatives in both evolution equations

## Uniform concentration assumption

The assumption of a uniform concentration across the annular gap is non-trivial. One way this might arise is via Taylor dispersion. In $[85,92]$ the authors showed that in the limit $\varepsilon P e_{G} \rightarrow 0$, where $\varepsilon$ is the inverse aspect ratio of the annulus and $P e_{G}$ is the Péclet number, (based on the gap width), diffusion effects dominate and a uniform or near uniform concentration field across the gap is achieved. For our problem, the Peclet numbers based on the gap width are in the range $P e_{G} \sim 10^{8}-10^{10}$ and the aspect ratio in a typical well can not justify $\varepsilon P e_{G} \ll 1$.

Therefore, we are certainly not in the Taylor dispersion regime. Indeed, from the perspective of having large $P e_{G}$, we might expect to be in the almost immiscible regime. Here either the displacements are effective and the uniform concentration assumption is reasonable, or otherwise long fingers are likely to form, with thin miscible diffuse interfacial layers of thickness $P e_{G}^{-1 / 2}$. Although such fingers are observed in the lab setting, we believe this scenario to be unlikely in the well for a number of practical reasons.

First, along the well one generally has centralisers fitted at regular intervals. These will disrupt and mix the flow across the gap. Second, even without centralisers, the existence of a long finger extending stably over 10 's or 100 's of meters is implausible. There are various instabilities that may occur. For example, fingering \& frontal instabilities were studied in [46, 47] in the context of the downward miscible displacement of one fluid by a lighter and less viscous fluid, between parallel plates. If frontal instabilities don't occur and a finger develops, there have been many studies of interfacial instabilities in shear flows, see e.g. [5, 42, 67] for an overview. In general, such flows are not robustly stable to all wavelengths and rheology combinations. The situation is not apparently helped by having a thin diffuse interface layer over which fluid properties change. Recently in, [23] the authors studied miscible multilayer flows where the viscosity varies rapidly between two constant values. As the layer thickness decreased, they showed that interfacial instabilities arise in the same way as for the immiscible fluid limit (with no surface tension). Thirdly, for the cases that may be stable in an idealised mathematical or lab setting, we must recall that the annulus outer wall is perhaps rough with small scale unevenness, that the pump pulsations will naturally occur, and that there are also likely to be azimuthal secondary flows.

It is almost impossible to judge the effects of all the above effects, and our assumption of an homogeneous concentration is really just a statement that our model is not designed to study small scale flow features on the gap scale.

## Chapter 2. Model derivation

### 2.3 Reduced shear flow models

At this point we would like to proceed in the same way as in [10] and find a transient displacement Hele-Shaw type of model. Unfortunately, the model (2.13)-(2.16) does not lead to a closure relation between the modified pressure gradient and the gap-averaged velocity. Thus, it is only a simpler 3D model of the flow. To circumvent this difficulty we develop an approximation to (2.14) \& (2.15) in which the solution remains close to the solution of the transient model, i.e. the difference between the solutions of both models is $O(\epsilon)$; and allows a closure relation between pressure gradients and the velocity.

Our starting point for the model derivation is the unsteady leading order shear-flow model, (2.14) \& (2.15), which we write as:

$$
\begin{equation*}
\rho \boldsymbol{u}_{t}=\boldsymbol{G}+\frac{\partial}{\partial y} \boldsymbol{\tau} \tag{2.25}
\end{equation*}
$$

where $\boldsymbol{u}=(v, w)$ and

$$
\begin{equation*}
G=\left(G_{\phi}, G_{\xi}\right) \equiv\left(-\frac{\partial p}{\partial \phi}+g_{\phi},-\frac{\partial p}{\partial \xi}+g_{\xi}\right) \tag{2.26}
\end{equation*}
$$

The system (2.25) is supplemented with the conditions (2.13) \& (2.16) and

$$
\begin{equation*}
\left(g_{\phi}, g_{\xi}\right)=\left(\frac{\rho \sin \beta \sin \pi \phi}{S t^{*}},-\frac{\rho \cos \beta}{S t^{*}}\right) \tag{2.27}
\end{equation*}
$$

As well as the system (2.25), which we shall refer to as the transient system, we shall consider the steady system, $(2.6) \&(2.7)$, which we write as:

$$
\begin{equation*}
0=\boldsymbol{G}_{s}+\frac{\partial}{\partial y} \boldsymbol{\tau}_{s} \tag{2.28}
\end{equation*}
$$

The subscript $s$ is used to indicate that we are seeking a velocity field $\boldsymbol{u}_{s}=\left(v_{s}, w_{s}\right)$ that satisfies (2.28), where potentially $\boldsymbol{G}_{\boldsymbol{s}} \neq \boldsymbol{G}$, (e.g. because the fluid concentration will be advected differently by $\boldsymbol{u}_{s}$ than by $\left.\boldsymbol{u}\right)$. As with (2.25), the pressure and concentration are independent of $y$. In addition to (2.28) the incompressibility condition (2.16) is satisfied:

$$
0=\frac{\partial u_{s}}{\partial y}+\frac{\partial v_{s}}{\partial \phi}+\frac{\partial w_{s}}{\partial \xi}
$$

## Chapter 2. Model derivation

Finally, we consider the gap-averaged approximation to (2.25):

$$
\begin{equation*}
\rho \bar{u}_{a, t}=\boldsymbol{G}_{a}+\frac{\partial}{\partial y} \boldsymbol{\tau}_{a} \tag{2.29}
\end{equation*}
$$

where again the subscript $a$ denotes that we seek a velocity field $\boldsymbol{u}_{a}=\left(v_{a}, w_{a}\right)$, and that potentially $\boldsymbol{G}_{a} \neq \boldsymbol{G}$; (2.29) is supplemented with the incompressibility condition:

$$
\begin{equation*}
0=\frac{\partial u_{a}}{\partial y}+\frac{\partial v_{a}}{\partial \phi}+\frac{\partial w_{a}}{\partial \xi} . \tag{2.30}
\end{equation*}
$$

The overbar indicates a gap-averaged quantity, e.g. $\overline{\boldsymbol{u}}_{a}=\left(\bar{v}_{a}, \bar{w}_{a}\right)$ :

$$
\bar{v}_{a}=\frac{1}{H} \int_{0}^{H} v_{a} \mathrm{~d} y, \quad \bar{w}_{a}=\frac{1}{H} \int_{0}^{H} w_{a} \mathrm{~d} y .
$$

It is important to note that the time evolution in (2.25), (2.28) and (2.29) is distinctly different. Starting with the steady problem (2.28), note that given $\boldsymbol{G}_{\boldsymbol{s}}$ at a point $(\phi, \xi)$, we are able to find a unique solution $\left(v_{s}, w_{s}\right)$, (and could then integrate the incompressibility condition to find $u_{a}$ ). All the time dependency in $\left(v_{s}, w_{s}\right)$ must come from the vectorfield $\boldsymbol{G}_{s}$, which is determined from (2.1), with appropriate boundary conditions, and the advection of the concentration field. When $G_{s}$ and $\bar{c}_{s}$ vary only slowly with time, $\left(v_{s}, w_{s}\right)$ are (pseudo-)steady on the $O(1)$ timescale. Considering now (2.25), we see that this is simply the parabolic extension of the elliptic system (2.28). If $\boldsymbol{G}$ and $\bar{c}$ are time invariant, $(v, w)$ will still evolve with time, i.e. (2.25) is an evolution problem for $(v, w)$, regardless of $\boldsymbol{G}$ and $\bar{c}$. Turning now to (2.29), we could interpret this as an integro-differential equation for $\left(v_{a}, w_{a}\right)$, but it is slightly unclear if, for specified $\boldsymbol{G}_{a}$ and $\bar{c}_{a}$, the evolution of $\left(v_{a}, w_{a}\right)$ is properly determined. Rather, given time derivatives of the averaged velocities ( $\bar{v}_{a, t}, \bar{w}_{a, t}$ ), we can determine unique ( $v_{a}, w_{a}$ ), (i.e. this is the same problem as (2.28) at fixed $t$ ). It remains to determine whether $(\bar{v}, \bar{w})$ evolve in a manner consistent with these solutions.

Apart from retention of the time derivative terms, the model and assumptions leading to (2.25) are identical with those in [10]. Inertial terms have been neglected under the asymptotic assumption, $\epsilon \ll 1$. We shall be interested in temporal changes in the annulus that occur over the slow timescale $\epsilon^{-1}$. It will however be the gap-averaged system (2.29) that forms the basis
for the generalisation of (2.1), that we consider later in chapter 3. Since (2.29) is not derived directly, we should like to estimate how close the solutions of (2.25) and (2.29) are. We consider first the case $n=1$, in which case the fluids are Bingham fluids. Later we discuss the case $n<1$.

### 2.3.1 Stability of $u$ and $u_{s}$, at $(\phi, \xi)$ :

For $n=1$, problems (2.25) and (2.28) are of a form considered by [20, 30]. Multiplying (2.28) by $\boldsymbol{u}_{s}$ and integrating with respect to $y$, we have:

$$
\begin{equation*}
a_{s}\left(\boldsymbol{u}_{s}, \boldsymbol{u}_{s}\right)+j_{s}\left(\boldsymbol{u}_{s}\right)=L\left(\boldsymbol{G}_{s} \cdot \boldsymbol{u}_{s}\right) \tag{2.31}
\end{equation*}
$$

where

$$
\begin{align*}
a_{s}(\boldsymbol{u}, \boldsymbol{v}) & =\kappa_{s} \int_{0}^{H} \boldsymbol{u}_{y} \cdot \boldsymbol{v}_{y} \mathrm{~d} y  \tag{2.32}\\
j_{s}(\boldsymbol{v}) & =\tau_{Y, s} \int_{0}^{H}\left|\boldsymbol{v}_{y}\right| \mathrm{d} y  \tag{2.33}\\
L(f) & =\int_{0}^{H} f \mathrm{~d} y \tag{2.34}
\end{align*}
$$

From this we can deduce that:

$$
\begin{equation*}
\left\|\boldsymbol{u}_{s}\right\|_{H_{s}^{1}} \leq \frac{H}{\kappa_{s} \pi}\left\|G_{s}\right\|_{L^{2}} \tag{2.35}
\end{equation*}
$$

from the Cauchy-Schwarz and Poincaré inequalities, where $\left\|\boldsymbol{u}_{s}\right\|_{H_{s}^{1}}$ is the semi-norm:

$$
\|\boldsymbol{v}\|_{H_{s}^{1}}=\left(\int_{0}^{H}\left|\boldsymbol{v}_{y}\right|^{2} \mathrm{~d} y\right)^{1 / 2}
$$

which is equivalent to the $H^{1}$-norm. Using the methods in [20, 30] we may assert that there exists a unique solution $\boldsymbol{u}_{s}$. Time-dependency enters $\boldsymbol{u}_{s}$ only through $\boldsymbol{G}_{s}$ and $\bar{c}_{s}$. We assume that $\boldsymbol{u}_{s}$ is differentiable with respect to each of the physical fluid properties, which depend on $\bar{c}_{s}$, and with respect to the components of $\boldsymbol{G}_{\boldsymbol{s}}$. Therefore, we may suppose that

$$
\begin{equation*}
\boldsymbol{u}_{s, t}=\epsilon \boldsymbol{b}(\epsilon t), \tag{2.36}
\end{equation*}
$$

for some vector $\boldsymbol{b}$.

Now we consider stability of (2.25). Multiplying (2.25) by ( $\boldsymbol{u}_{s}-\boldsymbol{u}$ ) and integrating with respect to $y$ we have:

$$
\begin{gather*}
\rho \int_{0}^{H} \boldsymbol{u}_{t} \cdot\left(\boldsymbol{u}_{s}-\boldsymbol{u}\right) \mathrm{d} y \geq \quad L\left(\boldsymbol{G} \cdot\left(\boldsymbol{u}_{s}-\boldsymbol{u}\right)\right)-a\left(\boldsymbol{u}, \boldsymbol{u}_{s}-\boldsymbol{u}\right) \\
-j\left(\boldsymbol{u}_{s}\right)+j(\boldsymbol{u}) \tag{2.37}
\end{gather*}
$$

where $a$ and $j$ are defined analogously to $a_{s}$ and $j_{s}$, but with non-subscripted rheological constants. Combining (2.36) with (2.28), we have:

$$
\begin{equation*}
\rho \boldsymbol{u}_{s, t}=\boldsymbol{G}_{s}+\boldsymbol{\tau}_{s, y}+\epsilon \rho b \tag{2.38}
\end{equation*}
$$

Multiplying by $\left(\boldsymbol{u}-\boldsymbol{u}_{s}\right)$ and integrating with respect to $y$ gives the following inequality:

$$
\begin{align*}
\rho \int_{0}^{H} \boldsymbol{u}_{s, t} \cdot\left(\boldsymbol{u}-\boldsymbol{u}_{s}\right) \mathrm{d} y \geq & L\left(\boldsymbol{G}_{s} \cdot\left(\boldsymbol{u}-\boldsymbol{u}_{s}\right)\right)-a_{s}\left(\boldsymbol{u}_{s}, \boldsymbol{u}-\boldsymbol{u}_{s}\right)  \tag{2.39}\\
& -j_{s}(\boldsymbol{u})+j_{s}\left(\boldsymbol{u}_{s}\right)+\epsilon \rho L\left(\boldsymbol{b} \cdot\left(\boldsymbol{u}-\boldsymbol{u}_{s}\right)\right)
\end{align*}
$$

Summing the two inequalities we have

$$
\begin{align*}
\frac{\rho}{2} \frac{\partial}{\partial t} \int_{0}^{H}\left|\boldsymbol{u}_{s}-\boldsymbol{u}\right|^{2} \mathrm{~d} y \leq & -a\left(\boldsymbol{u}_{s}-\boldsymbol{u}, \boldsymbol{u}_{s}-\boldsymbol{u}\right)+L\left(\left(\boldsymbol{G}-\boldsymbol{G}_{s}\right) \cdot\left(\boldsymbol{u}-\boldsymbol{u}_{s}\right)\right) \\
& +\epsilon \rho L\left(\boldsymbol{b} \cdot\left(\boldsymbol{u}_{s}-\boldsymbol{u}\right)\right)+\left(1-\frac{\kappa}{\kappa_{s}}\right) a_{s}\left(\boldsymbol{u}_{s}, \boldsymbol{u}-\boldsymbol{u}_{s}\right) \\
& +\left(1-\frac{\tau_{Y}}{\tau_{Y, s}}\right)\left[j_{s}(\boldsymbol{u})-j_{s}\left(\boldsymbol{u}_{s}\right)\right] \tag{2.40}
\end{align*}
$$

Using the Cauchy-Schwarz inequality we have:

$$
\begin{align*}
\frac{\rho}{2} \frac{\partial}{\partial t}\left\|u_{s}-\boldsymbol{u}\right\|_{L^{2}}^{2} \leq & -\kappa\left\|\boldsymbol{u}_{s}-\boldsymbol{u}\right\|_{H_{s}^{1}}^{2}  \tag{2.41}\\
& +\left(\left\|\boldsymbol{G}-\boldsymbol{G}_{s}\right\|_{L^{2}}+\epsilon \rho\|\boldsymbol{b}\|_{L^{2}}\right)\left\|\boldsymbol{u}_{s}-\boldsymbol{u}\right\|_{L^{2}} \\
& +\left(\left|\kappa_{s}-\kappa\right|\left\|\boldsymbol{u}_{s}\right\|_{H_{s}^{1}}+\left|\tau_{Y, s}-\tau_{Y}\right| H^{1 / 2}\right)\left\|\boldsymbol{u}_{s}-\boldsymbol{u}\right\|_{H_{s}^{1}}
\end{align*}
$$

We may split the first term on the right hand side in two. From the Poincaré inequality:

$$
\left\|\boldsymbol{u}_{s}-\boldsymbol{u}\right\|_{H_{s}^{1}} \geq \frac{\pi}{H}\left\|\boldsymbol{u}_{s}-\boldsymbol{u}\right\|_{L^{2}}
$$

Therefore either we have that:

$$
\begin{equation*}
\left\|u_{s}-u\right\|_{H_{s}^{1}}<\frac{2 H}{\kappa}\left(\frac{\left|\kappa_{s}-\kappa\right|}{\kappa_{s} \pi}\left\|G_{s}\right\|_{L^{2}}+\frac{\left|\tau_{Y, s}-\tau_{Y}\right|}{H^{1 / 2}}\right) \tag{2.42}
\end{equation*}
$$

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or we have the decay equation:

$$
\rho \frac{\partial}{\partial t}\left\|\boldsymbol{u}_{s}-\boldsymbol{u}\right\|_{L^{2}} \leq-\frac{\kappa \pi^{2}}{2 H^{2}}\left\|\boldsymbol{u}_{s}-\boldsymbol{u}\right\|_{L^{2}}+\left(\left\|\boldsymbol{G}-\boldsymbol{G}_{s}\right\|_{L^{2}}+\epsilon \rho\|\boldsymbol{b}\|_{L^{2}}\right)
$$

from which, for $t \in[0, T]$

$$
\begin{align*}
\left\|\boldsymbol{u}_{s}-\boldsymbol{u}\right\|_{L^{2}}(t) \leq & \int_{0}^{t}\left[\frac{\left\|\boldsymbol{G}-\boldsymbol{G}_{s}\right\|_{L^{2}}}{\rho}+\epsilon \rho\|\boldsymbol{b}\|_{L^{2}}\right] e^{-\lambda_{1}(t-s)} \mathrm{d} s \\
& +\left\|\boldsymbol{u}_{s}-\boldsymbol{u}\right\|_{L^{2}}(0) e^{-\lambda_{1} t} \tag{2.43}
\end{align*}
$$

with

$$
\lambda_{1}=\min _{t \in[0, T]}\left\{\frac{\kappa \pi^{2}}{2 H^{2} \rho}\right\} .
$$

In either case, we note that locally $\left\|\boldsymbol{u}_{s}-\boldsymbol{u}\right\|_{L^{2}}$ is bounded by the difference in the data between steady and transient problems.

### 2.3.2 Stability of $u_{a}$ and $u_{s}$, at $(\phi, \xi):$

We proceed analogously to the above, with the variational form of (2.29) \& (2.28). The latter we modify, to include the term $\overline{\boldsymbol{u}}_{s, t}$, i.e.

$$
\begin{align*}
\rho H \overline{\boldsymbol{u}}_{s, t} \cdot\left(\overline{\boldsymbol{u}}_{a}-\overline{\boldsymbol{u}}_{s}\right) \geq & L\left(\boldsymbol{G}_{s} \cdot\left(\boldsymbol{u}_{a}-\boldsymbol{u}_{s}\right)\right)-a_{s}\left(\boldsymbol{u}_{s}, \boldsymbol{u}_{a}-\boldsymbol{u}_{s}\right)  \tag{2.44}\\
& -j_{s}\left(\boldsymbol{u}_{a}\right)+j_{s}\left(\boldsymbol{u}_{s}\right)+\epsilon \rho H \overline{\boldsymbol{b}} \cdot\left(\overline{\boldsymbol{u}}_{a}-\overline{\boldsymbol{u}}_{s}\right) . \\
\rho H \overline{\boldsymbol{u}}_{a, t} \cdot\left(\overline{\boldsymbol{u}}_{s}-\overline{\boldsymbol{u}}_{a}\right) \geq & L\left(\boldsymbol{G}_{a} \cdot\left(\boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right)\right)-a_{a}\left(\boldsymbol{u}_{a}, \boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right) \\
& -j_{a}\left(\boldsymbol{u}_{s}\right)+j_{a}\left(\boldsymbol{u}_{a}\right) . \tag{2.45}
\end{align*}
$$

Summing these two inequalities, we have

$$
\begin{align*}
\frac{\rho}{2 H} \frac{\partial}{\partial t}\left\|\boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right\|_{L^{1}}^{2} \leq & -a_{a}\left(\boldsymbol{u}_{s}-\boldsymbol{u}_{a}, \boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right)+\epsilon \rho H \overline{\boldsymbol{b}} \cdot\left(\overline{\boldsymbol{u}}_{s}-\overline{\boldsymbol{u}}_{a}\right) \\
& +L\left(\left(\boldsymbol{G}_{s}-\boldsymbol{G}_{a}\right) \cdot\left(\boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right)\right) \\
& +\left(1-\frac{\kappa_{a}}{\kappa_{s}}\right) a_{s}\left(\boldsymbol{u}_{s}, \boldsymbol{u}_{a}-\boldsymbol{u}_{s}\right) \\
& +\left(1-\frac{\tau_{Y, a}}{\tau_{Y, s}}\right)\left[j_{s}\left(\boldsymbol{u}_{a}\right)-j_{s}\left(\boldsymbol{u}_{s}\right)\right] \tag{2.46}
\end{align*}
$$

Again we split the decay term in half, and bound above using the Cauchy-Schwarz and Poincaré inequalities. Either we have:

$$
\begin{align*}
\left\|\boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right\|_{H_{s}^{1}}< & \left|\kappa_{s}-\kappa_{a}\right| \frac{2 H\left\|\boldsymbol{G}_{s}\right\|_{L^{2}}}{\pi \kappa_{s} \kappa_{a}}+\left|\tau_{Y, s}-\tau_{Y, a}\right| \frac{2 H^{1 / 2}}{\kappa_{a}} \\
& +\left\|\boldsymbol{G}_{s}-\boldsymbol{G}_{a}\right\|_{L^{2}} \frac{2 H}{\pi \kappa_{a}} \tag{2.47}
\end{align*}
$$

or we have the decay equation:

$$
\frac{\rho}{H} \frac{\partial}{\partial t}\left\|\boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right\|_{L^{1}} \leq-\frac{\kappa_{a} \pi^{2}}{2 H^{3}}\left\|\boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right\|_{L^{1}}+\epsilon \rho|\overline{\boldsymbol{b}}|
$$

from which, for $t \in[0, T]$

$$
\begin{align*}
\left\|\boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right\|_{L^{1}}(t) \leq & \int_{0}^{t} \epsilon H|\overline{\boldsymbol{b}}| e^{-\lambda_{2}(t-s)} \mathrm{d} s \\
& +\left\|\boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right\|_{L^{1}}(0) e^{-\lambda_{2} t} \tag{2.48}
\end{align*}
$$

with

$$
\lambda_{2}=\min _{t \in[0, T]}\left\{\frac{\kappa_{a} \pi^{2}}{2 H^{2} \rho}\right\}
$$

Again we see that $\left\|u_{s}-u\right\|_{L^{1}}$ remains uniformly bounded. Note that we cannot expect results on $\left\|\boldsymbol{u}_{s}-\boldsymbol{u}_{a}\right\|_{L^{2}}$, since the averaged problem leads only to an equation for evolution of the mean velocity.

## Summary and further comments

Combining (2.43) \& (2.48), we see that $\left\|\boldsymbol{u}_{a}-\boldsymbol{u}\right\|_{L^{1}}(t)$ is bounded by differences in the data between the steady, transient and averaged problems. When these differences are uniformly $O(\epsilon)$ over the slow timescale, we can certainly expect that $\left\|\boldsymbol{u}_{a}-\boldsymbol{u}\right\|_{L^{1}}(t)=O(\epsilon)$. Since we work with the gap-averaged velocity fields, this is the type of approximation result that is needed, i.e. the averages are close provided $\epsilon$ is small. In words, this means that provided the flow has sufficient time for viscosity to act, before the fluid properties or other process features change locally, then it is valid to interchange the time derivative and averaging operators. This is valid in the sense that the averages of the velocity fields, taken across the gap, remain close.

More generally, we consider boundary conditions on the annulus that vary only on the slow timescale $\epsilon^{-1}$, on which timescale $\bar{c}$ also evolves. Decay of initial errors occurs on an $O(1)$
timescale, after which the errors are simply due to the differences in $G$ and the material properties. These are governed by the evolution of the gap-averaged velocity fields, which takes place on the slow timescale. We should therefore expect $\left\|\boldsymbol{u}_{a}-\boldsymbol{u}\right\|_{L^{1}}(t)=O(\epsilon)$ on a timescale $[0, T]$, where $T^{-1} \sim o(\epsilon)$, but $T^{-1} \sim \epsilon$ is unrealistic.

The exact timescale will of course depend on the particular flow studied. For example, for a stable single fluid flow, one would expect $\left\|\boldsymbol{u}_{a}-\boldsymbol{u}\right\|_{L^{1}}(t)=O(\epsilon)$ on the slow timescale. If we wish to consider instability of a steady solution, we must expect that the model breaks down as soon as $O(1)$ differences in the data evolve, i.e. depending on the growth timescale for the instability.

The arguments for $n<1$ are more difficult, especially for two-dimensional flows. To circumvent this, we may note that the velocity $\boldsymbol{u}$, governed by (2.25), decays to a vector in the direction of $\boldsymbol{G}$ on the $O(1)$ timescale, (i.e. the component orthogonal to $\boldsymbol{G}$ decays exponentially). Thus, we may consider a one-dimensional problem. For such a flow, we can proceed essentially as above. On summing the variational inequalities we use the bound:

$$
\int_{0}^{H}\left(\left|v_{y}\right|^{n-1} v_{y}-\left|u_{y}\right|^{n-1} u_{y}\right)\left(v_{y}-u_{y}\right) \mathrm{d} y \geq C \int_{0}^{H}\left|v_{y}-u_{y}\right|^{n+1} \mathrm{~d} y
$$

for $C=C(n, H)$; (see Zeidler [90]). Thereafter, we may use Holder's inequality to bound $\|u-v\|_{H_{s}^{1}}$ terms and the Poincaré inequality on $W^{1, n+1}(0, H)$ for the rest.

### 2.4 Eccentric annular Hele-Shaw displacement transient model

Finally, we are in a position to derive the gap-averaged transient model. We commence with a characterization of the fluid rheologies.

### 2.4.1 Rheological assumptions

The fluids are considered to be shear thinning generalised Newtonian fluids, and typically have a yield stress. After the Hele-Shaw scaling, the principal components of the rate of strain are $\dot{\gamma}_{\phi y} \sim v_{y}$ and $\dot{\gamma}_{\xi y} \sim w_{y}$. Thus,

$$
\dot{\gamma} \sim\left[v_{y}^{2}+w_{y}^{2}\right]^{1 / 2}
$$

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The leading order shear stresses and rates of strain are related by a law of form:

$$
\begin{equation*}
\tau_{i j}=\eta(\dot{\gamma}) \dot{\gamma}_{i j}, . \quad i j=\phi y, \xi y \tag{2.49}
\end{equation*}
$$

where $\eta$ is referred to as the effective viscosity. When the fluid has a yield stress, $\eta(\dot{\gamma}) \rightarrow \infty$ as $\dot{\gamma} \rightarrow 0$, and $\tau_{i j}$ becomes indeterminate.

In $[10,57,58,59]$ it is assumed that the fluids are Herschel-Bulkley fluids:

$$
\begin{align*}
\tau_{i j} & =\left[\kappa \dot{\gamma}^{n-1}+\frac{\tau_{Y}}{\dot{\gamma}}\right] \dot{\gamma}_{i j}, & & \tau>\tau_{Y},  \tag{2.50}\\
\dot{\gamma} & =0, & & \tau \leq \tau_{Y},
\end{align*}
$$

where $\kappa, n$ and $\tau_{Y}$ are the consistency, power-law index and yield stress of the fluid, respectively, (which will depend on $\bar{c}$ ). Since the fluids are shear-thinning, we consider $0<n \leq 1$. Here we modify the Herschel-Bulkley model slightly, and will replace (2.50) with

$$
\begin{align*}
\tau_{i j} & =\left[\mu_{\infty}+\kappa \dot{\gamma}^{n-1}+\frac{\tau_{Y}}{\dot{\gamma}}\right] \dot{\gamma}_{i j}, & & \tau>\tau_{Y},  \tag{2.51}\\
\dot{\gamma} & =0, & & \tau \leq \tau_{Y},
\end{align*}
$$

where $\mu_{\infty}$ is referred to as the high-shear viscosity. There are 2 motivations for this modification.

First from the physical perspective, use of (2.50) implies an effective viscosity:

$$
\eta(\dot{\gamma})=\kappa \dot{\gamma}^{n-1}+\frac{\tau_{Y}}{\dot{\gamma}}
$$

and therefore $\eta(\dot{\gamma}) \rightarrow 0$ as $\dot{\gamma} \rightarrow \infty$. This is unrealistic, since suspension viscosities commonly approach a Newtonian plateau for large $\dot{\gamma}$, i.e. the Herschel-Bulkley model is really intended to model low-shear behaviour. The second motivation is purely mathematical. Without the highshear viscosity we are later forced to work in subspaces of $L^{1+n}$ and $W^{1,1+n}$, rather than $L^{2}$ and $H^{1}$. Although most of what we prove in the Hilbert space setting can be extended to $L^{1+n}$ and $W^{1,1+n}$, it is less convenient and involves additional analysis. From the numerical perspective this is anyway unnecessary, since numerically we end up working in finite dimensional subspaces, which lie in both $\left\{L^{1+n}, W^{1,1+n}\right\}$ and in $\left\{L^{2}, H^{1}\right\}$.

In practice we shall suppose that $\mu_{\infty}$ is small, by which we mean that at typical shear rates,

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(which are $O_{S}(1)$ due to the scaling), we expect

$$
\mu_{\infty} \ll \kappa \dot{\gamma}^{n-1}+\frac{\tau_{Y}}{\dot{\gamma}} .
$$

In this sense we may regard $\mu_{\infty}$ as a small regularisation parameter. We emphasise however that this is not the common regularisation that is used frequently to avoid proper treatment of the yield stress. Here it is vital that the yield stress behaviour is retained, since occurrence of static fluid regions in the annulus is commonplace in cementing.

## Fluid concentration effects

Finally, if we have intermediate fluid concentrations occurring in the annulus, it will be necessary to specify the physical and rheological properties of the mixture. We consider this purely as a problem of model closure. For simplicity we use linear interpolation of the physical properties. Whilst this is certainly not realistic for all fluids combinations, there is no universally accepted closure. Nothing that follows in the model derivation or analysis depends explicitly on the closure law for the mixture, i.e. provided the closure is reasonably smooth and therefore measured mixture properties could be utilised later. The numerical simulation results evidently will depend on the mixture law adopted.

$$
\begin{align*}
\tau_{Y}(\bar{c}) & =\tau_{Y, 2} \bar{c}+\tau_{Y, 1}(1-\bar{c}), \quad \kappa(\bar{c})=\kappa_{2} \bar{c}+\kappa_{1}(1-\bar{c}),  \tag{2.52}\\
n(\bar{c}) & =n_{2} \bar{c}+n_{1}(1-\bar{c}), \quad \rho(\bar{c})=\rho_{2} \bar{c}+\rho_{1}(1-\bar{c}) . \tag{2.53}
\end{align*}
$$

### 2.4.2 Hele-Shaw model derivation

We first integrate (2.16) across the annular gap to eliminate $u$ :

$$
0=\frac{\partial}{\partial \phi}[H \bar{v}]+\frac{\partial}{\partial \xi}[H \bar{w}],
$$

which prompts definition of a stream function $\Psi(\phi, \xi, t)$, for the gap-averaged flow:

$$
\begin{equation*}
H \bar{w}=\frac{\partial \Psi}{\partial \phi}, \quad H \bar{v}=-\frac{\partial \Psi}{\partial \xi} \tag{2.54}
\end{equation*}
$$

Turning now to the system (2.29), we observe that the pressure and gravitational terms do not vary with $y$. Therefore, integrating twice with respect to $y$ we have:

$$
\begin{equation*}
\boldsymbol{u}_{a}(y, \phi, \xi, t)=\left[\boldsymbol{G}_{a}-\rho \overline{\boldsymbol{u}}_{a, t}\right] \int_{y}^{H} \frac{\tilde{y}}{\eta(\tilde{y})} \mathrm{d} \tilde{y} \tag{2.55}
\end{equation*}
$$

where $\boldsymbol{u}=(v, w)$ and

$$
\begin{equation*}
\boldsymbol{G}_{a}=\left(-\frac{\partial p}{\partial \phi}+\frac{\rho \sin \beta \sin \pi \phi}{S t^{*}},-\frac{\partial p}{\partial \xi}-\frac{\rho \cos \beta}{S t^{*}}\right) . \tag{2.56}
\end{equation*}
$$

Therefore $\boldsymbol{u}_{a}$ is instantaneously parallel to $\boldsymbol{G}_{a}-\rho \overline{\boldsymbol{u}}_{a, t}$. Writing $s(y)=\left|\boldsymbol{u}_{a}\right|(y)$ and $A=$ $\left|\boldsymbol{G}_{a}-\rho \overline{\boldsymbol{u}}_{a, t}\right|$, at each fixed $(\phi, \xi, t)$, the speed $s$ is related to the modified pressure gradient $A$ via the one-dimensional shear flow problem:

$$
-A=\frac{\mathrm{d}}{\mathrm{~d} y}\left[\eta\left(\left|\frac{\mathrm{~d} s}{\mathrm{~d} y}\right|\right) \frac{\mathrm{d} s}{\mathrm{~d} y}\right] .
$$

For generalised Newtonian fluids this problem is straightforwardly solved, either analytically, or numerically by simple quadrature. However, we are interested in the gap-averaged speed:

$$
\begin{align*}
|\nabla \Psi| & =H \bar{s}=\int_{0}^{H} s(y) \mathrm{d} y=-\int_{0}^{H} y \frac{\mathrm{~d} s}{\mathrm{~d} y}(y) \mathrm{d} y \\
& =\int_{0}^{H} y \dot{\gamma}(y) \mathrm{d} y=\frac{1}{A^{2}} \int_{T_{Y}}^{A H} \tau \dot{\gamma}(\tau) \mathrm{d} \tau \tag{2.57}
\end{align*}
$$

where from (2.51), for $\tau \geq \tau_{Y}$,

$$
\tau(\dot{\gamma})=\mu_{\infty} \dot{\gamma}+\kappa \dot{\gamma}^{n}+\tau_{Y}
$$

and $\dot{\gamma}(\tau)$ is obtained by inverting this monotone function. We may observe straightforwardly that $\dot{\gamma}(\tau)$ is a strictly monotone $C^{\infty}$ function of $\tau>\tau_{Y}$. If $n<1$, then as $\tau \rightarrow \tau_{Y}, \dot{\gamma}(\tau) \rightarrow 0$ and the Herschel-Bulkley term dominates:

$$
\dot{\gamma} \sim\left[\frac{\tau-\tau_{Y}}{\kappa}\right]^{1 / n} .
$$

As $\tau \rightarrow \infty$, the high-shear viscous term dominates:

$$
\dot{\gamma} \sim \frac{\tau-\tau_{Y}}{\mu_{\infty}} .
$$

In between these limits numerical inversion is needed to define $\dot{\gamma}(\tau)$. If $n=1$, then

$$
\dot{\gamma}=\frac{\tau-\tau_{Y}}{\mu_{\infty}+\kappa} \quad \forall \tau \geq \tau_{Y} .
$$

The expression (2.57) defines the closure relationship between the gap-averaged flow rate $|\nabla \Psi|$ and the modified pressure gradient $A$. Evidently, as $A \rightarrow \tau_{Y} / H$ the range of the integral vanishes and $|\nabla \Psi| \rightarrow 0$. Since the integrand, $\tau \dot{\gamma}(\tau)$, is strictly positive, increasing faster than linear and is $C^{\infty}$, for $\tau>\tau_{Y} / H$ we see that $|\nabla \Psi|(A)$ is also $C^{\infty}$ and increases strictly monotonically, for $A>\tau_{Y} / H$. Inverting, this relation, we may write $A=A(|\nabla \Psi|)$. For $|\nabla \Psi|>0$ we have that $A(|\nabla \Psi|)$ is strictly positive, (bounded strictly below by $\left.\tau_{Y} / H\right)$, and strictly monotone. It is convenient to separate the yield'stress effects. Below, we shall write

$$
A(|\nabla \Psi|)=\chi(|\nabla \Psi|)+\frac{\tau_{Y}}{H}
$$

Thus, $\chi(|\nabla \Psi|)$ represents the part of the modified pressure gradient surplus to that required to overcome the yield stress locally. We may rewrite (2.57) as:

$$
\begin{equation*}
|\nabla \Psi|=\frac{1}{\left[\chi+\tau_{Y} / H\right]^{2}} \int_{\tau_{Y}}^{\chi H+\tau_{Y}} \tau \dot{\gamma}(\tau) \mathrm{d} \tau \tag{2.58}
\end{equation*}
$$

Although we have focused on the relation between $|\nabla \Psi|$ and $A$, (equivalently $\chi$ ), we note that via the constitutive law and the limits on the integration in (2.57) \& (2.58), we have the following parametric dependency of these functions.

$$
|\nabla \Psi|=|\nabla \Psi|\left(\chi ; H, \tau_{Y}, \kappa, n, \mu_{\infty}\right), \quad \chi=\chi\left(|\nabla \Psi| ; H, \tau_{Y}, \kappa, n, \mu_{\infty}\right)
$$

Returning to (2.55), we average across the half-gap, $y \in[0, H]$, to give:

$$
\begin{equation*}
\overline{\boldsymbol{u}}_{a}=\left[\boldsymbol{G}_{a}-\rho \overline{\boldsymbol{u}}_{a, t}\right] \frac{1}{H} \int_{0}^{H} \int_{y}^{H} \frac{\tilde{y}}{\eta(\tilde{y})} \mathrm{d} \tilde{y} \mathrm{~d} y \tag{2.59}
\end{equation*}
$$

from which we see that $\overline{\boldsymbol{u}}_{a}$ is also instantaneously parallel to the vector $\left[\boldsymbol{G}-\rho \overline{\boldsymbol{u}}_{a, t}\right]$ :

$$
\begin{equation*}
\frac{\left(-\frac{\partial \Psi}{\partial \xi}, \frac{\partial \Psi}{\partial \phi}\right)}{|\nabla \Psi|}=\frac{\left(\rho \frac{\partial^{2} \Psi}{\partial \xi \partial t}-\frac{\partial p}{\partial \phi}+g_{\phi},-\rho \frac{\partial^{2} \Psi}{\partial \phi \partial t}-\frac{\partial p}{\partial \xi}+g_{\xi}\right)}{A} \tag{2.60}
\end{equation*}
$$

For $|\nabla \Psi|>0$, replacing $A$ with $\chi+\tau_{Y} / H$, this implies that

$$
\begin{aligned}
\rho \frac{\partial^{2} \Psi}{\partial \xi \partial t}-\frac{\partial p}{\partial \phi}+\frac{\rho \sin \beta \sin \pi \phi}{S t^{*}} & =-\left(\frac{\chi(|\nabla \Psi|)+\tau_{Y} / H}{|\nabla \Psi|}\right) \frac{\partial \Psi}{\partial \xi} \\
-\rho \frac{\partial^{2} \Psi}{\partial \phi \partial t}-\frac{\partial p}{\partial \xi}-\frac{\rho \cos \beta}{S t^{*}} & =\left(\frac{\chi(|\nabla \Psi|)+\tau_{Y} / H}{|\nabla \Psi|}\right) \frac{\partial \Psi}{\partial \phi}
\end{aligned}
$$

Cross-differentiating to eliminate the pressure, we finally arrive at:

$$
\begin{equation*}
\nabla \cdot\left[\rho \nabla \Psi_{t}\right]=-\nabla \cdot\left[\left(\frac{\chi(|\nabla \Psi|)+\tau_{Y} / H}{|\nabla \Psi|}\right) \nabla \Psi+\mathbf{f}\right], \tag{2.61}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathbf{f}=\left(\rho(\bar{c}) \frac{\cos \beta}{S t^{*}}, \rho(\bar{c}) \frac{\sin \beta \sin \pi \phi}{S t^{*}}\right) \tag{2.62}
\end{equation*}
$$

If $|\nabla \Psi|=0$, we may still cross-differentiate to eliminate the pressure, except that the right-hand side of this system is indeterminate. We may write:

$$
\begin{equation*}
\nabla \cdot\left[\rho \nabla \Psi_{t}\right]=-\nabla \cdot[\mathbf{S}+\mathbf{f}], \tag{2.63}
\end{equation*}
$$

where

$$
\begin{align*}
\mathbf{S}=\left(\frac{\chi(|\nabla \Psi|)+\tau_{Y} / H}{|\nabla \Psi|}\right) \nabla \Psi & \Leftrightarrow|\mathbf{S}|>\tau_{Y} / H  \tag{2.64}\\
|\nabla \Psi|=0 & \Leftrightarrow|\mathbf{S}| \leq \tau_{Y} / H . \tag{2.65}
\end{align*}
$$

We note also that:

$$
\begin{equation*}
\mathbf{S}=\left(-\rho \frac{\partial^{2} \Psi}{\partial \phi \partial t}-\frac{\partial p}{\partial \xi}-\frac{\rho \cos \beta}{S t^{*}},-\rho \frac{\partial^{2} \Psi}{\partial \xi \partial t}+\frac{\partial p}{\partial \phi}-\frac{\rho \sin \beta \sin \pi \phi}{S t^{*}}\right) \tag{2.66}
\end{equation*}
$$

Equation (2.63) is the classical formulation of the evolution problem for $\Psi$, (effectively giving the gap-averaged velocity in the annulus). We shall consider this in a more rigorous setting below in $\S 3.1$.

For some intuition, observe that if $n=1$, we consider a constant concentration, zero yield stress and assume that the annulus is concentric, then (2.63) is simply:

$$
\rho \triangle \Psi_{t}=-3\left(\kappa+\mu_{\infty}\right) \triangle \Psi
$$

with suitable boundary conditions and initial condition.

### 2.4.3 Boundary conditions

Equation (2.63) is supplemented with the following boundary conditions:

$$
\begin{equation*}
\cdot \Psi(0, \xi, t)=0 . \tag{2.67}
\end{equation*}
$$

Chapter 2. Model derivation

$$
\begin{equation*}
\Psi(1, \xi, t)=Q(t), \tag{2.68}
\end{equation*}
$$

on the wide $(\phi=0)$ and narrow ( $\phi=1$ ) sides of the annulus, respectively. Here $Q(t)=O_{S}(1)$ represents the total flow rate through the annulus, appropriately scaled.

Conditions at the ends of the annulus are harder to specify, and depend largely on the situation that we are modelling. In general we shall suppose that large variations in the fluid concentration occur away from the ends, i.e. we are interested in displacement phenomena away from the ends. If we consider our constant geometry section to be a section of the well, then appropriate are:

$$
\begin{align*}
S_{\xi} & =0 \Rightarrow \frac{\partial \Psi}{\partial \xi}(\phi, 0, t)=0  \tag{2.69}\\
S_{\xi} & =0 \Rightarrow \frac{\partial \Psi}{\partial \xi}(\phi, Z, t)=0 \tag{2.70}
\end{align*}
$$

Alternatively, if we model a lab-scale pilot experiment, we may impose e.g. the uniform inflow condition

$$
\begin{equation*}
\Psi(\phi, 0, t)=Q(t) \tag{2.71}
\end{equation*}
$$

in place of (2.69), retaining the outflow condition (2.70). Finally, if we consider $Z \gg 1$ so that the flow close to the ends of the annulus is far from any concentration variations, we may calculate appropriate one-dimensional flows at the ends, which correspond to stream functions: $\Psi_{i n}(\phi)$ at $z=0$, and $\Psi_{\text {out }}(\phi)$ at $z=Z$, respectively. We might then impose

$$
\begin{align*}
\Psi(\phi, 0, t) & =\Psi_{i n}(\phi, t)  \tag{2.72}\\
\Psi(\phi, Z, t) & =\Psi_{o u t}(\phi, t) \tag{2.73}
\end{align*}
$$

in place of (2.69) \& (2.70). The Dirichlet conditions (2.72) \& (2.73) are easiest to handle analytically, and we assume this below unless otherwise stated.

The system (2.63) with boundary conditions (2.67), (2.68), (2.72), (2.73), and equation (2.24)for the concentration will be the main object of study of the thesis. In the next chapter we will establish theoretical results that relate to the well-posedness of this model. In chapter 5 we solve this model numerically to study static mud channel removal. Chapter 4 addresses interfacial stability. For this study a fluid concentration is not as easy to work with as is an interface. Therefore, a variant of the model is derived below.

### 2.5 Interface tracking model derivation

Later in chapter 4 we consider a version of model (2.63)-(2.65) in which the fluid is separated cleanly by an interface into two fluids of different constant physical properties. Below we derive this model.

## Derivation and jump conditions at the interface

From $\S 2.4 .2$ we have,

$$
\begin{equation*}
\nabla \cdot\left[\rho \nabla \Psi_{t}\right]=-\nabla \cdot[S+f] \tag{2.74}
\end{equation*}
$$

Let us assume for the moment that $\Psi \in C^{1}(\Omega)$. Now, let us multiply (2.74) by a test function $v$ such that,

$$
v: \boldsymbol{R}^{2} \times[0, \infty) \rightarrow \boldsymbol{R}
$$

is smooth and with compact support on $\Omega$.

After multiplying (2.74) by $v$ and integrating by parts, we have

$$
\begin{equation*}
0=\int_{\Omega}\left(\rho \nabla \Psi_{t}+\boldsymbol{S}+\boldsymbol{f}\right) \cdot \nabla v \mathrm{~d} \Omega=-\int_{\Omega} \nabla \cdot\left(\rho \nabla \Psi_{t}+\boldsymbol{S}+\boldsymbol{f}\right) v \mathrm{~d} \Omega \tag{2.75}
\end{equation*}
$$

We derived this equality assuming that $\Psi \in C^{1}(\Omega)$, but note that (2.75) has meaning even if $\Psi$ is only bounded.

Suppose now that $\Psi$ is smooth on either side of a smooth curve $C$ dividing $\Omega$ into $\Omega_{1}$ and $\Omega_{2}$, where $\Omega_{1}$ is that part of $\Omega$ on the left of the curve which has only fluid 1 on it and $\Omega_{2}$ that part on the right of the curve with fluid 2 on it. Suppose that $\Psi$ is a solution of (2.74), as we will prove in chapter 3 , and assuming that its first derivatives are uniformly continuous in $\Omega_{1}$ and
$\Omega_{2}$ we can choose a test function $v$ with compact support in $\Omega_{1}$, thus (2.75) becomes,

$$
\begin{equation*}
0=\int_{\Omega_{1}}\left(\rho \nabla \Psi_{t}+\boldsymbol{S}_{1}+\boldsymbol{f}_{1}\right) \cdot \nabla v \mathrm{~d} \Omega=-\int_{\Omega_{1}} \nabla \cdot\left(\rho \nabla \Psi_{t}+\boldsymbol{S}_{1}+\boldsymbol{f}_{1}\right) v \mathrm{~d} \Omega \tag{2.76}
\end{equation*}
$$

where the subscript 1 means that we are considering the properties of fluid 1 only. Because $\Omega_{1}$ is full of fluid 1 then $\nabla \cdot f_{1}=0$, thus

$$
\begin{equation*}
\rho_{1} \Delta \Psi_{t}=-\nabla \cdot S_{1} \quad \text { in } \Omega_{1} \tag{2.77}
\end{equation*}
$$

Likewise,

$$
\begin{equation*}
\rho_{2} \Delta \Psi_{t}=-\nabla \cdot S_{2} \quad \text { in } \Omega_{2} \tag{2.78}
\end{equation*}
$$

Now let us select a test function $v$ with compact support in $\Omega$, but which does not necessarily vanish along the curve $C$. Then, using (2.75) we have

$$
\begin{align*}
0=\int_{\Omega}\left(\rho \nabla \Psi_{t}+\boldsymbol{S}+\boldsymbol{f}\right) \cdot \nabla v \mathrm{~d} \Omega & =-\int_{\Omega_{1}}\left(\rho_{1} \nabla \Psi_{t}+\boldsymbol{S}_{1}+\boldsymbol{f}_{1}\right) \cdot \nabla v \mathrm{~d} \Omega_{1} \\
& +\int_{\Omega_{2}}\left(\rho_{2} \nabla \Psi_{t}+\boldsymbol{S}_{2}+\boldsymbol{f}_{2}\right) \cdot \nabla v \mathrm{~d} \Omega_{2} \tag{2.79}
\end{align*}
$$

Now, since $v$ has compact support within $\Omega$, we have

$$
\begin{align*}
0=\int_{\Omega}\left(\rho_{1} \nabla \Psi_{t}+\boldsymbol{S}_{1}+\boldsymbol{f}_{1}\right) \cdot \nabla v \mathrm{~d} \Omega_{1} & =-\int_{\Omega_{1}}\left(\rho_{1} \Delta \Psi_{t}+\nabla \cdot\left(\boldsymbol{S}_{1}+\boldsymbol{f}_{1}\right)\right) v \mathrm{~d} \Omega_{1} \\
& +\int_{C}\left[\left(\rho_{1} \nabla \Psi_{t}+\boldsymbol{S}_{1}+\boldsymbol{f}_{1}\right) \cdot \boldsymbol{n}_{1}\right] v \mathrm{~d} l \\
& =\int_{C}\left[\left(\rho_{1} \nabla \Psi_{t}+\boldsymbol{S}_{1}+\boldsymbol{f}_{1}\right) \cdot \boldsymbol{n}_{1}\right] v \mathrm{~d} l \tag{2.80}
\end{align*}
$$

because of (2.77) and $n_{1}$ denotes the outward normal to $C$ from $\Omega_{1}$ to $\Omega_{2}$.

Similarly we have

$$
\begin{align*}
0=\int_{\Omega}\left(\rho_{2} \nabla \Psi_{t}+\boldsymbol{S}_{2}+\boldsymbol{f}_{2}\right) \cdot \nabla v \mathrm{~d} \Omega_{2} & =-\int_{\Omega_{2}}\left(\rho_{2} \Delta \Psi_{t}+\nabla \cdot\left(\boldsymbol{S}_{2}+\boldsymbol{f}_{2}\right)\right) v \mathrm{~d} \Omega_{1} \\
& +\int_{C}\left[\left(\rho_{2} \nabla \Psi_{t}+\boldsymbol{S}_{2}+\boldsymbol{f}_{2}\right) \cdot \dot{\boldsymbol{n}_{2}}\right] v \mathrm{~d} l \\
& =\int_{C}\left[\left(\rho_{2} \nabla \Psi_{t}+\boldsymbol{S}_{2}+\boldsymbol{f}_{2}\right) \cdot \boldsymbol{n}_{2}\right] v \mathrm{~d} l \tag{2.81}
\end{align*}
$$

Adding (2.80) and (2.81) we have

$$
\begin{equation*}
\left.\left(\rho_{k} \nabla \Psi_{t}+\boldsymbol{S}_{k}+\boldsymbol{f}_{k}\right) \cdot \boldsymbol{n}\right|_{1} ^{2}=0 \tag{2.82}
\end{equation*}
$$

along $C$.
Thus, our interface model can be written as follow,

$$
\begin{equation*}
\rho_{k} \Delta \Psi_{t}+\nabla \cdot \mathbf{S}_{k}=0, \quad(\phi, \xi) \in \Omega_{k}, k=1,2 \tag{2.83}
\end{equation*}
$$

where the index $k$ is used to denote the fluid. Equation (2.83) is supplemented with the following boundary conditions:

$$
\begin{gather*}
\Psi(0, \xi, t)=0  \tag{2.84}\\
\Psi(1, \xi, t)=Q(t) \tag{2.85}
\end{gather*}
$$

on the wide ( $\phi=0$ ) and narrow ( $\phi=1$ ) sides of the annulus, respectively. Here $Q(t)$ represents the total flow rate through the annulus, appropriately scaled. For much of this section we shall assume an infinitely long annulus, with suitable conditions on $\Psi$ as $\xi \rightarrow \pm \infty$. The two fluid domains, $\Omega_{k}, k=1,2$, are separated by an interface, i.e. the smooth curve $C$, that we represent as the following level set:

$$
\begin{equation*}
F(\phi, \xi, t)=0 . \tag{2.86}
\end{equation*}
$$

The function $F(\phi, \xi, t)$ satisfies a gap-averaged kinematic equation:

$$
\begin{equation*}
\frac{H}{\epsilon} \frac{\partial F}{\partial t}-\frac{\partial \Psi}{\partial \xi} \frac{\partial F}{\partial \phi}+\frac{\partial \Psi}{\partial \phi} \frac{\partial F}{\partial \xi}=0 \tag{2.87}
\end{equation*}
$$

where $H(\phi)$ is the half-gap width and $\epsilon$ denotes the ratio of viscous to convective timescales. At the interface the following two jump/continuity conditions are satisfied,

$$
\begin{align*}
\left.\Psi\right|_{1} ^{2} & =0  \tag{2.88}\\
{\left.\left[\left(\frac{\rho_{k} \cos \beta}{S t^{*}}, \frac{\rho_{k} \sin \beta \sin \pi \phi}{S t^{*}}\right)+\rho_{k} \nabla \Psi_{t}+\mathbf{S}_{k}\right] \cdot \nabla F\right|_{1} ^{2} } & =0 \tag{2.89}
\end{align*}
$$

The first of these is simply continuity of $\Psi$, whereas (2.89) gives the jump in normal derivative of $\Psi$.

## Equivalence with concentration model

For simple and suitably smooth interface configurations, the kinematic condition for the interface can be related to the concentration equation, as follows. Let us assume that for each $\xi$ and $t$ we can define

$$
\phi=\left\{\phi_{M}(\xi, t): \bar{c}\left(\phi_{M}, \xi, t\right)=0.5\right\}
$$

Assume that the region where the concentration field changes smoothly from 1 to zero is $O\left(\delta_{d}\right)$, see Figure 2.2, where $\delta_{d}$ is the diffusive layer thickness, which is of $O\left(1 / P e^{1 / 2}\right)$. We know from [85, 92] that if we consider the limit $P e \rightarrow \infty$ we approach the immiscible limit of our model, and we can consider our domain to be separated cleanly by an interface into two domains which contain fluids of different constant physical properties, see Figure 2.2. Thus we can define our interface as

$$
\phi_{i}(\xi, t) \sim \phi_{M}(\xi, t)+O\left(\delta_{d}\right) .
$$



Figure 2.2: Example of concentration field at fixed $\xi$ and $t$.

Note that if we assume that the fluids are separated cleanly by $\phi_{i}(\xi, t)$, we have that $\bar{c}(\phi, \xi, t)=1$ for $\phi \in\left[0, \phi_{i}(\xi, t)\right)$ and $\forall \xi, t$, therefore

$$
\begin{align*}
\int_{0}^{\phi_{i}}(H \bar{w} \bar{c}) \mathrm{d} \phi & =\Psi\left(\phi_{i}, \xi, t\right) \\
\Rightarrow \frac{\partial}{\partial \xi} \int_{0}^{\phi_{i}}(H \bar{w} \bar{c}) \mathrm{d} \phi & =\frac{\partial \Psi}{\partial \phi} \frac{\partial \phi_{i}}{\partial \xi}+\frac{\partial \Psi}{\partial \xi} \quad \text { at } \phi=\phi_{i} \tag{2.90}
\end{align*}
$$

and

$$
\begin{equation*}
\frac{\partial}{\partial \xi} \int_{0}^{\phi_{M}}(H \bar{w} \bar{c}) \mathrm{d} \phi=\int_{0}^{\phi_{M}} \frac{\partial(H \bar{w} \bar{c})}{\partial \xi} \mathrm{d} \phi+\frac{1}{2} \frac{\partial \Psi}{\partial \phi} \frac{\partial \phi_{i}}{\partial \xi} \quad \text { at } \phi=\phi_{M} \tag{2.91}
\end{equation*}
$$

but

$$
\begin{array}{r}
\frac{\partial}{\partial \xi} \int_{0}^{\phi_{M}}(H \bar{w} \bar{c}) \mathrm{d} \phi=\frac{\partial}{\partial \xi} \int_{0}^{\phi_{i}-O\left(\delta_{d}\right)}(H \bar{w} \bar{c}) \mathrm{d} \phi=\int_{0}^{\phi_{i}-O\left(\delta_{d}\right)} \frac{\partial(H \bar{w} \bar{c})}{\partial \xi} \mathrm{d} \phi+\frac{1}{2} \frac{\partial \Psi}{\partial \phi} \frac{\partial \phi_{i}}{\partial \xi}-O\left(\delta_{d}\right)  \tag{2.92}\\
\text { at } \phi=\phi_{M}
\end{array}
$$

Subtracting (2.90) and (2.92) we have

$$
\begin{equation*}
\int_{0}^{\phi_{M}} \frac{\partial(H \bar{w} \bar{c})}{\partial \xi} \mathrm{d} \phi+\frac{\partial}{\partial \xi} \int_{\phi_{i}-O\left(\delta_{d}\right)}^{\phi_{i}}(H \bar{w} \bar{c}) \mathrm{d} \phi=\frac{1}{2} \frac{\partial \Psi}{\partial \phi} \frac{\partial \phi_{i}}{\partial \xi}+\frac{\partial \Psi}{\partial \xi}+O\left(\delta_{d}\right) \quad \text { at } \phi=\phi_{M} \tag{2.93}
\end{equation*}
$$

Similarly, we can get

$$
\begin{align*}
\int_{0}^{\phi_{M}} \frac{1}{\epsilon} \cdot \frac{\partial(H \bar{c})}{\partial t} \mathrm{~d} \phi & =\frac{H}{2 \epsilon} \frac{\partial \phi_{i}}{\partial t}+O\left(\delta_{d}\right) \quad \text { at } \phi=\phi_{M}  \tag{2.94}\\
\int_{0}^{\phi_{M}} \frac{\partial(H \bar{v} \bar{c})}{\partial \phi} \mathrm{d} \phi & =-\frac{1}{2} \frac{\partial \Psi}{\partial \xi} \quad \text { at } \phi=\phi_{M} \tag{2.95}
\end{align*}
$$

Integrating (2.24) with respect to $\phi$ from 0 to $\phi_{M}$, taking the limit $\delta_{d} \rightarrow 0$, i.e. the immiscible limit, and substituting (2.93), (2.94) and (2.95), we get

$$
\begin{equation*}
\frac{H}{\epsilon} \frac{\partial \phi_{i}}{\partial t}+\frac{\partial \Psi}{\partial \xi}+\frac{\partial \Psi}{\partial \phi} \frac{\partial \phi_{i}}{\partial \xi}=0 \quad \text { at } \phi=\phi_{i} \tag{2.96}
\end{equation*}
$$

Therefore a natural choice for $F(\phi, \xi, t)$ defined in (2.86) will be,

$$
F(\phi, \xi, t)=\phi-\phi_{i}(\xi, t)
$$

## Chapter 3

## Qualitative results

Mathematical validation of our model is presented in this chapter. We prove that the evolution equation (2.63) is well-posed. For the Herschel-Bulkley model, regularised by the high-shear viscosity discussed in §2.4.1, we can work in a Hilbert space setting. This provides a considerable simplification to the analysis.

Now, we present an outline of the chapter.
i) In $\S 3.1$, first we present some preliminary results on the behavior of $\chi(|\nabla \Psi|)$ and $\Psi$ which will be needed for the existence and uniqueness proof. Secondly, we define the variational problem associated with the system (2.63), we also define the subgradient of the corresponding functional for the minimization problem (3.26). Finally, we follow closely the results found in [24] (section 9.6). The section referenced introduces the theory of certain nonlinear semigroups, generated by convex functions; theory also known as gradient flows. Using these results we prove that a solution for (2.63) uniquely exists and that:

$$
\Psi \in C\left([0, \infty) ; H^{1}(\Omega)\right), \quad \Psi_{t} \in L^{\infty}\left([0, \infty) ; H^{1}(\Omega)\right)
$$

ii) In $\S 3.2$ we consider the flow of a fluid with a fixed set of physical and rheological properties along the annulus, i.e. these may depend on $(\phi, \xi)$ and a fixed concentration field, i.e. $\epsilon \rightarrow$ 0 . We assume that one of the fluid properties undergoes a continuous change everywhere in the annulus, for example from a density field $\rho_{1}$ to a density field $\rho_{2}$, while the other properties remain unchanged. We present various results on continuity of our solution with respect to the rheological and physical properties of the problem. The results for

## Chapter 3. Qualitative results

the transient model are summarized below,

1. Let $\rho_{1} \rightarrow \rho_{2}$, then for $t \in[0, T]$

$$
\begin{equation*}
\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}} T+\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}(0) \geq K_{2}\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2} \tag{3.1}
\end{equation*}
$$

2. If $\tau_{Y, 1} \rightarrow \tau_{Y, 2}$, we have

$$
\begin{equation*}
\left\|\tau_{1 Y}-\tau_{2 Y}\right\|_{L^{2}} T+\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}(0) \geq C_{2}\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2} \tag{3.2}
\end{equation*}
$$

3. Finally if $\kappa_{1} \rightarrow \kappa_{2}$ we get

$$
\begin{equation*}
\left\|\kappa_{1}-\kappa_{2}\right\|_{L^{2}} T+\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}(0) \geq C_{2}\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2} \tag{3.3}
\end{equation*}
$$

where the constants do not depend on the solutions or the rheological parameter that varies, i.e. we have continuity with respect to the initial data and physical parameters.
iii) In $\S 3.3$ we derive a number of qualitative results for the steady and unsteady problems. These results do have a physical interpretation, which we give, but also serve as a useful test in verifying that our numerical procedure is correct. We show that the total energy dissipation rate in the annulus increases with any of the yield stress, consistency or high shear viscosity. Also we present the decay rate of the solution for the transient model to the solution of the steady model, taking into account that we have a fixed concentration field, i.e. $\epsilon \rightarrow 0$,

$$
\left\|\nabla\left(\Psi-\Psi_{S}\right)\right\|_{L^{2}} \leq\left\|\nabla\left(\Psi-\Psi_{S}\right)\right\|_{L^{2}}(0) e^{-K t}
$$

where

$$
K=\frac{2 C}{\rho_{\min }} \quad \text { for } \quad C \leq 1
$$

### 3.1 Existence \& uniqueness of $\Psi$

Here we shall prove existence of a solution $\Psi(\phi, \xi, t)$. We are concerned primarily with the evolution of $\Psi$ in $t$. According to (2.24) the concentration $\bar{c}$ evolves slowly on the timescale $\epsilon^{-1}$. Consequently we will regard $\bar{c}$ as fixed in time, (i.e. formally we consider the limit $\epsilon \rightarrow 0$ ), saving

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for future work the analysis of the fully coupled system. Thus, our focus is on establishing that we have a well defined evolution problem for $\Psi$.

The annular flow domain is $\Omega=(0,1) \times(0, Z)$. We state first some physically motivated assumptions, that we will assume throughout.

A1 The concentration $\bar{c}(\phi, \xi) \in L^{\infty}(\Omega)$, and is bounded by 0 and 1. The physical properties of the fluid are all smooth functions of $\bar{c} ; \mu_{\infty}, \kappa, \rho$ and $n$ are strictly positive, $\tau_{Y}$ is semi-positive. All are bounded above and in particular $n \leq 1$, (shear thinning fluids).

A2 The flow rate through the annulus, $Q(t)$, and the various pressure gradients in the annulus are bounded. With A1 above, this implies that $\left(\mathbf{S}+\rho \nabla \Psi_{t}\right) \in\left[L^{\infty}(\Omega)\right]^{2}$, since we have that

$$
\begin{equation*}
\mathbf{S}+\rho \nabla \Psi_{t}=\left(-\frac{\partial p}{\partial \xi}-\frac{\rho \cos \beta}{S t^{*}}, \frac{\partial p}{\partial \phi}-\frac{\rho \sin \beta \sin \pi \phi}{S t^{*}}\right) \tag{3.4}
\end{equation*}
$$

A3 The annulus eccentricity $e$ satisfies $0 \leq e<1$, which implies that $1+e \geq H(\phi) \geq 1-e>0$.

Since we often consider a displacement between 2 fluids, with intermediate concentrations confined to the interior, at times we shall also assume the following.

A4 As $\xi \rightarrow 0$ and $\xi \rightarrow Z$, the physical properties of the fluid, $\tau_{Y}, \mu_{\infty}, \kappa, \rho$ and $n$ all approach constant values.

### 3.1.1 Preliminary results

We commence with a number of preliminary results.

## Properties of $\chi(|\nabla \Psi|)$

Proposition 1 For $|\nabla \Psi|>0$ we have that $\chi(|\nabla \Psi|)$ is $C^{\infty}$, strictly positive and strictly monotone; $\chi(|\nabla \Psi|) \rightarrow 0$ as $|\nabla \Psi| \rightarrow 0$.

Proof

This follows from the properties of $A(|\nabla \Psi|)$. See the discussion in $\S 2.4 .2$.

## Chapter 3. Qualitative results

Proposition 2 The function $\chi(|\nabla \Psi|)$ is bounded below by $\chi_{N}(|\nabla \Psi|), \chi_{B}(|\nabla \Psi|)$ and $\chi_{H B}(|\nabla \Psi|)$, defined implicitly for $|\nabla \Psi| \geq 0$ as follows:

$$
\begin{align*}
& |\nabla \Psi|=\frac{H^{3} \chi_{N}}{3 \mu_{\infty}}  \tag{3.5}\\
& |\nabla \Psi|=\frac{H^{3} \chi_{B}^{2}}{3 \mu_{\infty}} \frac{\left(\chi_{B}+1.5 \tau_{Y} / H\right)}{\left(\chi_{B}+\tau_{Y} / H\right)^{2}},  \tag{3.6}\\
& |\nabla \Psi|=\frac{H^{m+2}}{(m+2) \kappa} \frac{\chi_{H B}^{m+1}}{\left(\chi_{H B}+\tau_{Y} / H\right)^{2}}\left(\chi_{H B}+\frac{m+2}{m+1} \frac{\tau_{Y}}{H}\right), \quad m=1 / n \tag{3.7}
\end{align*}
$$

Proof
These results follow from (2.58) and the definition of $\dot{\gamma}(\tau)$. We prove only (3.6), the others following in analogous fashion. We bound $\tau(\dot{\gamma})$ below for $\tau \geq \tau_{Y}$, by neglecting the term $\dot{\kappa} \dot{\gamma}^{n}$, i.e.

$$
\tau(\dot{\gamma}) \geq \mu_{\infty} \dot{\gamma}+\tau_{Y} \quad \Rightarrow \quad \dot{\gamma}(\tau) \leq \frac{\tau-\tau_{Y}}{\mu_{\infty}} .
$$

Inserting into (2.58):

$$
|\nabla \Psi| \leq \frac{1}{\mu_{\infty}\left[\chi+\tau_{Y} / H\right]^{2}} \int_{\tau_{Y}}^{\chi H+\tau_{Y}} \tau\left(\tau-\tau_{Y}\right) \mathrm{d} \tau=\frac{H^{3} \chi^{2}}{3 \mu_{\infty}} \frac{\left(\chi+1.5 \tau_{Y} / H\right)}{\left(\chi+\tau_{Y} / H\right)^{2}} .
$$

Therefore, $\forall|\nabla \Psi| \geq 0$ we have

$$
\frac{\chi_{B}^{2}\left(\chi_{B}+1.5 \tau_{Y} / H\right)}{\left(\chi_{B}+\tau_{Y} / H\right)^{2}} \leq \frac{\chi^{2}\left(\chi+1.5 \tau_{Y} / H\right)}{\left(\chi+\tau_{Y} / H\right)^{2}},
$$

and consequently $\chi_{B} \leq \chi$, from the monotonicity of this function. The bound with $\chi_{N}$ comes from neglecting the yield stress terms as well as the term $\kappa \dot{\gamma}^{n}$. The bound using $\chi_{H B}$ comes from neglecting only the terms $\mu_{\infty} \dot{\gamma}$.

## Remarks:

1. The above inequalities are sharp for $|\nabla \Psi|>0$.
2. The asymptotic behaviour as $|\nabla \Psi| \rightarrow \infty$ is of most interest. For $\chi_{H B}$, following [57], we have $\chi_{H B} \sim|\nabla \Psi|^{n}$, whereas evidently $\chi_{N} \sim|\nabla \Psi|$ and $\chi_{B} \sim|\nabla \Psi|$.


Figure 3.1: The function $\chi\left(|\nabla \Psi|, H, \tau_{Y}, \kappa, \mu_{\infty}, n\right)$ with parameters $\tau_{Y}=n=1, e=0.55$ and $\kappa=\mu_{\infty}=1 / 2$.
3. It is also relatively straightforward to provide an upper bound for $\chi$. For example, writing

$$
\tau(\dot{\gamma}) \leq 2 \max \left\{\mu_{\infty} \dot{\gamma}, \kappa \dot{\gamma}^{n}\right\}+\tau_{Y}
$$

implies that $\dot{\gamma}(\tau) \geq \dot{\gamma}_{m}(\tau)$, where

$$
\dot{\gamma}_{m}(\tau)=\min \left\{\frac{\tau-\tau_{Y}}{2 \mu_{\infty}},\left[\frac{\tau-\tau_{Y}}{2 \kappa}\right]^{1 / n}\right\}
$$

Therefore, defining $\chi_{M}$ implicitly via

$$
|\nabla \Psi|=\frac{1}{\left[\chi_{M}+\tau_{Y} / H\right]^{2}} \int_{\tau_{Y}}^{\chi_{M} H+\tau_{Y}} \tau \dot{\gamma}_{m}(\tau) \mathrm{d} \tau
$$

leads to $\chi(|\nabla \Psi|) \leq \chi_{M}(|\nabla \Psi|)$. Further, we can see that at large $\tau$,

$$
\dot{\gamma}_{m}(\tau)=\frac{\tau-\tau_{Y}}{2 \mu_{\infty}}
$$

so that

$$
|\nabla \Psi| \sim \frac{H^{3} \chi_{M}^{2}}{6 \mu_{\infty}} \frac{\left(\chi_{M}+1.5 \tau_{Y} / H\right)}{\left(\chi_{M}+\tau_{Y} / H\right)^{2}} \Rightarrow \chi_{M} \sim \frac{6 \mu_{\infty}|\nabla \Psi|}{H^{3}}
$$

With the lower bounds in the above proposition, this demonstrates that $\chi(|\nabla \Psi|)$ is linear in the limit $|\nabla \Psi| \rightarrow \infty$.

## Chapter 3. Qualitative results

## Behaviour of $\|\Psi\|$

Lemma 1 Provided that assumptions A1, A2 $8 \mathcal{A} 3$ hold, then the solution $\Psi$ of (2.63), satisfying boundary conditions (2:67)-(2.70), lies in the space

$$
\Psi \in L^{\infty}\left([0, \infty], H^{1}(\Omega)\right)
$$

Proof

Multiplying (2.63) by $\Psi$ and integrating over $\Omega$, using the divergence theorem:

$$
\begin{equation*}
\int_{\Omega} \rho \nabla \Psi \cdot \nabla \Psi_{t} \mathrm{~d} \Omega=\int_{\partial \Omega} \Psi\left[\rho \nabla \Psi_{t}+\mathbf{S}\right] \cdot \nu \mathrm{d} s-\int_{\Omega} \nabla \Psi \cdot \mathbf{S}-\Psi \nabla \cdot \mathbf{f} \mathrm{d} \Omega \tag{3.8}
\end{equation*}
$$

where $\boldsymbol{\nu}$ denotes the outward normal to $\Omega$. Using (3.4), we have

$$
\begin{align*}
\int_{\partial \Omega} \Psi\left[\rho \nabla \Psi_{t}+\mathbf{S}\right] \cdot \nu \mathrm{d} s & =\int_{\partial \Omega} \Psi\left(-\frac{\partial p}{\partial \xi}+g_{\xi}, \frac{\partial p}{\partial \phi}-g_{\phi}\right) \cdot \nu \mathrm{d} s \\
& \leq\left\|\nabla p-\left(g_{\phi}, g_{\xi}\right)\right\|_{L^{\infty}(\partial \Omega)} \int_{\partial \Omega}|\Psi| \mathrm{d} s \\
& \leq C_{0}\left\|\nabla p-\left(g_{\phi}, g_{\xi}\right)\right\|_{L^{\infty}(\partial \Omega)}\|\Psi\|_{H_{s}^{1}} \tag{3.9}
\end{align*}
$$

The first bound follows from assumptions A1 \& A2, using the Holder inequality. The last line follows from,

$$
\begin{aligned}
\int_{-L}^{L}\left|\frac{\partial \Psi}{\partial \xi}\right| \mathrm{d} \phi & \geq\left|\Psi_{i n}\right|-\left|\Psi_{o u t}\right| \\
\Rightarrow \int_{\Omega}\left|\frac{\partial \Psi}{\partial \xi}\right| \mathrm{d} \Omega & \geq \int_{\partial \Omega_{1}}\left|\Psi_{i n}\right| \mathrm{d} \phi+\int_{\partial \Omega_{3}}\left|\Psi_{o u t}\right| \mathrm{d} \phi \\
\Rightarrow \int_{\Omega}\left|\frac{\partial \Psi}{\partial \phi}\right| \mathrm{d} \Omega & \geq \int_{\partial \Omega_{2}}|\Psi(1, \xi, t)| \mathrm{d} \xi+\int_{\partial \Omega_{4}}|\Psi(0, \xi, t)| \mathrm{d} \xi \\
\Rightarrow 2 \int_{\Omega}|\nabla \Psi| \mathrm{d} \Omega & \geq \int_{\partial \Omega}|\Psi| \mathrm{d} s
\end{aligned}
$$

From proposition 2 we have that

$$
\begin{equation*}
\chi(|\nabla \Psi|) \geq \frac{3 \mu_{\infty}}{H^{3}}|\nabla \Psi| \tag{3.10}
\end{equation*}
$$

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Combining all this: ,

$$
\begin{align*}
\frac{\mathrm{d}}{\mathrm{~d} t} \int_{\Omega} \frac{\rho}{2}|\nabla \Psi|^{2} \mathrm{~d} \Omega \leq & C_{0}\left\|\nabla p-\left(g_{\phi}, g_{\xi}\right)\right\|_{L^{\infty}(\partial \Omega)}\|\Psi\|_{H_{s}^{1}}+\|\nabla \cdot \mathbf{f}\|_{L^{2}}\|\Psi\|_{L^{2}} \\
& -3 \inf _{\Omega}\left\{\frac{\mu_{\infty}}{H^{3}}\right\}\|\Psi\|_{H_{s}^{1}}^{2}-\inf _{\Omega}\left\{\frac{\tau_{Y}}{H}\right\}\|\Psi\|_{H_{s}^{1}} \tag{3.11}
\end{align*}
$$

where

$$
\|\Psi\|_{H_{s}^{1}}=\left[\int_{\Omega}|\nabla \Psi|^{2} \mathrm{~d} \Omega\right]^{1 / 2} .
$$

Since $\Psi=0$ along $\phi=0$, the seminorm $\|\Psi\|_{H_{s}^{1}}$ is equivalent to $\|\Psi\|_{H^{1}}$, and evidently $\|\Psi\|_{L^{2}} \leq\|\Psi\|_{H^{1}}$. Therefore, we can find constants $C_{1} \geq 0$ and $C_{2}>0$, for which

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} t} \int_{\Omega} \frac{\rho}{2}|\nabla \Psi|^{2} \mathrm{~d} \Omega \leq C_{1}\|\Psi\|_{H^{1}}-C_{2}\|\Psi\|_{H^{1}}^{2} \tag{3.12}
\end{equation*}
$$

Integrating (3.12) with respect to $t$, we can find $C_{3}>0$ for which

$$
\begin{aligned}
C_{3}\|\Psi\|_{H^{1}}^{2}(t) & \leq \inf _{\Omega}\left\{\frac{\rho}{2}\right\}\|\Psi\|_{H_{3}^{1}}^{2}(t) \\
& \leq\left.\int_{\Omega} \frac{\rho}{2}|\nabla \Psi|^{2} \mathrm{~d} \Omega\right|_{t=0}+\int_{0}^{t} C_{1}\|\Psi\|_{H^{1}}(s)-C_{2}\|\Psi\|_{H^{1}}^{2}(s) \mathrm{d} s
\end{aligned}
$$

We see that the integrand becomes negative if

$$
\|\Psi\|_{H^{1}}(t) \geq \frac{C_{1}}{C_{2}}
$$

and consequently $\|\Psi\|_{H^{1}}(t)$ is bounded for all $t>0$.

## $I(u)$ and $\partial I[u]$

We denote by $V_{I}$ the subspace of $H^{1}(\Omega)$ containing functions that satisfy boundary conditions (2.67), (2.68), (2.72) and (2.73). The space $V_{I}$ is non-empty, since for example $\Psi^{*} \in V_{I}$, where

$$
\begin{equation*}
\Psi^{*}=\Psi_{\text {in }}(\phi, t) \frac{\rho_{\text {out }}(\phi)-\rho}{\rho_{\text {out }}(\phi)-\rho_{\text {in }}(\phi)}+\Psi_{\text {out }}(\phi, t) \frac{\rho-\rho_{\text {in }}(\phi)}{\rho_{\text {out }}(\phi)-\dot{\rho}_{\text {in }}(\phi)}, \tag{3.13}
\end{equation*}
$$

and $\rho_{\text {in }}(\phi) \& \rho_{\text {out }}(\phi)$, are the density at the inflow and outflow, $\xi_{=}=0, Z$, respectively. Note that the boundary streamfunctions, $\Psi_{i n} \& \Psi_{o u t}$, must both satisfy

$$
\Psi_{\text {in }}(1, t)=\Psi_{\text {out }}(1, t)=Q(t)
$$

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$$
\Psi_{\text {in }}(0, t)=\Psi_{\text {out }}(0, t)=0 .
$$

We also denote by $V_{I, 0}$ the subspace of $H^{1}(\Omega)$ containing functions that are zero at $\partial \Omega$ and equipped with the $L^{2}(\Omega)$ inner product. Note that $V_{I, 0}$ is a Hilbert space and in fact $V_{I, 0}=$ $H_{0}^{1}(\Omega)$, but for notation we work with $V_{I, 0}$.

Obviously $V_{I}$ is an affine space, i.e.

$$
V_{I}=\Psi^{*}+V_{I, 0},
$$

for any $\Psi^{*} \in V_{I}$. For $\Psi^{*} \in V_{I}$ and $u \in V_{I, 0}$ we consider the following functional, $I(u)$ :

$$
\begin{equation*}
I[u]:=\left\{I_{1}[u]+I_{2}[u] \quad u \in V_{I, 0},\right. \tag{3.14}
\end{equation*}
$$

where

$$
\begin{equation*}
I_{k}[u]=\int_{\Omega} L_{k}(\nabla u) \mathrm{d} \Omega, \quad k=1,2, \tag{3.15}
\end{equation*}
$$

with

$$
\begin{align*}
& L_{1}(\nabla u)=\frac{1}{2} \int_{0}^{\left|\nabla \Psi^{*}+\nabla u\right|^{2}} \frac{\chi\left(s^{1 / 2}\right)}{s^{1 / 2}} \mathrm{~d} s+\nabla\left(\Psi^{*}+u\right) \cdot \mathbf{f}  \tag{3.16}\\
& L_{2}(\nabla u)=\frac{\tau_{Y}}{H}\left|\nabla \Psi^{*}+\nabla u\right| \tag{3.17}
\end{align*}
$$

Except where stated, we shall regard $\Psi^{*}$ as fixed. Later, where we make comparisons between solutions, we shall explicitly denote the dependence on $\Psi^{*}$. The subdifferential of $I, \partial I$, is defined as follows:

$$
\begin{equation*}
\partial I[u]:=\left\{v \in H_{0}^{1}(\Omega): I[w] \geq I[u]+(v, w-u) \forall w \in V_{I, 0}\right\}, \tag{3.18}
\end{equation*}
$$

where $(\cdot, \cdot)$ is the inner product in $L^{2}(\Omega)$.

Proposition 3 The functional $I(u)$ is strictly convex, proper and lower semi-continuous; $\partial I[u]$ is monotone.

Proof

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For convexity we may follow the same steps as in [57], (proposition 2). Note that $L(\nabla u)=L_{1}(\nabla u)+L_{2}(\nabla u)$ is convex. Therefore, (see [17]: theorem 3.1, p. 66 and theorem 3.4, p. 74$), L(\nabla u)$ is weakly lower semi-continuous. The lower semicontinuity follows from the weak lower semi-continuity. Lastly, since $I[u]$ is convex, proper and lower semi-continuous, monotonicity of $\partial I[u]$ follows from [24], (theorem 1, p. 524).

### 3.1.2 Definition of $\partial I_{1, L^{2}}[u]$ and $\partial I_{2, L^{2}}[u]$

We now consider characterization ${ }^{1}$ of $\partial I[u]_{1, L^{2}}$. We consider $u, w \in H_{0}^{1}(\Omega)$. The functional $I_{1}[u]$ is Gâteaux-differentiable. Consequently, $\partial I_{1, L^{2}}[u]=I_{1, L^{2}}^{\prime}[u]$, and we find that

$$
\begin{equation*}
I_{1}[w] \geq I_{1}[u]-\int_{\Omega}(w-u) \nabla \cdot\left(\frac{\chi\left(\left|\nabla \Psi^{*}+\nabla u\right|\right)}{\left|\nabla \Psi^{*}+\nabla u\right|}\left(\nabla \Psi^{*}+\nabla u\right)+\mathbf{f}\right) \mathrm{d} \Omega \tag{3.19}
\end{equation*}
$$

for $u \in V_{I, 0}, \forall w \in V_{I, 0}$.

If $u \in V_{I, 0}$, then $\partial I_{2, L^{2}}[u]$ is characterized by:

$$
-\nabla \cdot\left(\frac{\tau_{Y}}{H} \frac{\left(\nabla \Psi^{*}+\nabla u\right)}{\left|\nabla \Psi^{*}+\nabla u\right|}\right)
$$

(see [1]), which is set-valued for $\left|\nabla \Psi^{*}+\nabla u\right|=0$. Thus

$$
\begin{equation*}
I_{2}[w] \geq I_{2}[u]-\int_{\Omega}(w-u) \nabla \cdot\left(\frac{\tau_{Y}}{H} \frac{\left(\nabla \Psi^{*}+\nabla u\right)}{\left|\nabla \Psi^{*}+\nabla u\right|}\right) \mathrm{d} \Omega \tag{3.20}
\end{equation*}
$$

We now combine these 2 terms.

Proposition 4 Let $I[u]=I_{1}[u]+I_{2}[u]$, with $I_{1}[u]$ Gâteaux-differentiable and $\partial I_{1}[u]$ single valued. Then if $v_{1} \in \partial K_{1}[u]$

$$
v=v_{1}+v_{2} \in \partial\left(I_{1}[u]+I_{2}[u]\right) \Leftrightarrow v_{2} \in \partial I_{2}[u]
$$

Proof
${ }^{1}$ The sub script $L^{2}$ means we are working with the $L^{2}$ norm, and the first index implies e.g. $I_{1}$.
$\Rightarrow$ : Because $v_{1} \in \partial I_{1}[u]$ and $I_{1}[u]$ is Gâteaux-differentiable, we have

$$
\begin{equation*}
I_{1}[u]-I_{1}[w] \geq\left(w-u,-v_{1}\right), \quad \forall w \in V_{I, 0} \tag{3.21}
\end{equation*}
$$

i.e. because $\partial I_{1}[u]$ is single valued. We know that $v_{1}+v_{2} \in \partial\left(I_{1}[u]+I_{2}[u]\right)$, thus

$$
\begin{equation*}
I_{1}[w]-I_{1}[u]+I_{2}[w]-I_{2}[u] \geq\left(w-u, v_{1}+v_{2}\right) \forall w \in V_{I, 0} \tag{3.22}
\end{equation*}
$$

Adding (3.21) and (3.22) we get:

$$
\begin{equation*}
I_{2}[w]-I_{2}[u] \geq\left(w-u, v_{2}\right) \quad \forall w \in V_{I, 0} \tag{3.23}
\end{equation*}
$$

Therefore, $v_{2} \in \partial I_{2}[u]$.
$\Leftrightarrow$ : By assumption $\partial I_{1}[u]$ is single valued and $v_{1} \in \partial I_{1}[u]$. Therefore, because $v_{2} \in \partial I_{2}[u],(3.22)$ and (3.23) are satisfied. Hence $v_{1}+v_{2} \in \partial\left(I_{1}[u]+I_{2}[u]\right)$.

Thus, it follows that for fixed $\Psi^{*}, u \in V_{I, 0}, \forall w \in V_{I, 0}$ :

$$
\begin{equation*}
I[w] \geq I[u]-\int_{\Omega}(w-u) \nabla \cdot(\mathbf{S}[u]+\mathbf{f}) \mathrm{d} \Omega \tag{3.24}
\end{equation*}
$$

i.e. $v \in \partial I_{L^{2}}[u] \Rightarrow v \in-\nabla \cdot(\mathbf{S}[u]+\mathbf{f})$.

We may now consider the steady problem for $\Psi$, which may be written as:

$$
\begin{equation*}
\nabla \cdot(\mathbf{S}+\mathbf{f})=0 \tag{3.25}
\end{equation*}
$$

with bouindary conditions (2.67), (2.68), (2.72) and (2.73).

### 3.1.3 Steady state problem

-Theorem 1 [Steady state problem] There exists a unique solution $\Psi_{s} \in V_{I}$ to (3.25), where $\Psi=\Psi^{*}+u_{s}$, ' and $u_{s}$ is the minimiser of:

$$
\begin{equation*}
\inf _{v \in V_{t, 0}} I[v] . \tag{3.26}
\end{equation*}
$$

Proof
This is essentially the same result as in [57] except that, due to adoption of (2.51), we are now in the Hilbert space setting.

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### 3.1.4 Transient problem

To prove existence and uniqueness of the transient model we need to define the following operator. Let $\rho$ satisfy the assumptions A1-A3, and consider the elliptic problem

$$
E[z]=-v
$$

where

$$
\begin{equation*}
E[z]=\nabla \cdot[\rho \nabla z], \quad \text { in } \Omega, \quad z=0 \text { on } \partial \Omega . \tag{3.27}
\end{equation*}
$$

Because $\rho_{\text {min }} \leq \rho(\bar{c}) \leq \rho_{\text {max }}, E$ is an strictly elliptic operator with bounded coefficients, thus for $v \in H^{-1}(\Omega)$, this problem has a unique solution $z \in H_{0}^{1}(\Omega)$, (see [29], Th. 8.3 and Corollary 8.7 pg . 171), and we define $E^{-1}: H^{-1}(\Omega) \mapsto H_{0}^{1}(\Omega)$ as this solution, i.e. $E^{-1}[v] \doteq z$, i.e. if $\rho=$ constant, then $E^{-1}=\triangle^{-1}$.

Note that because $\nabla \cdot(\mathbf{S}[u]+\mathbf{f})$ is the characterization of $\partial I_{L^{2}}[u]$, this means that $\nabla \cdot(\mathbf{S}[u]+\mathbf{f}) \subset$ $L^{2}(\Omega)$, which is compactly embedded in $H^{-1}(\Omega)$.

From (3.24) we can write:

$$
\begin{aligned}
\int_{\Omega}(w-u) \nabla \cdot(\mathbf{S}[u]+\mathbf{f}) \mathrm{d} \Omega & =\int_{\Omega}(w-u) E E^{-1} \nabla \cdot(\mathbf{S}[u]+\mathbf{f}) \mathrm{d} \Omega \\
& =\int_{\Omega}(w-u) \nabla \cdot\left[\rho \nabla\left(E^{-1} \nabla \cdot(\mathbf{S}[u]+\mathbf{f})\right)\right] \mathrm{d} \Omega
\end{aligned}
$$

integrating by parts we have:

$$
\int_{\Omega}(w-u) \nabla \cdot(\mathbf{S}[u]+\mathbf{f}) \mathrm{d} \Omega=-\int_{\Omega} \nabla(w-u) \rho \nabla\left(E^{-1} \nabla \cdot(\mathbf{S}[u]+\mathbf{f})\right) \mathrm{d} \Omega,
$$

because of the definition of $E^{-1}$.

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We may note that since $0<\rho_{\min } \leq \rho \leq \rho_{\max }$, the associated seminorm defined by the inner product, $\langle x, x\rangle$

$$
\langle v, w\rangle_{V_{\rho, 0}}=\int_{\Omega} \rho \nabla v \cdot \nabla w \mathrm{~d} \Omega, \quad \forall v, w \in V_{I, 0},
$$

is equivalent to $\|\cdot\|_{H^{1}}$, because of the zero boundary conditions. Note that this is just a weighted seminorm in $H_{0}^{1}(\Omega)$ which we shall denote $\|\cdot\|_{V_{\rho, 0}}$.

Therefore the subdifferential of $I[u]$ with $V_{\rho, 0}$ seminorm and inner product is defined as:

$$
\partial I_{\dot{H}_{p}^{1}}[u]:=\left\{v \in V: I[w] \geq I[u]+\langle v, w-u\rangle \forall w \in V_{I, 0}\right\} .
$$

With characterization:

$$
I[w] \geq I[u]+\int_{\Omega} \rho \nabla(w-u) \nabla\left(E^{-1} \nabla \cdot(\mathbf{S}[u]+\mathbf{f})\right) \mathrm{d} \Omega,
$$

i.e. $v \in \partial I_{\dot{H}_{\rho}^{1}}[u] \Rightarrow v \in E^{-1} \nabla \cdot(\mathbf{S}[u]+\mathbf{f})$.

Proposition 5 If $u \in D\left(\partial I_{L^{2}}\right)$ then $u \in D\left(\partial I_{V_{0,0}}\right)$.

Proof
Suppose $u \in D\left(\partial I_{L^{2}}\right)$; which implies that $\exists v \in \partial I_{L^{2}}[u]$. Thus

$$
\begin{aligned}
& I[w] \geq I[u]-\int_{\Omega}(w-u) v \mathrm{~d} \Omega \quad \forall w \in V_{I, 0} \\
\Rightarrow & I[w] \geq I[u]-\int_{\Omega}(w-u) E E^{-1} v \mathrm{~d} \Omega \quad \forall w \in V_{I, 0} \\
\Rightarrow & I[w] \geq I[u]+\int_{\Omega} \rho \nabla(w-u) \nabla\left(E^{-1} v\right) \mathrm{d} \Omega \quad \forall w \in V_{I, 0}
\end{aligned}
$$

Because of the definition of the operator $E[z]$ we have,

$$
\Rightarrow I[w] \geq I[u]+\int_{\Omega} \rho \nabla(w-u) \nabla z \mathrm{~d} \Omega \quad \forall w \in V_{I, 0} .
$$

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Therefore $z \in \partial I_{V_{\rho, 0}}[u]$ and because $v \in \partial I_{L^{2}}[u]$ we can characterize $z$ by, $z \in$ $E^{-1} \nabla \cdot(\mathbf{S}[u]+\mathbf{f})$. By definition, $u \in D(\partial I)$, provided that $\partial I[u] \neq \emptyset$. Because $\exists z \in \partial I_{V_{\rho, 0}}[u]$ this implies that $u \in D\left(\partial I_{V_{\rho, 0}}\right)$.

### 3.1.5 Existence of a solution

We are now in a position to demonstrate existence of a solution. Consider the differential equation:

$$
\left\{\begin{array}{l}
u^{\prime}(t)+A[u(t)] \ni 0 \quad t \geq 0  \tag{3.28}\\
u(0)=u_{0}
\end{array}\right.
$$

where $u_{0} \in H$ is given and $H$ is a Hilbert space. $A=\partial I$ is the subgradient of $I$, which may be nonlinear and perhaps multi-valued. Our result follows from the application of the following result.

Theorem 2 For each $u_{0} \in D(\partial I)$ there exists a unique function

$$
u \in C([0, \infty) ; H) \text { with } u^{\prime} \in L^{\infty}(0, \infty ; H)
$$

such that,

1. $u(0)=u_{0}$,
2. $u(t) \in D(\partial I)$ for each $t>0$, and
3. $-u^{\prime}(t) \in \partial I$ for a.e. $t \geq 0$.

Proof
For a proof see [24] pages 529-533.
Remember that we have two characterizations of $\partial I$ one with the norm $\|\cdot\|_{L^{2}}$ and one with the seminorm $\|\cdot\|_{V_{p, 0}}$. We showed that $\partial I_{L^{2}}[u] \neq \emptyset$. Then, by Proposition 5 we have that if $u_{0} \in D\left(\partial I_{L^{2}}\right)$ then $u_{0} \in D\left(\partial I_{V_{\rho, 0}}\right)$.

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Therefore, applying Theorem 2 to this initial condition, we have that there exists a unique function

$$
u \in C\left([0, \infty) ; V_{I, 0}\right) \text { with } u^{\prime} \in L^{\infty}\left(0, \infty ; V_{I, 0}\right),
$$

such that,

1. $u(0)=u_{0}$,
2. $u(t) \in D\left(\partial I_{V_{p, 0}}\right)$ for each $t>0$, and
3. $-u^{\prime}(t) \in \partial I_{V_{p, 0}}$ for a.e. $t \geq 0$.

This implies that $-u^{\prime}(t) \in E^{-1} \nabla \cdot(\mathbf{S}[u]+\mathbf{f})$, thus

$$
-u^{\prime}(t) \in E^{-1} \nabla \cdot(\mathbf{S}[u]+\mathbf{f}) \Rightarrow-E\left[u^{\prime}(t)\right] \in \nabla \cdot(\mathbf{S}[u]+\mathbf{f}) .
$$

Therefore (2.63) has a unique solution $u \in C\left([0, \infty) ; V_{I, 0}\right)$ and, (see earlier comments), $V_{I, 0}$ is $H_{0}^{1}(\Omega)$.

### 3.2 Continuity

### 3.2.1 Preliminary results

We commence with some qualitative results for the derivative of $\chi$ with respect to some rheological properties.

Behaviour of $\frac{\partial \chi}{\partial T_{Y}}, \frac{\partial X}{\partial \kappa}$
Proposition 6 For a fixed gap-averaged speed, $|\nabla \Psi|$, and given $\tau_{Y}^{*} \in\left[\tau_{Y, 1}, \tau_{Y, 2}\right]$, then

$$
\begin{equation*}
\frac{\partial \chi}{\partial \tau_{Y}} \leq \frac{\chi}{\tau_{Y}} . \tag{3.29}
\end{equation*}
$$

Proof
Now $A=\chi(|\nabla \Psi|)+\tau_{Y} / H$ is given implicitly by:

$$
|\nabla \Psi|=\frac{1}{A^{2}} \int_{\tau_{Y}}^{A H} \tau \dot{\gamma}(\tau) \mathrm{d} \tau .
$$

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Because $|\nabla \Psi|$ is fixed the derivative with respect to any parameter $q$ is:

$$
\begin{align*}
0 & =\frac{-2 A^{\prime}}{A^{3}} \int_{\tau_{Y}}^{A H} \tau \dot{\gamma}(\tau) \mathrm{d} \tau+\frac{A^{\prime} H^{2} \dot{\gamma}(A H)}{A}+\frac{1}{A^{2}} \int_{\tau_{Y}}^{A H} \tau \dot{\gamma}^{\prime}(\tau) \mathrm{d} \tau \\
& =\frac{A^{\prime}}{A^{3}} \int_{\tau_{Y}}^{A H} \tau^{2} \frac{\mathrm{~d}}{\mathrm{~d} \tau} \dot{\gamma}(\tau) \mathrm{d} \tau+\frac{1}{A^{2}} \int_{\tau_{Y}}^{A H} \tau \dot{\gamma}^{\prime}(\tau) \mathrm{d} \tau \tag{3.30}
\end{align*}
$$

where the ' denotes differentiation with respect to $q$. Thus,

$$
\begin{equation*}
A^{\prime}=A\left(\frac{-\int_{\tau_{Y}}^{A H} \tau \dot{\gamma}^{\prime}(\tau) \mathrm{d} \tau}{\int_{\tau_{Y}}^{A H} \tau^{2} \frac{\mathrm{~d}}{\mathrm{~d} \tau} \dot{\gamma}(\tau) \mathrm{d} \tau}\right) \tag{3.31}
\end{equation*}
$$

Differentiating the constitutive law with respect to $\tau_{Y}$,

$$
\tau(\dot{\gamma})=\mu_{\infty} \dot{\gamma}+\kappa \dot{\gamma}^{n}+\tau_{Y}
$$

at fixed $\tau$, we find:

$$
\frac{\partial \dot{\gamma}}{\partial \tau_{Y}}=-\frac{1}{\mu_{\infty}+\kappa n \dot{\gamma}^{n-1}}<0
$$

note that this is true for every $\tau \in\left(\tau_{Y}, A H\right]$. Differentiating the constitutive law with respect to $\tau$ we have:

$$
\frac{\partial \dot{\gamma}}{\partial \tau}=\frac{1}{\mu_{\infty}+\kappa n \dot{\gamma}^{n-1}}
$$

Let

$$
f(\dot{\gamma})=\mu_{\infty}+\kappa n \dot{\gamma}^{n-1}
$$

then

$$
\begin{align*}
A^{\prime} & \leq \frac{A}{\tau_{Y}}  \tag{3.32}\\
\Rightarrow \chi^{\prime}+\frac{1}{H} & \leq \frac{\chi}{\tau_{Y}}+\frac{1}{H} \tag{3.33}
\end{align*}
$$

(3.32) holds because $\tau(\dot{\gamma}) \geq \tau_{Y}$. Therefore

$$
\begin{equation*}
\chi^{\prime} \leq \frac{\chi}{\tau_{Y}} \cdot \square \tag{3.34}
\end{equation*}
$$

## Chapter 3. Qualitative results

Proposition 7 For a fixed gap-averaged speed, $|\nabla \Psi|$, and given $\kappa^{*} \in\left[\kappa_{1}, \kappa_{2}\right]$, then

$$
\begin{equation*}
\frac{\partial \chi}{\partial \kappa} \leq \frac{\chi}{\kappa}+\frac{\tau_{Y}}{H \kappa} . \tag{3.35}
\end{equation*}
$$

Proof
Differentiating the constitutive law with respect to $\kappa$,

$$
\tau(\dot{\gamma})=\mu_{\infty} \dot{\gamma}+\kappa \dot{\gamma}^{n}+\tau_{Y}
$$

at fixed $\tau$, we find:

$$
\frac{\partial \dot{\gamma}}{\partial \kappa}=-\frac{\dot{\gamma}^{n}}{\mu_{\infty}+\kappa n \dot{\gamma}^{n-1}}<0
$$

note that this is true for every $\tau \in\left(\tau_{Y}, A H\right]$. Recall that,

$$
\frac{\partial \dot{\gamma}}{\partial \tau}=\frac{1}{\mu_{\infty}+\kappa n \dot{\gamma}^{n-1}} .
$$

As in Proposition 6, let

$$
f(\dot{\gamma})=\mu_{\infty}+\kappa n \dot{\gamma}^{n-1},
$$

and note that

$$
\frac{\tau}{\kappa} \geq \dot{\gamma}^{n}
$$

Thus,

$$
\begin{align*}
A^{\prime} & \leq \frac{A}{\kappa}  \tag{3.36}\\
\Rightarrow \chi^{\prime} & \leq \frac{\chi}{\kappa}+\frac{\tau_{Y}}{H \kappa} \tag{3.37}
\end{align*}
$$

Chapter 3. Qualitative results
Proposition 8 For a fixed gap-averaged speed, $|\nabla \Psi|$, and given $\mu_{\infty}^{*} \in\left[\mu_{\infty, 1}, \mu_{\infty, 2}\right]$, then

$$
\begin{equation*}
\frac{\partial \chi}{\partial \mu_{\infty}} \leq \frac{\chi}{\mu_{\infty}}+\frac{\tau_{Y}}{H \mu_{\infty}} \tag{3.38}
\end{equation*}
$$

Proof
Differentiating the constitutive law with respect to $\kappa$,

$$
\tau(\dot{\gamma})=\mu_{\infty} \dot{\gamma}+\kappa \dot{\gamma}^{n}+\tau_{Y}
$$

at fixed $\tau$, we find:

$$
\frac{\partial \dot{\gamma}}{\partial \mu_{\infty}}=-\frac{\dot{\gamma}}{\mu_{\infty}+\kappa n \dot{\gamma}^{n-1}}<0,
$$

note that this is true for every $\tau \in\left(\tau_{Y}, A H\right]$. The proof follows in an analogous way as in Proposition (7),

$$
\begin{align*}
A^{\prime} & \leq \frac{A}{\mu_{\infty}}  \tag{3.39}\\
\Rightarrow \chi^{\prime} & \leq \frac{\chi}{\mu_{\infty}}+\frac{\tau_{Y}}{H \mu_{\infty}} \tag{3.40}
\end{align*}
$$

Proposition 9 Let $A, B \in \boldsymbol{R}^{2}$, then

$$
\begin{equation*}
(|A|+|B|)^{2-p}\left(|A|^{p-2}(A, A-B)+|B|^{p-2}(B, B-A)\right) \geq C|A-B|^{2} \tag{3.41}
\end{equation*}
$$

for $C \leq 1$ and $1<p \leq 2$.

Proof
From now on we will denote $a=|A|, b=|B|, c=|A-B|$. Because we have an inner product we can write:

$$
(A, A-B)=|a||c| \cos \alpha
$$

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where $\alpha$ is the angle between $A$ and $A-B$. In a similar way we have:

$$
\begin{array}{r}
(B, B-A)=|b||c| \cos \beta \\
(A, B)=|a||b| \cos \gamma
\end{array}
$$

where $\beta$ is the angle between $B$ and $A-B$ and $\gamma$ the angle between $A$ and $B$.


Figure 3.2: a) $\gamma>\pi / 2$, b) $\beta \leq \pi / 4, \gamma \leq \pi / 2$, c) $\beta>\pi / 4, \gamma \leq \pi / 2$.

Case a) $\gamma>\pi / 2$, Figure 3.2a.
Clearly $(A, A-B)>0,(B, B-A)>0$, then

$$
|A-\dot{B}|^{2}=(A, A-B)+(B, B-A)
$$

and because $(A, A-B)>0$ and $(B, B-A)>0$ we have:

$$
\begin{align*}
& (A, A-B)=|A|^{p-2}(A, A-B)|A|^{2-p} \leq|A|^{p-2}(A, A-B)(|A|+|B|)^{2-p} \\
& (B, B-A)=|B|^{p-2}(B, B-A)|B|^{2-p} \leq|B|^{p-2}(B, B-A)(|A|+|B|)^{2-p} \tag{3.42}
\end{align*}
$$

adding (3.43) and (3.43) we have the result with $C=1$.

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Case b) $\beta \leq \pi / 4, \gamma \leq \pi / 2$, Figure 3.2b.
For this case, clearly (see Figure 3.3)


Figure 3.3: $\beta \leq \pi / 4, \gamma \leq \pi / 2$

$$
b>a+d=a+\cos \tilde{\alpha} c>a+\frac{c}{2},
$$

where $\tilde{\alpha}=\pi-\alpha$.
Thus,

$$
\begin{aligned}
&\left.\frac{(|A|}{}+|B|\right)^{2-p}\left(|A|^{p-2}(A, A-B)+|B|^{p-2}(B, B-A)\right) \\
& \quad=(a+b)^{2-p}\left(a^{p-1} \cos \alpha+b^{p-1} \cos \beta\right) \\
& \quad \geq(a+b)^{2-p}\left(b^{p-1} \cos \beta-a^{p-1} \cos \beta\right) \\
&=(a+b)^{2-p}(b-a)(p-1) \tilde{b}^{p-2} \cos \beta \quad \text { for } \tilde{b} \in[a, b] \\
&=\left(a^{2-p}+(2-p) \tilde{b}^{1-p} b\right)(b-a)(p-1) \tilde{b}^{p-2} \cos \beta \\
& \geq(2-p) \tilde{b}^{2-p}(b-a)(p-1) \tilde{b}^{p-2} \cos \beta \\
& \geq \frac{(b-a)(2-p)(p-1)}{2} \\
& \geq \tilde{N} c,
\end{aligned}
$$

where $\tilde{N}=1 / 4(2-p)(p-1)<1$.
Case c) $\beta>\pi / 4, \gamma \leq \pi / 2$, Figure 3.2c.

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For this case we express $\cos \beta$ as:

$$
\begin{aligned}
\cos \beta & =\cos (\tilde{\alpha}-\gamma) \\
& =\cos \tilde{\alpha} \cos \gamma+\sin \tilde{\alpha} \sin \gamma \\
& \simeq \cos \tilde{\alpha}-\frac{1}{2} \gamma^{2} \cos ^{2} \tilde{\gamma}+\sin \tilde{\alpha} \gamma \\
& \simeq \cos \tilde{\alpha}+\sin \tilde{\alpha} \gamma
\end{aligned}
$$

because $\gamma$ is small.

Then,

$$
\begin{aligned}
&\left.\frac{(|A|}{}+|B|\right)^{2-p}\left(|A|^{p-2}(A, A-B)+|B|^{p-2}(B, B-\bar{A})\right) \\
&=\frac{(a+b)^{2-p}\left(b^{p-1} \cos \beta-\left.\right|^{2}\right.}{c} \\
& \geq \frac{(a+b)^{2-p} \tilde{N}\left(b^{p-1} \gamma+\left(b^{p-1}-a^{p-1}\right)|\cos \alpha|\right)}{c} \\
& \geq \frac{b^{2-p} \tilde{N} b^{p-1} \gamma}{c} \\
&=\frac{\tilde{N} b \sin \gamma}{c} \\
&=\frac{\tilde{N} b \sin \alpha}{b} \\
& \geq \frac{\tilde{N}}{2}
\end{aligned}
$$

where $\tilde{N}=1 / 2$.

Proposition 10 Let $\Psi_{1}, \Psi_{2} \in H^{1}(\Omega)$, be two solutions for (2.63)-(2.65) for concentration fields $\bar{c}_{1}$ and $\bar{c}_{2}$ respectively, then

$$
\begin{equation*}
\int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)-\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega \geq C| | \nabla\left(\Psi_{1}-\Psi_{2}\right) \|_{L^{2}(\Omega)}^{2}, \tag{3.44}
\end{equation*}
$$

for a constant $C$ which does not depend on $\Psi_{1}$ or $\Psi_{2}$.

Proof

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Now, assume that exists $\Omega^{+} \subset \Omega$ such that

$$
\nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)>0 \text { and } \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right)>0 \text { in } \Omega^{+}
$$

then, using Proposition 2 we have

$$
\begin{aligned}
& \frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|} \nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \geq \frac{3 \mu_{1, \infty}}{H^{3}} \nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \\
& \frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|} \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right) \geq \frac{3 \mu_{2, \infty}}{H^{3}} \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right)
\end{aligned}
$$

Therefore,

$$
\begin{equation*}
\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|} \nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)+\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|} \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right) \geq C\left|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right|^{2} \tag{3.45}
\end{equation*}
$$

Without loss of generality assume that $\nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)<0$ in some region $\Omega^{-} \subset \Omega$, and just for notation let us write $\chi_{1}=\chi\left(\left|\nabla \Psi_{1}\right|\right)$ and $\chi_{2}=\chi\left(\left|\nabla \Psi_{2}\right|\right)$ thus,

$$
\begin{align*}
& \frac{\chi_{1}}{\left|\nabla \Psi_{1}\right|} \nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)+\frac{\chi_{2}}{\left|\nabla \Psi_{2}\right|} \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right)= \\
+ & \frac{1}{2} \frac{\chi_{1}}{\left|\nabla \Psi_{1}\right|^{1-\varepsilon}}\left(\left|\nabla \Psi_{1}\right|^{-\varepsilon} \nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)+\left|\nabla \Psi_{2}\right|^{-\varepsilon} \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right)\right) \\
+ & \left.\left|\nabla \Psi_{2}\right|^{-\varepsilon} \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right)\right)\left(\frac{1}{2} \frac{\chi_{2}}{\left|\nabla \Psi_{2}\right|^{1-\varepsilon}}-\frac{1}{2} \frac{\chi_{1}}{\left|\nabla \Psi_{1}\right|^{1-\varepsilon}}\right) \\
+ & \frac{1}{2} \frac{\chi_{2}}{\left|\nabla \Psi_{2}\right|^{1-\varepsilon}}\left(\left|\nabla \Psi_{1}\right|^{-\varepsilon} \nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)+\left|\nabla \Psi_{2}\right|^{-\varepsilon} \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right)\right) \\
+ & \left.\left|\nabla \Psi_{1}\right|^{-\varepsilon} \nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)\right)\left(\frac{1}{2} \frac{\chi_{1}}{\left|\nabla \Psi_{1}\right|^{1-\varepsilon}}-\frac{1}{2} \frac{\chi_{2}}{\left|\nabla \Psi_{2}\right|^{1-\varepsilon}}\right) \tag{3.46}
\end{align*}
$$

for $0<\varepsilon<1$.
Now we shall prove that

$$
\begin{equation*}
\frac{\chi}{|\nabla \Psi|^{1-\varepsilon}} \tag{3.47}
\end{equation*}
$$

is an increasing monotone function. Let us take the derivative of (3.47) with respect to $|\nabla \Psi|$, thus what we will have to prove is

$$
\chi^{\prime}-(1-\varepsilon) \frac{\chi}{|\nabla \Psi|}>0
$$

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Using the fact that $\chi$ is linear as $|\Psi| \rightarrow \infty$ (see remarks after Proposition 2) and that,

$$
\chi \sim \frac{6 \mu_{\infty}}{H^{3}}|\nabla \Psi|
$$

we have that in the limit $|\Psi| \rightarrow \infty$,

$$
\chi^{\prime} \sim \frac{6 \mu_{\infty}}{H^{3}}
$$

because $\chi \in C^{\infty}$ with respect to $|\nabla \Psi|$. Therefore,

$$
\begin{equation*}
\chi^{\prime}-(1-\varepsilon) \frac{\chi}{|\nabla \Psi|} \sim \frac{6 \mu_{\infty}}{H^{3}}(1-(1-\varepsilon))>0 \tag{3.48}
\end{equation*}
$$

Now, using Proposition 2 and (3.7) we can find that in the limit $|\Psi| \rightarrow 0$

$$
|\nabla \Psi| \sim C \chi^{m+1}
$$

thus,

$$
\begin{aligned}
\chi & \sim C|\nabla \Psi|^{1 /(m+1)} \\
\chi^{\prime} & \sim C \frac{1}{m+1}|\nabla \Psi|^{1 /(m+1)-1}
\end{aligned}
$$

then

$$
\begin{equation*}
\chi^{\prime}-(1-\varepsilon) \frac{\chi}{|\nabla \Psi|} \sim C|\nabla \Psi|^{1 /(m+1)-1}\left(\frac{1}{m+1}-(1-\varepsilon)\right) \tag{3.49}
\end{equation*}
$$

which is greater than zero if

$$
\varepsilon>1-\frac{1}{m+1}
$$

For values of $a \leq|\nabla \Psi| \leq b$, we have,

$$
\begin{aligned}
& K_{\min } \leq \chi \leq K_{\max } \\
& B_{\min } \leq \chi^{\prime} \leq B_{\max }
\end{aligned}
$$

because $\chi \in C^{\infty}$ with respect to $|\nabla \Psi|$ and we are in a closed interval, thus

$$
\begin{equation*}
\chi^{\prime}-(1-\varepsilon) \frac{\chi}{|\nabla \Psi|} \geq B_{\min }-(1-\varepsilon) K_{\max } \tag{3.50}
\end{equation*}
$$

which is greater than zero if

$$
\varepsilon>1-\frac{B_{\min }}{K_{\max }}
$$

Therefore,

$$
\frac{\chi}{|\nabla \Psi|^{1-\varepsilon}}
$$

is an increasing monotone function for some range of $\varepsilon$ sufficiently close to 1 .
Using the assumption that $\nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)<0$ in $\Omega^{-}$, we have that $\left|\nabla \Psi_{2}\right| \geq\left|\nabla \Psi_{1}\right|$ in $\Omega^{-}$, then we have

$$
\begin{align*}
& \frac{\chi_{1}}{\left|\nabla \Psi_{1}\right|} \nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)+\frac{\chi_{2}}{\left|\nabla \Psi_{2}\right|} \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right) \geq \\
& \frac{1}{2}\left(\frac{\chi_{1}}{\left|\nabla \Psi_{1}\right|^{1-\varepsilon}}+\frac{\chi_{2}}{\left|\nabla \Psi_{2}\right|^{1-\varepsilon}}\right)\left(\left|\nabla \Psi_{1}\right|^{-\varepsilon} \nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)\right. \\
+ & \left.\left|\nabla \Psi_{2}\right|^{-\varepsilon} \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right)\right) \tag{3.51}
\end{align*}
$$

Using (3.5) in Proposition 2 and Proposition 9 we have,

$$
\begin{align*}
& \frac{1}{2}\left(\frac{\chi_{1}}{\left|\nabla \Psi_{1}\right|^{1-\varepsilon}}+\frac{\chi_{2}}{\left|\nabla \Psi_{2}\right|^{1-\varepsilon}}\right)\left(\left|\nabla \Psi_{1}\right|^{-\varepsilon} \nabla \Psi_{1} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right)\right. \\
+ & \left.\left|\nabla \Psi_{2}\right|^{-\varepsilon} \nabla \Psi_{2} \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right)\right) \geq C \frac{\left|\nabla \Psi_{1}\right|^{\varepsilon}+\left|\nabla \Psi_{2}\right|^{\varepsilon}}{\left|\nabla \Psi_{1}\right|^{\varepsilon}+\left|\nabla \Psi_{2}\right|^{\varepsilon}}\left|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right|^{2} \tag{3.52}
\end{align*}
$$

Therefore, the result (3.44) holds.

### 3.2.2 Continuity with respect to rheological and physical properties

Consider the flow of a fluid with a fixed set of physical and rheological properties along the annulus, i.e. these may depend on $(\phi, \xi)$. Also, we are interested in the limit $\epsilon \rightarrow 0$, therefore the concentration is fixed in time. Suppose that one of the fluid properties varies continuously, for example from a density field $\rho_{1}$ to a density field $\rho_{2}$, while the other properties remain unchanged. We would like that our solution $\Psi_{1}$ also varies continuously to $\Psi_{2}, \Psi_{1}$ being the solution for $\rho_{1}$ and $\Psi_{2}$ the solution for $\rho_{2}$. We present below a number of such continuity results of the solution with respect to the rhelogical parameters $\rho, \tau_{Y}, \kappa$ and $\mu_{\infty}$ and the initial conditions. For the rest of this section we will assume that the boundary function $\Psi^{*}$ is Lipchitz continuous with respect to the rheological parameters.

### 3.2.3 Steady State

## Density continuity

Note that for this case the only density dependence comes from the boundary function (3.13), $\Psi^{*}$. The function $\chi$ depends on all the rheological parameters, which we will fix for this case, but density. The variational formulation of the steady state problem is:
$\int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi^{*}+\nabla u\right|\right)}{\left|\nabla \Psi^{*}+\nabla u\right|}\left(\nabla \Psi^{*}+\nabla u\right)\right) \cdot \nabla(v-u)+\frac{\tau_{Y}}{H}\left(\left|\nabla \Psi^{*}+\nabla v\right|-\left|\nabla \Psi^{*}+\nabla u\right|\right)+\mathbf{f} \cdot \nabla(v-u) \mathrm{d} \Omega \geq 0$.

Let $v=u_{1}$, meaning that $u_{1}$ is the solution of our problem for $\rho\left(c_{1}\right)$ and we denote the solution for $\rho\left(c_{2}\right) u=u_{2}$, then:

$$
\begin{equation*}
\int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)\right) \cdot \nabla\left(u_{1}-u_{2}\right)+\frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{2}^{*}+\nabla u_{1}\right|-\left|\nabla \Psi_{2}\right|\right)+\mathbf{f}_{2} \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \geq 0 \tag{3.54}
\end{equation*}
$$

Now, let $v=u_{2}$ and $u=u_{1}$, then:

$$
\begin{equation*}
\int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(u_{2}-u_{1}\right)+\frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{1}^{*}+\nabla u_{2}\right|-\left|\nabla \Psi_{1}\right|\right)+\mathrm{f}_{1} \cdot \nabla\left(u_{2}-u_{1}\right) \mathrm{d} \Omega \geq 0 . \tag{3.55}
\end{equation*}
$$

Note that,

$$
\begin{align*}
\int_{\Omega} \frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{2}^{*}+\nabla u_{1}\right|\right) \mathrm{d} \Omega & \leq \int_{\Omega} \frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{2}^{*}-\nabla \Psi_{1}^{*}\right|+\left|\nabla \Psi_{1}\right|\right) \mathrm{d} \Omega  \tag{3.56}\\
\int_{\Omega} \frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{1}^{*}+\nabla u_{2}\right|\right) \mathrm{d} \Omega & \leq \int_{\Omega} \frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{2}^{*}-\nabla \Psi_{1}^{*}\right|+\left|\nabla \Psi_{2}\right|\right) \mathrm{d} \Omega . \tag{3.57}
\end{align*}
$$

Substituting (3.56) and (3.56) into (3.54) and (3.55) respectively and adding the resulting variatonial inequalities, we get:

$$
\begin{align*}
& \int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \\
+ & 2 \int_{\Omega} \frac{\gamma_{Y}}{H}\left(\left|\nabla \Psi_{2}^{*}-\nabla \Psi_{1}^{*}\right|\right)+\left(\mathbf{f}_{2}-\mathbf{f}_{1}\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \geq 0 . \tag{3.58}
\end{align*}
$$

## Chapter 3., Qualitative results

Now, let us keep every functional in terms of $\nabla \Psi_{k}$ and $\nabla \Psi_{k}^{*}$, for $k=1,2$. Thus, we have

$$
\begin{align*}
& \int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega= \\
& \int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega \\
+ & \int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega . \tag{3.59}
\end{align*}
$$

Similarly,

$$
\begin{equation*}
\int_{\Omega}\left(\mathbf{f}_{2}-\mathbf{f}_{1}\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega=\int_{\Omega}\left(\mathbf{f}_{2}-\mathbf{f}_{1}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega+\int_{\Omega}\left(\mathbf{f}_{2}-\mathbf{f}_{1}\right) \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega \tag{3.60}
\end{equation*}
$$

Let us consider the following functional:

$$
F(v)=\frac{1}{2} \int_{0}^{\left|\nabla \Psi^{*}+\nabla v\right|} \frac{\chi\left(s^{1 / 2}\right)}{s^{1 / 2}} \mathrm{~d} s
$$

This functional is convex, lower semi-continuous, Gâteaux differentiable and

$$
F^{\prime}(v)=\frac{\chi\left(\left|\nabla \Psi^{*}+\nabla v\right|\right)}{\left|\nabla \Psi^{*}+\nabla v\right|}\left(\nabla \Psi^{*}+\nabla v\right)
$$

Using Proposition 5.4 in [22] we have:

$$
\begin{aligned}
& F\left(u_{1}\right)-F\left(u_{2}\right) \geq\left\langle F^{\prime}\left(u_{2}\right), \Psi_{1}-\Psi_{2}\right\rangle \\
& F\left(u_{2}\right)-F\left(u_{1}\right) \geq\left\langle F^{\prime}\left(u_{1}\right), \Psi_{2}-\Psi_{1}\right\rangle
\end{aligned}
$$

where

$$
\left\langle F^{\prime}\left(u_{1}\right), \Psi_{2}-\Psi_{1}\right\rangle=\int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|} \nabla \Psi\right) \cdot \nabla\left(\Psi_{2}-\Psi_{1}\right) \mathrm{d} \Omega
$$

Therefore,

$$
\begin{equation*}
\int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|} \nabla \Psi_{1}-\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|} \nabla \Psi_{2}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega \geq 0 \tag{3.61}
\end{equation*}
$$

## Chapter 3. Qualitative results

Thus from (3.58) we have

$$
\begin{align*}
& \int_{\Omega}\left(\mathbf{f}_{2}-\mathbf{f}_{1}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega+\int_{\Omega}\left(\mathbf{f}_{2}-\mathbf{f}_{1}\right) \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega+\int_{\Omega} 2 \frac{\tau_{Y}}{H}\left|\nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right)\right| \mathrm{d} \Omega \\
+ & \int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega \geq \\
& \int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|} \nabla \Psi_{1}-\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|} \nabla \Psi_{2}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega \geq 0 . \tag{3.62}
\end{align*}
$$

Using Cauchy-Schwarz inequality, A1 \& A4, the Lipchitz continuity of $\Psi^{*}$ with respect to $\rho$, the facts that

$$
\frac{\chi(|\nabla \Psi|)}{|\nabla \Psi|} \nabla \Psi \in L^{2}(\Omega),
$$

and $\Psi_{1}, \Psi_{2} \in H^{1}(\Omega),(3.62)$ becomes

$$
\begin{equation*}
C\left\|\rho_{1}-\rho_{2}\right\|_{L^{2}} \geq \int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|} \nabla \Psi_{1}-\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|} \nabla \Psi_{2}\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \geq 0 . \tag{3.63}
\end{equation*}
$$

Therefore, using the result in Proposition 10, we have

$$
\begin{equation*}
\left\|\rho_{1}-\rho_{2}\right\|_{L^{2}} \geq C\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2} . \tag{3.64}
\end{equation*}
$$

## Yield stress continuity

As stated in section $\S 2.4 .2$ we know that $\chi=\chi\left(|\nabla \Psi| ; H, \tau_{Y}, \kappa, n, \mu_{\infty}\right)$, thus we will denote

$$
\chi_{k}(|\nabla \Psi|)=\chi\left(|\nabla \Psi| ; H, \tau_{Y k}, \kappa, n, \mu_{\infty}\right) \quad \text { for } k=1,2 .
$$

Without loss of generality, assume $\tau_{Y 1}>\tau_{Y 2}$, doing Taylor expansion for $\chi_{1}(|\nabla \Psi|)$ around $\tau_{Y 2}$ and neglecting $O\left(\left(\tau_{Y 1}-\tau_{Y 2}\right)^{2}\right)$, we have:

$$
\begin{equation*}
\chi\left(\left|\nabla \Psi_{1}\right|, \tau_{Y 1}\right)=\chi\left(\left|\nabla \Psi_{1}\right|, \tau_{Y 2}\right)+\frac{\partial}{\partial \tau_{Y}} \chi\left(\left|\nabla \Psi_{1}\right|, \tau_{Y}^{*}\right)\left(\tau_{Y 1}-\tau_{Y 2}\right) . \tag{3.65}
\end{equation*}
$$

Letting $v=u_{1}$ and $u=u_{2}$ in (3.53) we have:

$$
\begin{equation*}
\int_{\Omega}\left(\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|} \nabla \Psi_{2}\right) \cdot \nabla\left(u_{1}-u_{2}\right)+\frac{\tau_{Y 2}}{H}\left(\left|\nabla \Psi_{2}^{*}+\nabla u_{1}\right|-\left|\nabla \Psi_{2}\right|\right)+\mathbf{f} \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \geq 0 . \tag{3.66}
\end{equation*}
$$

## Chapter 3. Qualitative results

Using the triangle inequality, the right hand side of (3.66) is less than or equal to

$$
\begin{equation*}
\int_{\Omega}\left(\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|} \nabla \Psi_{2}\right) \cdot \nabla\left(u_{1}-u_{2}\right)+\frac{\tau_{Y 2}}{H}\left(\left|\nabla \Psi_{2}^{*}-\nabla \Psi_{1}^{*}\right|+\left|\nabla \Psi_{1}\right|-\left|\nabla \Psi_{2}\right|\right)+\mathbf{f} \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega . \tag{3.67}
\end{equation*}
$$

Now let $v=u_{2}$ and $u=u_{1}$ in (3.53) and use the triangle inequality as in (3.67), thus

$$
\begin{equation*}
\int_{\Omega}\left(\frac{\chi_{1}\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|} \nabla \Psi_{1}\right) \cdot \nabla\left(u_{2}-u_{1}\right)+\frac{\tau_{Y 1}}{H}\left(\left|\nabla \Psi_{2}^{*}-\nabla \Psi_{1}^{*}\right|+\left|\nabla \Psi_{2}\right|-\left|\nabla \Psi_{1}\right|\right)+\mathbf{f} \cdot \nabla\left(u_{2}-u_{1}\right) \mathrm{d} \Omega \geq 0 . \tag{3.68}
\end{equation*}
$$

Adding (3.67) and (3.68) we get:

$$
\begin{align*}
& \int_{\Omega}\left(\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi_{1}\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \\
+ & \int_{\Omega} \frac{\left(\tau_{Y 2}-\tau_{Y 1}\right)}{H}\left(\left|\nabla \Psi_{1}\right|-\left|\nabla \Psi_{2}\right|\right) \mathrm{d} \Omega \\
+ & \int_{\Omega} \frac{\left(\tau_{Y 2}+\tau_{Y 1}\right)}{H}\left(\left|\nabla \Psi_{1}^{*}-\nabla \Psi_{2}^{*}\right|\right) \mathrm{d} \Omega \geq 0 . \tag{3.69}
\end{align*}
$$

Using Cauchy-Schwarz inequality and the fact that $\tau_{Y} \in L^{\infty}(\Omega)$, the Lipchitz continuity of $\Psi^{*}$ with respect to $\tau_{Y}$ and the fact that $\Psi \in H^{1}(\Omega)$, we can find upper bounds for the second and the third integrals on the left-hand side of (3.69).

$$
\begin{align*}
\int_{\Omega} \frac{\left(\tau_{Y 2}-\tau_{Y 1}\right)}{H}\left(\left|\nabla \Psi_{1}\right|-\left|\nabla \Psi_{2}\right|\right) \mathrm{d} \Omega & \leq \int_{\Omega}\left|\tau_{Y 2}-\tau_{Y 1}\right|\left|\nabla \Psi_{1}-\nabla \Psi_{2}\right| \mathrm{d} \Omega \\
& \leq\left\|\tau_{Y 1}-\tau_{Y 2}\right\|_{L^{2}}\left\|\Psi_{1}-\Psi_{2}\right\|_{\dot{H}^{1}} \\
& \leq C\left\|\tau_{Y 1}-\tau_{Y 2}\right\|_{L^{\infty}}  \tag{3.70}\\
\int_{\Omega} \frac{\left(\tau_{Y 2}+\tau_{Y 1}\right)}{H}\left|\nabla \Psi_{1}^{*}-\nabla \Psi_{2}^{*}\right| \mathrm{d} \Omega & \leq C \int_{\Omega}\left|\tau_{Y 2}-\tau_{Y 1}\right|\left(\left|\tau_{Y 2}+\tau_{Y 1}\right|\right) \mathrm{d} \Omega \\
& \leq\left\|\tau_{Y 1}-\tau_{Y 2}\right\|_{L^{2}}\left\|\tau_{Y 1}+\tau_{Y 2}\right\|_{L^{2}} \\
& \leq C\left\|\tau_{Y 1}-\tau_{Y 2}\right\|_{L^{2}} \leq C| | \tau_{Y 1}-\tau_{Y 2} \|_{L^{\infty}} \tag{3.71}
\end{align*}
$$

Substituting (3.65) into the first integral of the left-hand side of (3.69) we have:

$$
\begin{align*}
& \left\|\tau_{\gamma_{1}}-\tau_{Y 2}\right\|_{L^{2}}| | \Psi_{1}-\Psi_{2}\left\|_{\dot{H}^{1}}+C\right\| \tau_{Y 1}-\tau_{Y 2} \|_{L^{2}} \\
& \quad+\int_{\Omega}\left(\frac{\left(\tau_{Y 2}-\tau_{Y_{1} 1}\right)}{\left|\nabla \Psi_{1}\right|} \frac{\partial \chi^{*}}{\partial \tau_{Y}} \nabla \Psi_{1}\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \\
& \quad+\int_{\Omega}\left(\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi_{2}\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(\Psi_{1}^{*}-\Psi_{2}^{*}\right) \mathrm{d} \Omega \\
& \quad \geq \int_{\Omega}\left(\frac{\chi_{2}\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|} \nabla \Psi_{1}-\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|} \nabla \Psi_{2}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega \geq 0 . \tag{3.72}
\end{align*}
$$

The last inequality on (3.72) holds because now $\chi(|\nabla \Psi|)$ has the same physical parameters and we can use (3.61).

We can bound

$$
\begin{aligned}
& \int_{\Omega}\left(\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi_{2}\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(\Psi_{1}^{*}-\Psi_{2}^{*}\right) \mathrm{d} \Omega \\
+ & \int_{\Omega}\left(\frac{\left(\tau_{Y 2}-\tau_{Y 1}\right)}{\left|\nabla \Psi_{1}\right|} \frac{\partial \chi^{*}}{\partial \tau_{Y}} \nabla \Psi_{1}\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega
\end{aligned}
$$

from above by:
$\int_{\Omega 2}\left|\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi_{2}\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right|\left|\nabla \Psi_{1}^{*}-\Psi_{2}^{*}\right| \mathrm{d} \Omega+\int_{\Omega}\left|\tau_{Y 2}-\tau_{Y 1}\right|\left|\frac{\partial \chi^{*}}{\partial \tau_{Y}}\right|\left|\nabla\left(u_{1}-u_{2}\right)\right| \mathrm{d} \Omega$.

Using the result in Proposition 6, we get

$$
\begin{align*}
& \int_{\Omega}\left|\tau_{Y, 1}-\tau_{Y, 2}\right|\left|\frac{\partial \chi^{*}}{\partial \tau_{Y}}\right|\left|\nabla\left(u_{1}-u_{2}\right)\right| \mathrm{d} \Omega \leq C \int_{\Omega}\left|\tau_{Y, 1}-\tau_{Y, 2}\right|\left|\chi\left(\left|\nabla \Psi_{1}\right|\right)+\frac{1}{H}\right|\left|\nabla\left(u_{1}-u_{2}\right)\right| \mathrm{d} \Omega \\
\leq & C \int_{\Omega}\left|\tau_{Y, 1}-\tau_{Y, 2}\right| \chi\left(\left|\nabla \Psi_{1}\right|\right)\left|\nabla\left(u_{1}-u_{2}\right)\right| \mathrm{d} \Omega+C \int_{\Omega}\left|\tau_{Y, 1}-\tau_{Y, 2}\right|\left|\nabla\left(u_{1}-u_{2}\right)\right| \mathrm{d} \Omega \\
\leq & C \int_{\Omega}\left|\tau_{Y, 1}-\tau_{Y, 2}\right| \chi\left(\left|\nabla \Psi_{1}\right|\right)\left|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right| \mathrm{d} \Omega+C \int_{\Omega}\left|\tau_{Y, 1}-\tau_{Y, 2}\right| \chi\left(\left|\nabla \Psi_{1}\right|\right)\left|\nabla\left(\Psi_{1}^{*}-\Psi_{2}^{*}\right)\right| \mathrm{d} \Omega \\
+ & C \int_{\Omega}\left|\tau_{Y, 1}-\tau_{Y, 2}\right|\left|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right| \mathrm{d} \Omega+C \int_{\Omega}\left|\tau_{Y, 1}-\tau_{Y, 2}\right|\left|\nabla\left(\Psi_{1}^{*}-\Psi_{2}^{*}\right)\right| \mathrm{d} \Omega \tag{3.73}
\end{align*}
$$

## Chapter 3. Qualitative results

Using Hölder's inequality and the fact that $\tau_{Y, k, \text { min }} \leq \tau_{Y, k} \leq \tau_{Y, k, \max }$ for $k=1,2$, recall that the rheological properties approach a constant value as $\xi \rightarrow 0$ and $\xi \rightarrow Z$, we have

$$
\begin{align*}
\int_{\Omega}\left|\tau_{Y, 1}-\tau_{Y, 2}\right| \chi\left(\left|\nabla \Psi_{1}\right|\right)\left|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right| \mathrm{d} \Omega & \leq\left\|\tau_{Y, 1}-\tau_{Y, 2}\right\|_{L^{\infty}}\left\|\chi\left(\left|\nabla \Psi_{1}\right|\right)\right\|_{L^{2}}\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}} \\
& \leq C\left\|\tau_{Y, 1}-\tau_{Y, 2}\right\|_{L^{\infty}} \tag{3.74}
\end{align*}
$$

because $\Psi_{k} \in H^{1}(\Omega)$ and $\chi$ is a $C^{\infty}$ function of $|\nabla \Psi|$. Similarly,

$$
\begin{equation*}
\int_{\Omega}\left|\tau_{Y, 1}-\tau_{Y, 2}\right| \chi\left(\left|\nabla \Psi_{1}\right|\right)\left|\nabla\left(\Psi_{1}^{*}-\Psi_{2}^{*}\right)\right| \mathrm{d} \Omega \leq\left\|\tau_{Y, 1}-\tau_{Y, 2}\right\|_{L^{\infty}}^{2}\left\|\chi\left(\left|\nabla \Psi_{1}\right|\right)\right\|_{L^{2}} \leq C\left\|\tau_{Y, 1}-\tau_{Y, 2}\right\|_{L^{\infty}}^{2} \tag{3.75}
\end{equation*}
$$

because the Lipchitz continuity of $\Psi_{k}^{*}$.
Now, see that

$$
\begin{align*}
& \int_{\Omega}\left|\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi_{2}\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right|\left|\nabla \Psi_{1}^{*}-\Psi_{2}^{*}\right| \mathrm{d} \Omega \\
\leq & C\left\|\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi_{2}\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right\|_{L^{2}}\left\|\tau_{Y 1}-\tau_{Y 2}\right\|_{L^{2}} \\
\leq & C\left\|\tau_{\gamma_{Y}}-\tau_{Y 2}\right\|_{L^{2}} \leq C\left\|\tau_{Y 1}-\tau_{Y 2}\right\|_{L^{\infty}} . \tag{3.76}
\end{align*}
$$

Assume that $\left\|\tau_{Y 1}-\tau_{Y 2}\right\|_{L^{\infty}}<\delta \ll 1$ then

$$
\left\|\tau_{Y_{1}}-\tau_{Y 2}\right\|_{L^{\infty}}>\left\|\tau_{Y 1}-\tau_{Y 2}\right\|_{L^{\infty}}^{2}
$$

Using this fact and combining (3.70), (3.71), (3.74), (3.75) and (3.76) we have:

$$
\left\|\tau_{Y_{1}}-\tau_{\gamma_{2}}\right\|_{L^{\infty}} \geq \int_{\Omega}\left(\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi_{2}\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega \geq 0
$$

using Proposition 10 we have the continuity result,

$$
\begin{equation*}
\left\|\tau_{Y_{1}}-\tau_{Y 2}\right\|_{L^{\infty}} \geq C\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2} \tag{3.77}
\end{equation*}
$$

## Consistency and High-shear viscosity continuity

The results on continuity for consistency and high-shear viscosity can be derived basically in the same way as the one for yield stress continuity, thus we omit the proof and state the following results,

$$
\begin{align*}
\left\|\kappa_{1}-\kappa_{2}\right\|_{L^{\infty}} & \geq C\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2},  \tag{3.78}\\
\left\|\mu_{\infty, 1}-\mu_{\infty, 2}\right\|_{L^{\infty}} & \geq C\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2}, \tag{3.79}
\end{align*}
$$

### 3.2.4 Continuity for Evolution equation

To prove continuity with respect to the physical parameters for the transient model will use the same techniques and results from §3.2.3.

## Density continuity

The variational inequality for the evolution model is:

$$
\begin{align*}
\int_{\Omega} \rho \nabla \Psi_{t} \cdot \nabla(v-u) \mathrm{d} \Omega & +\int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi^{*}+\nabla u\right|\right)}{\left|\nabla \Psi^{*}+\nabla u\right|}\left(\nabla \Psi^{*}+\nabla u\right)\right) \cdot \nabla(v-u) \mathrm{d} \Omega \\
& +\int_{\Omega} \frac{\tau_{Y}}{H}\left(\left|\nabla \Psi^{*}+\nabla v\right|-\left|\nabla \Psi^{*}+\nabla u\right|\right)+\mathbf{f} \cdot \nabla(v-u) \mathrm{d} \Omega \geq 0 \tag{3.80}
\end{align*}
$$

Following the same technique as in $\S 3.2 .3$, we let $v=u_{1}, \dot{u}=u_{2}$ in (3.80), then letting $v=u_{2}$ and $u=u_{1}$ and adding the resulting variational inequalities we have:

$$
\begin{align*}
& \int_{\Omega} \rho_{2} \nabla \Psi_{2 t} \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega-\int_{\Omega} \rho_{1} \nabla \Psi_{1 t} \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \\
+ & \int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \\
+ & \int_{\Omega} 2 \frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{2}^{*}-\nabla \Psi_{1}^{*}\right|\right)+\left(\mathbf{f}_{2}-\mathbf{f}_{1}\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \geq 0 . \tag{3.81}
\end{align*}
$$

We already have a bound for the third integral in (3.81), the first and second integrals in (3.81) can be rewritten as:

$$
\begin{align*}
& \int_{\Omega} \rho_{2} \nabla \Psi_{2 t} \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega-\int_{\Omega} \rho_{1} \nabla \Psi_{1 t} \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \\
= & \int_{\Omega} \rho_{2}\left(\nabla \Psi_{2 t}-\nabla \Psi_{1 t}\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega+\int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \\
= & \int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega-\int_{\Omega} \rho_{2}\left(\nabla \Psi_{1 t}-\nabla \Psi_{2 t}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega \\
+ & \int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega-\int_{\Omega} \rho_{2}\left(\nabla \Psi_{1 t}-\nabla \Psi_{2 t}\right) \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega \\
= & \int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega-\frac{1}{2} \int_{\Omega} \rho_{2} \frac{\partial}{\partial t}\left|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right|^{2} \mathrm{~d} \Omega \\
+ & \int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega-\int_{\Omega} \rho_{2}\left(\nabla \Psi_{1 t}-\nabla \Psi_{2 t}\right) \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega \tag{3.82}
\end{align*}
$$

Combining (3.62) and (3.82) we have:

$$
\begin{align*}
& \int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega-\frac{1}{2} \int_{\Omega} \rho_{2} \frac{\partial}{\partial t}\left|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right|^{2} \mathrm{~d} \Omega \\
+ & \int_{\Omega}\left(\mathbf{f}_{2}-\mathbf{f}_{1}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega+\int_{\Omega} 2 \frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{2}^{*}-\nabla \Psi_{1}^{*}\right|\right) \mathrm{d} \Omega \\
+ & \int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega-\int_{\Omega} \rho_{2}\left(\nabla \Psi_{1 t}-\nabla \Psi_{2 t}\right) \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega \geq \\
& \int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)-\frac{\chi\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)\right) \cdot \nabla\left(u_{1}-u_{2}\right) \mathrm{d} \Omega \geq 0 \\
\Rightarrow & \int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega+\int_{\Omega}\left(\mathbf{f}_{2}-\mathbf{f}_{1}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega \\
+ & \int_{\Omega} 2 \frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{2}^{*}-\nabla \Psi_{1}^{*}\right|\right) \mathrm{d} \Omega+\int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega \\
- & \int_{\Omega} \rho_{2}\left(\nabla \Psi_{1 t}-\nabla \Psi_{2 t}\right) \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega \geq K \frac{\partial}{\partial t}\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2} \tag{3.83}
\end{align*}
$$

Because of condition, A. $1 \rho(\bar{c}) \in L^{\infty}(\Omega)$. Having this in mind, the fact that $\Psi_{k} \in H^{1}(\Omega)$, $\dot{\Psi}_{k}^{\prime}(t) \in L^{\infty}\left(0, \infty, H^{1}(\Omega)\right)$ for $k=1,2$, the Lipschitz continuity of $\Psi^{*}$ with respect to the

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physical parameters and using Hölder's inequality, we have:

$$
\begin{aligned}
\int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega & \leq\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}}\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}\left\|\nabla \Psi_{1 t}\right\|_{L^{2}} \\
& \leq C\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}} \\
\int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla \Psi_{1 t} \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega & \leq\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}}^{2}\left\|\nabla \Psi_{1 t}\right\|_{L^{2}} \\
& \leq C\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}}^{2} \\
\cdot \int_{\Omega}\left(\rho_{2}-\rho_{1}\right) \nabla\left(\Psi_{1, t}-\Psi_{2, t}\right) \cdot \nabla\left(\Psi_{2}^{*}-\Psi_{1}^{*}\right) \mathrm{d} \Omega & \leq\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}}^{2}\left\|\nabla\left(\Psi_{1, t}-\Psi_{2, t}\right)\right\|_{L^{2}} \\
& \leq C\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}}^{2} \\
\int_{\Omega}\left(\mathbf{f}_{2}-\mathbf{f}_{1}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega & \leq C\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}}\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}} \\
& \leq C\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}} \\
\cdot \int_{\Omega} 2 \frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{2}^{*}-\nabla \Psi_{1}^{*}\right|\right) \mathrm{d} \Omega & \leq C\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}}
\end{aligned}
$$

Combining all this we have,

$$
\begin{equation*}
\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}} \geq K \frac{\partial}{\partial t}\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2} \tag{3.84}
\end{equation*}
$$

Using Gronwall's inequality in (3.84), we have the continuity result:

$$
\begin{equation*}
\left\|\rho_{1}-\rho_{2}\right\|_{L^{\infty}} T+\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}(0) \geq K\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2} \tag{3.85}
\end{equation*}
$$

## Yield stress continuity

Using the variational inequality as above, for two different yield stress fields, $\tau_{Y, 1}, \tau_{Y, 2}$, we have

$$
\begin{align*}
& \int_{\Omega} \rho\left(\nabla \Psi_{2 t}-\nabla \Psi_{1 t}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega+\int_{\Omega} \rho\left(\nabla \Psi_{2 t}-\nabla \Psi_{1 t}\right) \cdot \nabla\left(\Psi_{1}^{*}-\Psi_{2}^{*}\right) \mathrm{d} \Omega \\
+ & \left\|\tau_{\dot{Y}_{1}}-\tau_{Y 2}\right\|_{L^{2}} \geq C \int_{\Omega}\left(\frac{\chi_{2}\left(\left|\nabla \Psi_{2}\right|\right)}{\left|\nabla \Psi_{2}\right|}\left(\nabla \Psi_{2}\right)-\frac{\chi_{2}\left(\left|\nabla \Psi_{1}\right|\right)}{\left|\nabla \Psi_{1}\right|}\left(\nabla \Psi_{1}\right)\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega \geq 0 \tag{3.86}
\end{align*}
$$

using the result in (3.77). Using the fact that $\Psi^{*}$ is Lipschitz continuous with respect to $\tau_{Y}$ and that $\Psi^{\prime}(t) \in H^{1}$, we can find an upper bound for the second integral in (3.86), thus

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$$
\begin{gather*}
\int_{\Omega} \rho\left(\nabla \Psi_{2 t}-\nabla \Psi_{1 t}\right) \cdot \nabla\left(\Psi_{1}-\Psi_{2}\right) \mathrm{d} \Omega+\int_{\Omega} \rho\left(\nabla \Psi_{2 t}-\nabla \Psi_{1 t}\right) \cdot \nabla\left(\Psi_{1}^{*}-\Psi_{2}^{*}\right) \mathrm{d} \Omega \\
\leq-\frac{\rho}{2} \frac{\partial}{\partial t}\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2}+C\left\|\tau_{1 Y}-\tau_{2 Y}\right\|_{L^{2}} \tag{3.87}
\end{gather*}
$$

Combining the results in (3.77) and (3.87), we have the following differential inequality:

$$
\begin{equation*}
\left\|\tau_{1 Y}-\tau_{2 Y}\right\|_{L^{2}} \geq C \frac{\partial}{\partial t}\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2} \tag{3.88}
\end{equation*}
$$

Using Gronwall's inequality in (3.88), we have the continuity result:

$$
\begin{equation*}
\left\|\tau_{1 Y}-\tau_{2 Y}\right\|_{L^{2}} T+\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}(0) \geq C\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2} \tag{3.89}
\end{equation*}
$$

## Consistency and High-shear viscosity continuity

As in $\S 3.2 .3$ the results on continuity for consistency and high-shear viscosity can be derived basically in the same way as the one for yield stress continuity, thus we omit the proof and we just state the results,

$$
\begin{gather*}
\left\|\kappa_{1}-\kappa_{2}\right\|_{L^{2}} T+\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}(0) \geq C\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}^{2}  \tag{3.90}\\
\left\|\mu_{\infty, 1}-\mu_{\infty, 2}\right\|_{L^{2}} T+\left\|\nabla\left(\Psi_{1}-\Psi_{2}\right)\right\|_{L^{2}}(0) \geq C\left\|\nabla\left(\Psi_{1}-\dot{\Psi_{2}}\right)\right\|_{L^{2}}^{2} \tag{3.91}
\end{gather*}
$$

### 3.3 Qualitative behaviour of solutions

Here we derive, purely formally, a number of qualitative results for the steady and unsteady problems. These results do have a physical interpretation, which we give, but also serve as a useful test in verifying that our numerical procedure is correct. First we discuss the functions $\Psi_{i n}$ and $\Psi_{\text {out }}$, and how these are computed for a given flow.

### 3.3.1 End conditions and test concentrations

The fluid concentration $\bar{c}$ manifests only in the physical properties $\rho(\bar{c}), \tau_{Y}(\bar{c}), \kappa(\bar{c}), n(\bar{c}) \&$ $\mu_{\infty}(\bar{c})$, via specified closure laws. We restrict our consideration to concentration fields $\bar{c}$ that become constant in some subdomains of $\Omega$, close to the ends $\xi=0, Z$, i.e. $\bar{c} \rightarrow 1$ as $\xi \rightarrow 0$ and $\bar{c} \rightarrow 0$ as $\xi \rightarrow Z$. This is physically sensible, when we study displacements that happen in the interior of $\Omega$. It follows that the physical properties $\rho(\bar{c}), \tau_{Y}(\bar{c}), \kappa(\bar{c}) \& n(\bar{c})$, are each constant in some neighbourhood of the ends. We denote these constant properties as follows:

$$
\begin{gathered}
\rho_{\text {in }}=\rho(1), \quad \tau_{Y, \text { in }}=\tau_{Y}(1), \quad \kappa_{\text {in }}=\kappa(1), \quad n_{\text {in }}=n(1), \quad \mu_{\infty, i n}=\mu_{\infty}(1) \\
\rho_{o u t}=\rho(0), \quad \tau_{Y, \text { out }}=\tau_{Y}(0), \quad \kappa_{o u t}=\kappa(0), \quad n_{o u t}=n(0), \quad \mu_{\infty, o u t}=\mu_{\infty}(0)
\end{gathered}
$$

The end condition $\Psi_{i n}(\phi, t)$ is the stream function that corresponds instantaneously to the steady parallel flow of a fluid with properties ( $\rho_{i n}, \tau_{Y, i n}, \kappa_{i n}, n_{i n}, \mu_{\infty, i n}$ ), through an infinitely long annulus, at flow rate $Q(t)$. To compute such a steady parallel flow, we observe that $\Psi$ and all flow variables are independent of $\xi$. Therefore, we have that

$$
\frac{\partial}{\partial \phi}\left(S_{\phi}+f_{\phi}\right)=0
$$

and since $f_{\phi}$ is a constant, this implies that $S_{\phi}$ is independent of $\phi$. Now, if $\Psi$ is independent of $\xi$, we have that:

$$
S_{\phi}=\left(\chi+\tau_{Y} / H\right) \operatorname{sgn}\left\{\frac{\partial \Psi}{\partial \phi}\right\}
$$

so that the axial speed is of one sign and that $A=\chi+\tau_{Y} / H$ is independent of $\phi$. We may recall that $A$ was the absolute value of the modified pressure gradient. We now substitute from $\oint 2.4 .2$ the closure laws $|\nabla \Psi|=|\nabla \Psi|\left(A ; H, \tau_{Y}, \kappa, n, \mu_{\infty}\right)$, and therefore write:

$$
\begin{equation*}
Q=\Psi_{i n}(1)=\int_{0}^{1} \frac{\partial \Psi_{i n}}{\partial \phi} \mathrm{~d} \phi=\int_{0}^{1}|\nabla \Psi|\left(A ; H, \tau_{Y, i n}, \kappa_{i n}, n_{i n}, \mu_{\infty, i n}\right) \mathrm{d} \phi \tag{3.92}
\end{equation*}
$$

The above is a nonlinear equation for $A$, which has a unique solution since we have seen that $|\nabla \Psi|$ increases with $A$, when the yield stress is exceeded. We denote by $A=A_{\text {in }}$ the solution of (3.92), and now define $\Psi_{i n}$ as follows:

$$
\begin{equation*}
\Psi_{i n}(\phi)=\int_{0}^{\phi} \frac{\partial \Psi_{i n}}{\partial \phi} \mathrm{~d} \phi=\int_{0}^{\phi}|\nabla \Psi|\left(A_{i n} ; H, \tau_{Y, i n}, \kappa_{i n}, n_{i n}, \mu_{\infty, i n}\right) \mathrm{d} \phi \tag{3.93}
\end{equation*}
$$

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Similarly, $\Psi_{\text {out }}(\phi, t)$ is the stream function that corresponds instantaneously to the steady parallel flow through an infinitely long annulus of a fluid with properties $\left(\rho_{o u t}, \tau_{Y, o u t}, \kappa_{o u t}, n_{o u t}, \mu_{\infty, \text { out }}\right)$, at flow rate $Q(t)$. We then define $\Psi_{\text {out }}$ in exactly the same way as we have $\Psi_{i n}$, but using the rheological properties $\left(\tau_{Y, \text { out }}, \kappa_{\text {out }}, n_{\text {out }}, \mu_{\infty, \text { out }}\right)$.

### 3.3.2 Steady state problem

In the classical setting, we characterise the steady problem as the solution $\Psi_{s} \in V_{I}$ of

$$
\begin{equation*}
0=\nabla \cdot(\mathbf{S}+\mathbf{f}) \tag{3.94}
\end{equation*}
$$

As discussed previously, this may be formulated more precisely as $\Psi_{s}=\Psi^{*}+u_{s}$, where $u_{s} \in V_{0}$ is the solution of the following minimisation:

$$
\begin{equation*}
\inf _{v \in V_{0}} I[v] \tag{3.95}
\end{equation*}
$$

where $I[u]$ is defined in (3.14). Note that if we choose $\Psi^{*}=\Psi_{s}$, then evidently $u_{s}=0$. Equally, from our consideration of the subgradient of $I[u]$, we have that $\forall w \in V_{0}$ :

$$
I[w]-I\left[u_{s}\right] \geq-\int\left(w-u_{s}\right) \nabla \cdot\left(\mathbf{S}\left[u_{s}\right]+\mathbf{f}\right) \mathrm{d} \Omega
$$

Multiplying (3.94) by $\Psi_{s}$, integrating over $\Omega$, and using the divergence theorem:

$$
\begin{equation*}
0=\int_{\partial \Omega} \Psi_{S} \mathbf{S}\left[u_{s}\right] \cdot \boldsymbol{\nu} \mathrm{d} s+\int_{\Omega} \Psi_{S} \nabla \cdot \mathbf{f}-\left[\chi\left(\left|\nabla \Psi_{S}\right|\right)+\tau_{Y} / H\right]\left|\nabla \Psi_{S}\right| \mathrm{d} \Omega \tag{3.96}
\end{equation*}
$$

from which we see that

$$
\begin{equation*}
I\left(u_{s}\right)=-\int_{\Omega} L_{4}\left(\left|\nabla \Psi_{S}\right|\right) \mathrm{d} \Omega+\int_{\partial \Omega} \Psi_{S} \mathbf{S}\left[u_{s}\right] \cdot \boldsymbol{\nu} \mathrm{d} s \tag{3.97}
\end{equation*}
$$

where

$$
\begin{equation*}
L_{4}\left(\left|\nabla \Psi_{S}\right|\right)=\left[\chi\left(\left|\nabla \Psi_{S}\right|\right)\left|\nabla \Psi_{S}\right|-\frac{1}{2} \int_{0}^{\left|\nabla \Psi_{s}\right|^{2}} \frac{\chi\left(s^{1 / 2}\right)}{s^{1 / 2}} \mathrm{~d} s\right] \geq 0 \tag{3.98}
\end{equation*}
$$

The function in (3.98) increases monotonically with $\left|\nabla \Psi_{S}\right|$ and is strictly positive for $\left|\nabla \Psi_{S}\right|>0$.

Proposition 11 The function in (3.98) increases monotonically with $\left|\nabla \Psi_{S}\right|$ and is strictly positive for $\left|\nabla \Psi_{S}\right|>0$.

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Proof

Taking the derivative of $L_{4}\left(\left|\nabla \Psi_{s}\right|\right)$ with respect to $\left|\nabla \Psi_{s}\right|$, we have:

$$
L_{4}^{\prime}=\chi^{\prime}\left(\left|\nabla \Psi_{s}\right|\right)\left|\nabla \Psi_{s}\right| \geq 0,
$$

with a minimum when $\left|\nabla \Psi_{s}\right|=0$.
Therefore $L_{4}\left(\left|\nabla \Psi_{s}\right|\right)$ is an increasing function and is strictly positive for $\left|\nabla \Psi_{s}\right|>0$.

## Comparison results

We consider a displacement between fluids 1 and 2 . The fluid properties are specified smooth functions of the concentration: $\rho(\bar{c}), \tau_{Y}(\bar{c}), \kappa(\bar{c}), n(\bar{c}) \& \mu_{\infty}(\bar{c})$, for $\bar{c} \in[0,1]$. These functions interpolate between fluid 1 properties, at $\bar{c}=1$, and fluid 2 properties at $\bar{c}=0$. For example suppose that for the same two base fluids we have two different fluid concentration fields, say $\bar{c}_{1}$ and $\bar{c}_{2}$. These may for example correspond to different stages during a displacement. The two corresponding steady solutions are denoted $\Psi_{S, 1}$ and $\Psi_{S, 2}$.

We assume that the flow rate $Q$ is identical for both flows under consideration and that the end boundary conditions on $\Psi$ are computed in accordance with the procedure described in $\S 3.3 .1$. Since the end conditions are dependent only on the (constant) geometry and on the rheological properties of the two base fluids, it follows that the same conditions, $\Psi_{\text {in }}$ and $\Psi_{o u t}$, are attained by $\Psi_{S, 1}$ and $\Psi_{S, 2}$ at each end of the annulus. Therefore, in homogenizing the boundary conditions, we may take the same $\Psi^{*}$ for each concentration field, $\bar{c}_{1}$ and $\bar{c}_{2}$. The consequence is that if $\Psi_{S, 1}=\Psi^{*}+u_{S, 1}$ and if $\Psi_{S, 2}=\Psi^{*}+u_{s, 2}$, then $u_{s, 1}$ is in the test space for $u_{S, 2}$, and vice versa.

Lemma 2 Under assumptions A1-A4, $I\left[u_{s, 1}, \bar{c}_{1}\right] \leq I\left[u_{s, 2}, \bar{c}_{2}\right]$ for any of the following situations:

1. $\tau_{Y}\left(\bar{c}_{1}\right) \leq \tau_{Y}\left(\bar{c}_{2}\right)$ a.e. $(\phi, \xi) \in \Omega$.

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2. $\kappa\left(\bar{c}_{1}\right) \leq \kappa\left(\bar{c}_{2}\right)$ a.e. $(\phi, \xi) \in \Omega$.
3. $\mu_{\infty}\left(\bar{c}_{1}\right) \leq \mu_{\infty}\left(\bar{c}_{2}\right)$ a.e. $(\phi, \xi) \in \Omega$.

In each case above, it is assumed that all other physical properties remain equal between the two fluids.

Proof
The proof rests simply on the monotonicity of $\chi(|\nabla \Psi|)+\tau_{Y} / H$ with respect to any of the above. For example, let us assume that if $\tau_{Y}\left(\bar{c}_{1}\right) \leq \tau_{Y}\left(\bar{c}_{2}\right)$ then

$$
\left[\chi(|\nabla \Psi|)+\tau_{Y} / H\right]\left(\bar{c}_{1}\right) \leq\left[\chi(|\nabla \Psi|)+\tau_{Y} / H\right]\left(\bar{c}_{2}\right) .
$$

Then we have:

$$
I\left[u_{s, 1}, \bar{c}_{1}\right]=\inf _{v \in V_{0}} I\left[v, \bar{c}_{1}\right] \leq \inf _{v \in V_{0}} I\left[v, \bar{c}_{2}\right]=I\left[u_{s, 2}, \bar{c}_{2}\right] .
$$

This result then follows, provided that $\chi(|\nabla \Psi|)+\tau_{Y} / H$ is monotone in the directions indicated, for fixed $|\nabla \Psi|$

Now $A=\chi(|\nabla \Psi|)+\tau_{Y} / H$ is given implicitly by:

$$
|\nabla \Psi|=\frac{1}{A^{2}} \int_{\tau_{Y}}^{A H} \tau \dot{\gamma}(\tau) \mathrm{d} \tau
$$

At fixed $|\nabla \Psi|$ we differentiate with respect to any parameter $q$ :

$$
\begin{aligned}
0 & =\frac{-2 A^{\prime}}{A^{3}} \int_{\tau_{Y}}^{A H} \tau \dot{\gamma}(\tau) \mathrm{d} \tau+\frac{A^{\prime} H^{2} \dot{\gamma}(A H)}{A}+\frac{1}{A^{2}} \int_{\tau_{Y}}^{A H} \tau \dot{\gamma}^{\prime}(\tau) \mathrm{d} \tau \\
& =\frac{A^{\prime}}{A^{3}} \int_{\tau_{Y}}^{A H} \tau^{2} \frac{\mathrm{~d}}{\mathrm{~d} \tau} \dot{\gamma}(\tau) \mathrm{d} \tau+\frac{1}{A^{2}} \int_{\tau_{Y}}^{A H} \tau \dot{\gamma}^{\prime}(\tau) \mathrm{d} \tau
\end{aligned}
$$

where the ' denotes differentiation with respect to $q$. The first integrand is positive, and therefore

$$
\operatorname{sgn}\left(A^{\prime}\right)=-\operatorname{sgn}\left(\int_{\tau_{Y}}^{A H} \tau \dot{\gamma}^{\prime}(\tau) \mathrm{d} \tau\right)
$$

Differentiating the constitutive law, at fixed $\tau$, we find:

$$
\begin{aligned}
\frac{\partial \dot{\gamma}}{\partial \tau_{Y}} & =-\frac{1}{\mu_{\infty}+\kappa n \dot{\gamma}^{n-1}}<0 \\
\frac{\partial \dot{\gamma}}{\partial \kappa} & =-\frac{\dot{\gamma}^{n}}{\mu_{\infty}+\kappa n \dot{\gamma}^{n-1}}<0 \\
\frac{\partial \dot{\gamma}}{\partial \mu_{\infty}} & =-\frac{\dot{\gamma}}{\mu_{\infty}+\kappa n \dot{\gamma}^{n-1}}<0
\end{aligned}
$$

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Consequently, $A$ increases with any of these parameters.

Remark: It follows that, for increasing yield stress, consistency or high-shear viscosity, the functional

$$
-\int_{\Omega} L_{4}\left(\left|\nabla \Psi_{S}\right|\right) \mathrm{d} \Omega+\int_{\partial \Omega} \Psi_{S} \mathbf{S}\left[u_{s}\right] \cdot \boldsymbol{\nu} \mathrm{d} s
$$

increases. When the ends of the annulus consist of regions of constant physical properties, the end contributions to the above boundary integral are negligible, since $\boldsymbol{S}_{\xi}=0$, and this integral simplifies to the drop in the modified pressure, along $\phi=1$, say $\triangle p$, where

$$
\Delta p=\left.\int_{0}^{L} S_{\phi}\right|_{\phi=1} \mathrm{~d} \xi=\int_{0}^{L}-\frac{\partial p}{\partial \xi}+\rho \cos \beta \mathrm{d} \xi=p(0)-p(L)+\int_{0}^{L} \rho \cos \beta \mathrm{~d} \xi
$$

Therefore we have that:

$$
-\int_{\Omega} L_{4}\left(\left|\nabla \Psi_{S}\right|, \bar{c}_{1}\right) \mathrm{d} \Omega+Q \triangle p\left(\bar{c}_{1}\right)<-\int_{\Omega} L_{4}\left(\left|\nabla \Psi_{S}\right|, \bar{c}_{2}\right) \mathrm{d} \Omega+Q \triangle p\left(\bar{c}_{2}\right)
$$

under the conditions above.

The above functional may be interpreted physically as the total energy dissipation rate in the annulus. Therefore Lemma 2 states that the dissipation rate increases with any of the yield stress, consistency or high shear viscosity. This is of course physically intuitive.

### 3.3.3 Transient problem

The main purpose of this section is to show that the solution of the transient model (2.63) goes to the solution of the pseudo-steady model (2.1) as $t \rightarrow \infty$. First we treat the problem analytically, showing that in fact $\left\|\Psi-\Psi_{s}\right\|_{L^{2}(\Omega)}$ decays as $t \rightarrow \infty$. Afterwards, we compute numerically the solution for both systems and present the behaviour of $\left\|\Psi-\Psi_{s}\right\|_{L^{2}(\Omega)}$. For simplicity we assume that the displacement in the annulus is steady and that the concentration field does not undergo any temporal changes, i.e. we have a fixed concentration field.

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## Analytical Results

For the rest of this section we assume that $\Psi^{*}=\Psi_{s}$ for the transient model. Thus, first we multiply (2.1) by $\left(\Psi-\Psi_{s}\right)$ and integrate by parts to get

$$
\begin{equation*}
\int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{s}\right|\right)}{\left|\nabla \Psi_{s}\right|} \nabla \Psi_{s}\right) \cdot \nabla\left(\Psi-\Psi_{s}\right)+\frac{\tau_{Y}}{H}\left(|\nabla \Psi|-\left|\nabla \Psi_{s}\right|\right)+\mathbf{f} \cdot \nabla\left(\Psi-\Psi_{s}\right) \mathrm{d} \Omega \geq 0 \tag{3.99}
\end{equation*}
$$

note that because $\Psi^{*}=\Psi_{s}$ then $\left(\Psi-\Psi_{s}\right) \in H_{0}^{1}(\Omega)$.

Now we multiply (2.63) by ( $\Psi_{s}-\Psi$ ) and integrate by parts, thus

$$
\begin{align*}
\int_{\Omega} \rho \nabla \Psi_{t} \cdot \nabla\left(\Psi_{s}-\Psi\right) \mathrm{d} \Omega & +\int_{\Omega}\left(\frac{\chi(|\nabla \Psi|)}{|\nabla \Psi|} \nabla \Psi\right) \cdot \nabla\left(\Psi_{s}-\Psi\right) \mathrm{d} \Omega \\
& +\int_{\Omega} \frac{\tau_{Y}}{H}\left(\left|\nabla \Psi_{s}\right|-|\nabla \Psi|\right)+\mathbf{f} \cdot \nabla\left(\Psi_{s}-\Psi\right) \mathrm{d} \Omega \geq 0 \tag{3.100}
\end{align*}
$$

Using the fact that

$$
\frac{\chi\left(\left|\nabla \Psi^{*}+\nabla v\right|\right)}{\left|\nabla \Psi^{*}+\nabla v\right|}\left(\left|\nabla \Psi^{*}+\nabla v\right|\right)
$$

is the Gateaux derivative of

$$
F(v)=\frac{1}{2} \int_{0}^{\left|\nabla \Psi^{*}+\nabla v\right|} \frac{\chi\left(s^{1 / 2}\right)}{s^{1 / 2}} \mathrm{~d} s
$$

We can use Proposition 5.4 in [22] and have the following result:

$$
\begin{aligned}
& F(\Psi)-F\left(\Psi_{s}\right) \geq\left\langle F^{\prime}\left(\Psi_{s}\right), \Psi-\Psi_{s}\right\rangle \\
& F\left(\Psi_{s}\right)-F(\Psi) \geq\left\langle F^{\prime}(\Psi), \Psi_{s}-\Psi\right\rangle
\end{aligned}
$$

where

$$
\left\langle F^{\prime}(\Psi), \Psi_{s}-\Psi\right\rangle=\int_{\Omega}\left(\frac{\chi(|\nabla \Psi|)}{|\nabla \Psi|} \nabla \Psi\right) \cdot \nabla\left(\Psi_{s}-\Psi\right) \mathrm{d} \Omega
$$

## Chapter 3. Qualitative results

Thus,

$$
\begin{aligned}
& \left\langle F^{\prime}\left(\Psi_{s}\right)-F^{\prime}(\Psi), \Psi_{s}-\Psi\right\rangle \leq 0 \\
\Rightarrow & \int_{\Omega}\left(\frac{\chi\left(\left|\nabla \Psi_{S}\right|\right)}{\left|\nabla \Psi_{S}\right|} \nabla \Psi_{s}-\frac{\chi(|\nabla \Psi|)}{|\nabla \Psi|} \nabla \Psi\right) \cdot \nabla\left(\Psi_{s}-\Psi\right) \mathrm{d} \Omega \leq-C| | \nabla\left(\Psi-\Psi_{S}\right) \|_{L^{2}}^{2}
\end{aligned}
$$

because of Proposition 10.

Therefore, adding (3.99) \& (3.100) we have

$$
\begin{equation*}
\int_{\Omega} \rho \nabla \Psi_{t} \cdot \nabla\left(\Psi_{s}-\Psi\right) \mathrm{d} \Omega-C\left\|\nabla\left(\Psi-\Psi_{S}\right)\right\|_{L^{2}}^{2} \geq 0 \tag{3.101}
\end{equation*}
$$

Because $\frac{\partial \Psi_{s}}{\partial t}=0$, actually (3.101) states that

$$
\begin{equation*}
-C\left\|\nabla\left(\Psi-\Psi_{S}\right)\right\|_{L^{2}}^{2} \geq \frac{\rho_{\min }}{2} \frac{\partial}{\partial t} \int_{\Omega}\left|\nabla\left(\Psi_{s}-\Psi\right)\right|^{2} \mathrm{~d} \Omega \tag{3.102}
\end{equation*}
$$

Therefore, using Gronwall's inequality we have

$$
\begin{equation*}
\left\|\nabla\left(\Psi-\Psi_{S}\right)\right\|_{L^{2}} \leq\left\|\nabla\left(\Psi-\Psi_{S}\right)\right\|_{L^{2}}(0) e^{-K t} \tag{3.103}
\end{equation*}
$$

where

$$
K=\frac{2 C}{\rho_{\min }}
$$

## Numerical Results

Here and for all the numerical results in the rest of the thesis we consider the case when $\mu_{\infty}=0$. This is simpler to work with than $\mu_{\infty}>0$. As stated in $\S 2.4$, when $\mu_{\infty}=0$ we are no longer in a Hilbert space setting, instead the steady model has a solution which lives in the Banach space $W^{1,1+\frac{1}{n}}(\Omega)$. The solution of the transient model (2.63), is in $H^{1}(\Omega)$ for finite time.

On the other hand, recall that numerically we work in finite dimensional subspaces of $H^{1}(\Omega)$ -in the case of the FEM we consider basis functions in $H^{1}(\Omega)$. In chapter 4 we use a Chebyshev spectral method, as an outline (in chapter 4 we fully explain the method), we express our
solution as the sum of the first $n$ Chebyshev polynomials which are $C^{\infty}$ functions. Thus, both methods use finite dimensional subspaces of $H^{1}(\Omega)$, which is compactly embedded in $W^{1,1+\frac{1}{n}}(\Omega)$. We solve the system (2.63) using an augmented lagrangian algorithm, [26, 31], which fully describes the unyielded regions of the flow, if any. For a full description of the algorithm we refer to chapter 6 . For space discretisation we use a finite element method with triangular elements and linear basis functions. For time discretisation we use the implicit Backward Euler method which comes out naturally from the augmented lagrangian formulation. As an example of the decay property of the transient solution, we compute the evolution of $\Psi(\phi, \xi, t)$ from an initial condition,

$$
\Psi(\phi, \xi, 0)=\sin (\pi \phi) \sin (\xi \pi)
$$

Figure 3.4 a) shows the decay of $\left\|\Psi-\Psi_{s}\right\|_{L^{2}(\Omega)}$ as time evolves. The rheological and physical parameters for the problem are: $\tau_{Y, 1}=1, \kappa_{1}=2, \rho_{1}=1, m_{1}=2, \tau_{Y, 2}=1, \kappa_{2}=1, \rho_{2}=1$, $m_{2}=1, e=0.4$ and $\beta=0$. Figure 3.4 b ) shows the fixed concentration field, remember that we are interested in the changes of the velocity field on the fast timescale. Rheological parameters are interpolated as in (2.53). Figure 3.5 shows the contours of the stream function $\Psi$ for times $t=1,5,10$ and the contours of the steady state solution $\Psi_{s}$. Clearly this simulation agrees with the theoretical results stated in (3.103).


Figure 3.4: a) Decay of $\left\|\Psi-\Psi_{s}\right\|_{L^{2}(\Omega)}$. b) Fixed concentration field.

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Figure 3.5: Transient decay to steady state solution, contour plots of $\Psi$ at time $t=1,5,10$, and contour plot of the steady state solution $\Psi_{s}$.

## Chapter 4

## Stability of multi layer flows

As explained in chapter 1, for certain parameter combinations the annular displacement is ineffective. The displacement front advances faster on the wide side than on the narrow side and consequently develops a long finger. At long times; this flow will resemble a parallel multilayer flow, along the length of the annulus. It is of practical interest to understand if such flows are stable or unstable and that is the aim of this chapter. We use classical methods of hydrodynamic theory to analyze the linear stability of a parallel multi-layer base flow.
(i) In $\S 4.1$ we present the steady parallel multi-layer solutions that we consider as the base flows for our stability study. We identify 4 different types of base flow:

- Both fluids are yielded at the interface.
- Both fluids are yielded at the interface but a small static channel forms in the narrow part of the annulus.
- Fluid in the narrow part of the annulus is completely static.
- Fluid in the wide part of the annulus is unyielded at the interface.

These base solutions are governed by 9 dimensionless parameters. For the case where the narrow side is completely unyielded, we find an analytical expression to determine the minimum value of the interface position such that fluid 2 will be static and unyielded whenever the interface position is greater than this value.
(ii) In $\S 4.2$ we give the general linear stability problem, which is derived by considering a perturbation about the base state. The base state and perturbation are solutions of the
interface tracking model, described in $\S 2.5$. The general stability problem will depend on 14 dimensionless parameters.
(iii) In $\S 4.3$ we present stability results for the parallel flow of a single fluid along the annulus. We show that the flow is always linearly stable.
(iv) In $\S 4.4$ we use the method of normal modes and derive the eigenvalue problem that governs stability of the general case of a parallel flow of 2 fluids. We consider the 4 types of base flow identified in §4.1. We show that linear instabilities can arise only if both fluids are yielded at the interface.
(v) In full generality, the eigenvalue problem derived in $\S 4.4$, is governed by 14 dimensionless parameters. This makes any analytical understanding difficult. Therefore, in $\S 4.5$ we consider the simplified case of a concentric annulus, firstly for 2 Newtonian fluids and secondly for 2 power law fluids. This greatly reduces the parametric dependency and we are able to derive analytical expressions for the eigenvalues. This has use both in giving insight and as a test problem for our numerical code.
(vi) In $\S 4.6$ we analyze the full problem numerically. First we describe the method used, afterwards we present stability results for yield stress fluids. As expected, if the annulus is concentric, the results are comparable to the ones for Newtonian and power law fluids in an concentric annulus. In general it appears that introducing eccentricity has destabilizing effects.
(vii) From our numerical results in the previous sections, we observed that all significant changes occur for long wavelengths. Thus in $\S 4.7$ we study the long wavelength asymptotics of the stability problem. For long wavelengths it is possible to derive a semi-analytical expression for the eigenvalues. We also present an industrial application and make some comparisons with the results found in [59]. We find that interfacial instabilities are predicted to occur only in the parameter space where the model in [59] predicts an unsteady displacement, i.e. a long finger on the wide side.
(viii) Finally, in $\S 4.8$ we investigate the effects of a spatial developing instability for a single fluid and two fluids along the annulus. It appears that any non trivial spatial perturbation admits spatially growing instabilities for the single fluid problem. For the two fluid problem we give a condition for instability in terms of two eigenvalue problems. We have not analyzed this problem further and will be considered as future work.

In section $\S 2.4 .2$, and for the qualitative results in the previous chapter, we have considered the limit $\epsilon \rightarrow 0$ in which the convective timescale is much longer than the viscous timescale. Here we wish to study the effects of transient coupling between the interface and velocity field. Therefore, we shall not assume $\epsilon \ll 1$.

### 4.1 Parallel multi-layer flows

We study the stability of steady multi-layer flows, that arise as one fluid fingers past another during an annular displacement at unit flow rate. These basic flows are parallel to the $\xi$-axis, hence $\Psi=\Psi(\phi)$. We assume that the annulus is occupied by two fluids: fluid 1 occupying $\phi \in\left[0, \phi_{i}\right)$ and fluid 2 occupying $\phi \in\left(\phi_{i}, 1\right]$. The interface is denoted $\phi=\phi_{i}$. We denote the rheological parameters of fluid $k$ as $\tau_{k, Y}, m_{k} \& \kappa_{k}$ for $k=1,2$. Our solutions are steady parallel solutions of (2.83). Hence $S_{k}=\left(S_{k, \phi}, S_{k, \xi}\right)=\left(S_{k, \phi}, 0\right)$, and (2.83) implies that:

$$
\begin{equation*}
\frac{\partial}{\partial \phi} S_{k, \phi}=0 \tag{4.1}
\end{equation*}
$$

for the basic flow. By definition we have:

$$
\begin{equation*}
S_{k, \phi}=-\frac{\partial p}{\partial z}-\frac{\rho_{k} \cos \beta}{S t^{*}} \tag{4.2}
\end{equation*}
$$

and with (2.89), this implies that the axial pressure gradient is continuous at the interface:

$$
\left.\frac{\partial p}{\partial z}\right|_{1} ^{2}=0
$$

Therefore the fluid layers are acted on by a constant pressure gradient, modified by an axial static pressure that is different in each section due to the density jump. Since $S_{k, \phi}$ is independent of $\phi$ and $\xi$, (by definition, we look for solutions independent of $\xi$ ), following [59] we may write:

$$
\begin{equation*}
S_{1, \phi}=A, \quad S_{2, \phi}=A-b \tag{4.3}
\end{equation*}
$$

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where $b$ is the buoyancy parameter, given by:

$$
\begin{equation*}
b=\frac{\rho_{2}-\rho_{1}}{S t^{*}} \cos \beta \tag{4.4}
\end{equation*}
$$

Note that typically, in a cementing scenario where fingering occurs, $b$ is negative: the heavier cement channels past the more viscous drilling mud that is left behind on the narrow side. The constant $A$ represents the modified pressure gradient in the axial direction, within fluid 1 , and must be found as part of the base solution. In order to find $A$ we use (2.85), with $Q=1$ (for a fixed unit flow rate),

$$
\begin{equation*}
1=\Psi(1)=\left.\int_{0}^{\phi_{i}} \frac{\partial \Psi}{\partial \phi}\right|_{k=1} \mathrm{~d} \phi+\left.\int_{\phi_{i}}^{1} \frac{\partial \Psi}{\partial \phi}\right|_{k=2} \mathrm{~d} \phi \tag{4.5}
\end{equation*}
$$

where, assuming the fluids are Herschel-Bulkley fluids,

$$
\begin{align*}
\left.\frac{\partial \Psi}{\partial \phi}\right|_{k=1} & =\operatorname{sgn}(A) \begin{cases}0, & |A| \leq \tau_{1, Y} / H, \\
\frac{H^{m_{1}+2}}{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)} \frac{\left(|A|-\tau_{1, Y} / H\right)^{m_{1}+1}}{|A|^{2}}\left[|A|+\frac{\tau_{1, Y} / H}{m_{1}+1}\right] \\
|A|>\tau_{1, Y} / H,\end{cases}  \tag{4.6}\\
\left.\frac{\partial \Psi}{\partial \phi}\right|_{k=2}= & \operatorname{sgn}(A-b) \begin{cases}0, & |A-b| \leq \tau_{2, Y} / H, \\
\frac{H^{m_{2}+2}\left(|A-b|-\frac{\tau_{2, Y}}{H}\right)^{m_{2}+1}}{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)|A-b|^{2}}\left[|A-b|+\frac{\tau_{2, Y} / H}{m_{2}+1}\right] \\
& |A-b|>\tau_{2, Y} / H .\end{cases} \tag{4.7}
\end{align*}
$$

see [59]. Equation (4.5) is a nonlinear equation for $A$. It is straightforward to show that the integrals on the right-hand side increase strictly monotonically with $A$ and that (4.5) has a unique solution $A=A\left(\phi_{i}\right)$, depending only on the interface position, on $b$, and on the rheological parameters of the two fluids. Given each $A\left(\phi_{i}\right)$, we can construct $\Psi$ from (4.6) \& (4.7), by integrating with respect to $\phi$ from $\phi=0$.

We shall denote the basic solution, constructed as above, by $\Psi_{k, 0}$, (or simply $\Psi_{0}$ if only a single fluid is considered). In general $\Psi_{k, 0}(\phi)$ is continuous in $\phi$, with a jump in the first derivative (hence the axial velocity) at $\phi_{i}$. We present some characteristic basic flows below.

From (4.6) \& (4.7) we can see that the basic flow depends on 9 parameters, $\tau_{Y, k}, \kappa_{k}, m_{k}$, for $k=1,2, b, e$ and $\phi_{i}$. The buoyancy itself depends on 4 parameters, $\rho_{k}, S t^{*}$ and $\beta$. Although these do not appear in the base flow description, these parameters will appear individually in the stability problem that we derive later. In addition, this stability problem will contain a wavenumber and the time scale ratio $\epsilon$. Thus, in total our stability problem has a 14 parameter dependencies and this is the main difficulty of extracting useful information from the results. For simplicity we will consider the buoyancy as one parameter, and for the rest of the thesis we fix the Stokes number as $S t^{*}=0.1$.

Figure 4.1 shows the base parallel flow of two generalised Newitonian fluids with rheological parameters, $\tau_{Y, 1}=2, \kappa_{1}=2, m_{1}=2, \tau_{Y, 2}=1, \kappa_{2}=1, m_{2}=1, b=-9.5$, the eccentricity is 0.4 and the position of the interface, denoted as $\phi_{i}$, is at $\phi=0.5$. The inclination of the well is $\beta=\pi / 10$. As we can see, both fluids are fully yielded, the pressure gradient and buoyancy forces are large enough to over come both $\tau_{Y_{1}}$ and $\tau_{\gamma_{2}}$. The stream function is continuous and a discontinuity on the axial velocity is observed at the interface.

Figure 4.2 shows the basic parallel flow of a couple of fluids with rheological parameters, $\tau_{Y, 1}=$ $0.5, \kappa_{1}=1, m_{1}=2, \tau_{Y, 2}=0.6, \kappa_{2}=1, m_{2}=1, b=-9.5$, the eccentricity is 0.5 and $\phi_{i}=0.5$. The inclination of the well is $\beta=\pi / 10$. The pressure gradient and buoyancy forces are not sufficiently large to fully over come $\tau_{Y 2}$ along the narrow part of the annulus and a small mud channel forms.

Figure 4.3 shows a worst case scenario. Here the rheological parameters are, $\tau_{Y, 1}=1, \kappa_{1}=1$, $m_{1}=2, \tau_{Y, 2}=10, \kappa_{2}=1, m_{2}=1, b=-28.5$, the eccentricity is 0.8 and $\phi_{i}=0.5$. The inclination of the well is $\beta=\pi / 10$. Again, we increase the eccentricity of the annulus to 0.8 . and clearly the pressure gradient and buoyancy forces are not large enough to overcome $\tau_{Y 2}$ and fluid 2 if completely static.

Figure 4.4 is of purely mathematical interest, to our knowledge no experimental work has reported a case like this. The rheological parameters are, $\tau_{Y, 1}=7, \kappa_{1}=7, m_{1}=2, \tau_{Y, 2}=0.2$, $\kappa_{2}=1, m_{2}=1, b=0$, the eccentricity is 0.2 and $\phi_{i}=0.5$. For this case we have a fully vertical

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annulus. Note that if we increase the eccentricity, the region where fluid 1 is flowing gets wider and the plug zone will break.



Figure 4.1: Multilayer base flow, both fluids yielded at the interface. Rhelogical and physical parameters are: $\tau_{Y, 1}=2, \kappa_{1}=2, \rho_{1}=2, m_{1}=2, \tau_{Y, 2}=1, \kappa_{2}=1, \rho_{2}=1, m_{2}=1, e=0.4$, $\phi_{i}=0.5$ and $\beta=\pi / 10$.



Figure 4.2: Multilayer base flow, both fluids yielded at the interface, mud channel in the narrow side of the annulus. Rhelogical and physical parameters are: $\tau_{Y, 1}=0.5, \kappa_{1}=1, \rho_{1}=2, m_{1}=2$, $\tau_{Y, 2}=0.6, \kappa_{2}=1, \rho_{2}=1, m_{2}=1, e=0.5, \phi_{i}=0.5$ and $\beta=\pi / 10$.

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Figure 4.3: Multilayer base flow, fluid 2 unyielded at the interface. Rhelogical and physical parameters are: $\tau_{Y, 1}=1, \kappa_{1}=1, \rho_{1}=4, m_{1}=2, \tau_{Y, 2}=10, \kappa_{2}=1, \rho_{2}=1, m_{2}=1, e=0.8$, $\phi_{i}=0.5$ and $\beta=\pi / 10$.



Figure 4.4: Multilayer base flow, fluid 1 unyielded at the interface. Rhelogical and physical parameters are: $\tau_{Y, 1}=7, \kappa_{1}=7, \rho_{1}=1, m_{1}=2, \tau_{Y, 2}=0.2, \kappa_{2}=1, \rho_{2}=1, m_{2}=1, e=0.2$, $\phi_{i}=0.5$ and $\beta=0$.

### 4.1.1 Static mud channels

For certain combinations of dimensionless parameters, it is possible for one of the fluids to be stationary, as we can see in Fig. 4.3. Physically, this occurs when the modified pressure gradient $A$, (or $A-b$ in fluid 2 ), is not large enough to overcome the yield stress of the fluid. Since the modified pressure gradient is constant in each fluid layer and since $H(\phi)$ decreases, yield stress fluids become static at the largest values of $\phi$ in each layer, i.e. where the gap is smallest. It is therefore possible to have flows for which the fluid 1 layer or the fluid 2 layer is either partly or fully static. From an industrial perspective the flows within this category that have most interest are those in which fluid 2 is static, either partly or fully, i.e. there is a static mud channel.

Let us consider a basic multi-layer flow of two fluids with specified physical properties, computed as described above, and how the type of flow may vary for different choices of $\phi_{i}$. Since the minimal annular gap is at $\phi=1$, where $H=1-e$, there can only be a static "mud" channel in fluid 2 if there is a static layer as $\phi_{i} \rightarrow 1$. The condition for this to happen is that:

$$
\begin{equation*}
\left|A\left(\phi_{i}=1\right)-b\right|<\frac{\tau_{2, Y}}{1-e} \tag{4.8}
\end{equation*}
$$

Suppose that (4.8) is satisfied. We now ask what is the maximal azimuthal width of static mud channel that can exist. If the mud is static, then

$$
\begin{equation*}
1=\left.\int_{0}^{\phi_{i}} \frac{\partial \Psi}{\partial \phi}\right|_{k=1} \mathrm{~d} \phi \tag{4.9}
\end{equation*}
$$

i.e. all the unit flow rate must flow through fluid $1, \phi \in\left(0, \phi_{i}\right]$. Evidently as $\phi_{i} \rightarrow 0$, if the mud remains static, then $A\left(\phi_{i}\right) \rightarrow \infty$ since the unit flow rate is forced through an increasingly narrow azimuthal gap. It follows that for some $\phi_{i}$, there will be a solution to

$$
\begin{equation*}
\left|A\left(\phi_{i}\right)-b\right|=\frac{\tau_{2, Y}}{H\left(\phi_{i}\right)} \tag{4.10}
\end{equation*}
$$

which defines the smallest interface position for which the mud layer remains fully static. We denote this interface position by $\phi_{i}=\phi_{i, \min }$, and the maximal azimuthal width of mud channel is therefore $1-\phi_{i, \min }$. Let us now explore the parametric dependency of (4.8) and (4.10).

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First, suppose (4.8) holds, so that there is a static layer. For $\phi_{i}>\phi_{i, \min }$, the pressure gradient $A\left(\phi_{i}\right)$ is determined implicitly by (4.9). Evidently, $\operatorname{sgn}(A)=1$ and for $A H>\tau_{1, Y}$ we can rewrite (4.6) as

$$
\begin{equation*}
\left.\frac{\partial \Psi}{\partial \phi}\right|_{k=1}=\left(\frac{\tau_{1, Y}}{\kappa_{1}}\right)^{m_{1}} \frac{\left(\frac{A H}{\tau_{1, Y}}-1\right)^{m_{1}+1}\left(\frac{A H}{\tau_{1, Y}}+\frac{1}{m_{1}+1}\right)}{\left(m_{1}+2\right)\left(\frac{A H}{\tau_{1, Y}}\right)^{2}} \tag{4.11}
\end{equation*}
$$

which at each $\phi$ depends only on $e, m_{1}, A / \tau_{1, Y}$ and $B_{1}=\tau_{1, Y} / \kappa_{1}$. Therefore, we see that (4.9) implies

$$
\frac{A\left(\phi_{i}\right)}{\tau_{1, Y}}=f\left(e, B_{1}, m_{1}\right)
$$

Consequently, dividing (4.10) by $\tau_{1, Y}$, we have

$$
\begin{equation*}
H\left(\phi_{i, \min }\right)\left|f\left(e, B_{1}, m_{1}\right)-\frac{b}{\tau_{1, Y}}\right|=\frac{\tau_{2, Y}}{\tau_{1, Y}} \tag{4.12}
\end{equation*}
$$

i.e. $\phi_{i, \text { min }}$ depends on $\left(e, B_{1}, m_{1}, \Lambda_{b}, \Lambda_{Y}\right)$, where

$$
\begin{equation*}
\Lambda_{b}=\frac{b}{\tau_{1, Y}} \quad \Lambda_{Y}=\frac{\tau_{2, Y}}{\tau_{1, Y}} \tag{4.13}
\end{equation*}
$$

are the buoyancy ratio and yield stress ratio, respectively. The condition that (4.8) holds is now:

$$
\begin{equation*}
\Lambda_{Y}>(1-e)\left|f\left(e, B_{1}, m_{1}\right)-\frac{b}{\tau_{1, Y}}\right| \tag{4.14}
\end{equation*}
$$

i.e. we have static mud channels if and only if the yield stress ratio exceeds a critical value that is dependent on ( $e, B_{1}, m_{1}, \Lambda_{b}$ ).

Figures 4.5-4.9 show the effects of eccentricity on the magnitude of $\phi_{i, \min }$. As expected, if we increase the eccentricity of the annulus a mud channel will be left during the displacement in the narrow side of it. Rheological parameters play an important role as well, clearly if $\tau_{Y ; 2}>\tau_{Y, 1}$, i.e. $\Lambda_{Y}>1$, the pressure gradient necessary to overcome the yield stress of both fluids at the interface will not suffice to overcome the yield stress of fluid 2 in the narrow side of the annulus. Having as a result a decrease of the domain where $\phi_{i, \min }=1$. The same phenomena will occur if either $\kappa_{2}>\kappa_{1}$ and $\rho_{2}>\rho_{1}$.


Figure 4.5: Contour plot of $\phi_{i, \min }$, eccentricity vs. yield stress of fluid one, buoyancy parameter equal zero. Rhelogical and physical parameters are: $\kappa_{1}=0.5, \rho_{1}=1, m_{1}=1, \tau_{Y, 2}=0.5$, $\kappa_{2}=0.4, \rho_{2}=1, m_{2}=1$ and $\beta=0$. The zero level curve corresponds to the equality of condition (4.14), i.e. $\phi_{i, \min }=1$, the rest of the level curves correspond to different values of $\phi_{i, \min }$ when condition (4.14) is fulfilled.


Figure 4.6: Contour plot of $\phi_{i, \min }$, eccentricity vs. yield stress of fluid one, buoyancy parameter negative. Rhelogical and physical parameters are: $\kappa_{1}=0.5, \rho_{1}=1.1, m_{1}=1, \tau_{Y, 2}=0.5$, $\kappa_{2}=0.4, \rho_{2}=1, m_{2}=1$ and $\beta=0$. The zero level curve corresponds to the equality of condition (4.14), i.e. $\phi_{i, \min }=1$, the rest of the level curves correspond to different values of $\phi_{i, \min }$ when condition (4.14) is fulfilled.


Figure 4.7: Contour plot of $\phi_{i, \min }$, eccentricity vs. yield stress of fluid one, buoyancy parameter positive. Rhelogical and physical parameters are: $\kappa_{1}=0.5, \rho_{1}=1, m_{1}=1, \tau_{Y, 2}=0.5, \kappa_{2}=0.4$, $\rho_{2}=1.1, m_{2}=1$ and $\beta=0$. The zero level curve corresponds to the equality of condition (4.14), i.e. $\phi_{i, \min }=1$, the rest of the level curves correspond to different values of $\phi_{i, \min }$ when condition (4.14) is fulfilled.


Figure 4.8: Contour plot of $\phi_{i, m i n}$, eccentricity vs. yield stress of fluid one, equal consistency. Rhelogical and physical parameters are: $\kappa_{1}=0.5, \rho_{1}=1, m_{1}=1, \tau_{Y, 2}=0.5, \kappa_{2}=0.5, \rho_{2}=1$, $m_{2}=1$ and $\beta=0$. The zero level curve corresponds to the equality of condition (4.14), i.e. $\phi_{i, \min }=1$, the rest of the level curves correspond to different values of $\phi_{i, \min }$ when condition (4.14) is fulfilled.


Figure 4.9: Contour plot of $\phi_{i, \min }$, eccentricity vs. yield stress of fluid one, consistency of fluid 1 less than consistency of fluid 2. Rhelogical and physical parameters are: $\kappa_{1}=0.3, \rho_{1}=1$, $m_{1}=1, \tau_{Y, 2}=0.5, \kappa_{2}=0 . \dot{4}, \rho_{2}=1, m_{2}=1$ and $\beta=0$. The zero level curve corresponds to the equality of condition (4.14), i.e. $\phi_{i, \min }=1$, the rest of the level curves correspond to different values of $\phi_{i, \min }$ when condition (4.14) is fulfilled.

### 4.1.2 Summary of results

Solutions of a basic flow consists of solving a 1-D nonlinear equation, followd by integration of an algebraic relation to find $\Psi$. This computations depend on 9 different parameters, $\tau_{Y, k}$, $\kappa_{k}, m_{k}$, for $k=1,2, b, e$ and $\phi$. A case of industrial interest is when a static layer forms (see $\S 4.1 .1$ ). Such flows occur only for interfaces $\phi_{i}>\phi_{i, \min }$, and for certain combinations of parameters. We have shown that $\phi_{i, \min }$ depends only on 6 parameters, $\tau_{Y, k}$, for $k=1,2, \kappa_{1}$, $m_{1}, b$ and $e$. Therefore, provided that $\phi_{i}>\phi_{i, \min }$ the flows depend only on a reduced subset of the dimensionless parameters.

### 4.2 Stability of parallel multi-layer flows

For the rest of this chapter we will work with the interface tracking model, defined in $\S 2.5$. We consider the stability of any of the parallel multi-layer flows in $\S 4.1$. We denote the transient
interface by, $F=\phi-\phi_{i}(\xi, t)=0$. Our system of equations can be written as:

$$
\begin{align*}
\nabla \cdot\left[\mathbf{S}_{k}+\rho_{k} \nabla \Psi_{k, t}\right] & =0, \quad \text { in } \Omega_{k},  \tag{4.15}\\
\frac{H}{\epsilon} \frac{\partial \phi_{i}}{\partial t}+\frac{\partial \Psi_{k}}{\partial \xi}+\frac{\partial \Psi_{k}}{\partial \phi} \frac{\partial \phi_{i}}{\partial \xi} & =0, \quad \text { at } \phi=\phi_{i}(\xi, t)  \tag{4.16}\\
\Psi_{1}(0, \xi, t) & =0,  \tag{4.17}\\
\Psi_{2}(1, \xi, t) & =1,  \tag{4.18}\\
\left.\Psi_{k}\right|_{k=1} ^{k=2} & =0, \quad \text { at } \phi=\phi_{i}(\xi, t),  \tag{4.19}\\
\left.\left(\mathbf{S}_{k}+\rho_{k} \nabla \Psi_{k, t}\right) \cdot \mathbf{n}\right|_{k=1} ^{k=2} & =-b\left(1, \tan \beta \sin \pi \phi_{i}\right) \cdot \mathbf{n}, \quad \text { at } \phi=\phi_{i}(\xi, t), \tag{4.20}
\end{align*}
$$

We consider linear perturbation of these flows, assume normal mode expansions and transform to an eigenvalue problem, in the usual way. We denote the base flow solution from $\S 4.1$, by $\Psi_{k}=\Psi_{k, 0}(\phi) \& \mathbf{S}_{k}=\left(S_{k, \phi, 0}, 0\right)$, with interface position $\phi_{i}=\phi_{i, 0}$, and recall that $S_{k, \phi, 0}$ is constant in each domain. For $\delta \ll 1$ we assume the expansions:

$$
\begin{aligned}
\Psi_{k} & =\Psi_{k, 0}+\delta \Psi_{k, 1}+\delta^{2} \Psi_{k, 2}+\ldots \\
\mathbf{S}_{k} & =\left(S_{k, \phi, 0}, 0\right)+\delta \mathbf{S}_{k, 1}+\delta^{2} \mathbf{S}_{k, 2}+\ldots \\
\phi_{i} & =\phi_{i, 0}+\delta h+\ldots
\end{aligned}
$$

substitute into (4.15-4.20) and expand in powers of $\delta$.

$$
\begin{align*}
\nabla \Psi_{k} & =\left(\Psi_{k, 0, \phi}+\delta \Psi_{k, 1, \phi}+\ldots, \delta \Psi_{k, 1, \xi}+\ldots\right)  \tag{4.21}\\
|\nabla \Psi| & =\left[\left(\Psi_{k, 0, \phi}+\delta \Psi_{k, 1, \phi}+\ldots\right)^{2}+\left(\delta \Psi_{k, 1, \xi}+\ldots\right)^{2}\right]^{\frac{1}{2}} \\
& \sim\left|\Psi_{k, 0, \phi}\right|+\delta \operatorname{sgn}\left(\Psi_{k, 0, \phi}\right) \Psi_{k, 1, \phi}+O\left(\delta^{2}\right) \tag{4.22}
\end{align*}
$$

When the fluid is yielded we find:

$$
\begin{equation*}
\mathbf{S}_{k, 1}=\left[\chi_{k}^{\prime}\left(\left|\Psi_{k, 0, \phi}\right|\right) \Psi_{k, 1, \phi},\left(\chi_{k}\left(\left|\Psi_{k, 0, \phi}\right|\right)+\frac{\tau_{k, Y}}{H}\right) \frac{\Psi_{k, 1, \xi}}{\left|\Psi_{k, 0, \phi}\right|}\right] \tag{4.23}
\end{equation*}
$$

but otherwise $\mathbf{S}_{k}$ is indeterminate.

We see that the stability problem depends on 14 dimensionless parameters. First of all the base flows depend on the following 9 dimensionless parameters: $\tau_{Y, k}, \kappa_{k}, m_{k}$, for $k=1,2, b, e$ and
$\phi_{i}$. Now the buoyancy parameter $b$ is defined in terms of the two densities, $\rho_{k}$, the inclination $\beta$ and the Stokes number, St. The densities and the inclination appear individually in the above stability problem and therefore the full stability problem depends on 12 dimensionless parameters. Additionally, rather than solving the initial value problem above, we will take a normal mode expansion of the linear perturbation. This will introduce a 13 th dimensionless parameter, the wavenumber in the axial direction. Finally, in the kinematic equation for the interface motion appears a 14 th parameter $\epsilon$. Since this general situation is difficult to analyse, we start with some of the simplified scenarios. First we consider the stability of a single fluid along the annulus.

### 4.3 Single fluid problem

We commence with the flow of a single fluid along the annulus, and assume that the base flow of the single fluid is everywhere yielded, although the arguments following will also apply if the base flow is unyielded in a static layer on the narrow side of the annulus. The parameter dependence reduces greatly for this problem, now we only have 5 parameters to take into account: $\tau_{Y}, \kappa, m, e$ and the wavenumber $\alpha$. The linear stability equation for the single fluid is:

$$
\begin{equation*}
-\rho\left(\Delta \Psi_{1}\right)_{t}=\nabla \cdot\left[\chi^{\prime}\left(\left|\Psi_{0, \phi}\right|\right) \Psi_{1, \phi}, \frac{\chi\left(\left|\Psi_{0, \phi}\right|\right)+\frac{\tau_{Y}}{H}}{\left|\Psi_{0, \phi}\right|} \Psi_{1, \xi}\right] \tag{4.24}
\end{equation*}
$$

with boundary conditions:

$$
\Psi_{1}(0)=0=\Psi_{1}(1)
$$

We suppose a normal mode solution for $\Psi_{1}$ :

$$
\Psi_{1} \sim f(\phi) e^{i(\alpha \xi-s t)}
$$

Substituting $\Psi_{1}$ into (4.24) we have:

$$
\begin{equation*}
i s \rho\left(D^{2}-\alpha^{2}\right) f=D\left[D f \chi^{\prime}\left(\left|\Psi_{0, \phi}\right|\right)\right]-\alpha^{2} \frac{\chi\left(\left|\Psi_{0, \phi}\right|\right)+\frac{\tau_{Y}}{H}}{\left|\Psi_{0, \phi}\right|} f \tag{4.25}
\end{equation*}
$$

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where $D=\frac{d}{d \phi}$. Multiplying (4.25) by $f^{*}$, the complex conjugate of $f$, and integrating (by parts) with respect to $\phi$, we obtain:

$$
\begin{equation*}
i s\left[J_{1}+\alpha^{2} I_{1}\right]=J_{2}+\alpha^{2} I_{2}, \tag{4.26}
\end{equation*}
$$

where

$$
\begin{align*}
& I_{1}=\int_{0}^{1} \rho|f|^{2} \mathrm{~d} \phi  \tag{4.27}\\
& I_{2}=\int_{0}^{1}|f|^{2} \frac{\chi\left(\left|\Psi_{0, \phi}\right|\right)+\frac{\tau \gamma}{H}}{\left|\Psi_{0, \phi}\right|} \mathrm{d} \phi  \tag{4.28}\\
& J_{1}=\int_{0}^{1} \rho|D f|^{2} \mathrm{~d} \phi  \tag{4.29}\\
& J_{2}=\int_{0}^{1}|D f|^{2} \chi^{\prime} \mathrm{d} \phi, \tag{4.30}
\end{align*}
$$

all of which are real positive definite. If we suppose that the normal mode is temporal, i.e. $\alpha>0$, then clearly $s=$ is $s_{I}$ is imaginary and the growth rate $s_{I}<0$, so that the flow is linearly stable. Therefore, temporal instability can only occur for the multi-fluid problem, i.e. due to the presence of an interface and the jump conditions found there.

In Figure 4.10 we show the spectrum of problem (4.25) with rheological parameters, $\tau_{Y, 1}=1$, $\kappa_{1}=1, m_{1}=1, \rho_{1}=0.5$, on the left-hand-side and $\rho_{1}=1$, on the right-hand-side. We show our results in the following order, first we plot the spectrum as $\alpha \rightarrow 0$, from (4.25) we may simply derive the analytical bound for $s_{i}$, the upper and lower analytical bounds are defined by:

$$
\begin{equation*}
-\max \chi^{\prime} \leq s_{i} \leq-\min \chi^{\prime} \tag{4.31}
\end{equation*}
$$

The argument of $\chi^{\prime}$ is $\left|\Psi_{0, \phi}\right|(\phi)$ and the max and min are taken over $\phi \in[0,1]$. It has been shown that $\chi^{\prime}>0$, thus it is clear that the imaginary part of the temporal normal mode is always negative for $\alpha$ small .

Second, the pictures in the middle of Figure 4.10 show the maximum of the imaginary part of the eigenvalues for varying wave number $\alpha$.

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Figure 4.10: Spectrum of the temporal normal mode with lower and upper bounds as in (4.31) \& (4.32), figures on the left-hand-side show a fluid with rheological parameters: left-hand-side, $\tau_{Y, 1}=1, \kappa_{1}=1 ; \rho_{1}=0.5, m_{1}=1, e=0.4$ and $\beta=0$ on the right-hand-side we only change $\rho_{1}$ to 1 .

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Finally, we show the spectrum as $\alpha \rightarrow \infty$. Here from (4.25) we can derive:

$$
\begin{equation*}
-\frac{\max \left(\chi+\tau_{Y} / H\right)}{\min \left|\Psi_{0}^{\prime}\right|} \leq s_{i} \leq-\frac{\min \left(\chi+\tau_{Y} / H\right)}{\max \left|\Psi_{0}^{\prime}\right|}, \tag{4.32}
\end{equation*}
$$

as we know, $\chi+\tau_{Y} / H$ has been defined as the absolute value of the modified pressure gradient, thus a positive quantity.

### 4.4 Two-fluid eigenvalue problems

We have seen above that any single fluid parallel flow is linearly stable to spatially periodic temporal perturbations. We ask therefore, whether the presence of an interface between two parallel streams can linearly destabilise this flow. The linearised equation for $\Psi_{1}$ is

$$
\begin{equation*}
\nabla \cdot\left[\mathbf{S}_{k, 1}+\rho_{k} \nabla \Psi_{k, 1, t}\right]=0, \quad \text { in } \Omega_{k}, \tag{4.33}
\end{equation*}
$$

with $\mathbf{S}_{k, 1}$ defined by (4.23) when yielded. The kinematic equation and boundary conditions are:

$$
\begin{align*}
\frac{H}{\epsilon} h_{t}+\Psi_{k, 1, \xi}+\Psi_{k, 0, \phi} h_{\xi} & =0, \quad \text { at } \phi=\phi_{i, 0}  \tag{4.34}\\
\Psi_{1,1}(0, \xi, t) & =0  \tag{4.35}\\
\Psi_{2,1}(1, \xi, t) & =0 \tag{4.36}
\end{align*}
$$

The jump conditions (4.19) \& (4.20) are linearised, first about the basic flow and secondly onto the basic flow interface position. Thus, at $O(\delta)$, condition (4.19) becomes

$$
\begin{equation*}
\left.\left(\Psi_{k, 1}+h \Psi_{k, 0, \phi}\right)\right|_{k=1} ^{k=2}=0, \quad \text { at } \phi=\phi_{i, 0}, \tag{4.37}
\end{equation*}
$$

and we note that it is the derivative of this quantity with respect to $\xi$ that appears in (4.34), i.e. it is irrelevant which fluid is considered for the kinematic condition. Expanding $\left(\mathbf{S}_{k}+\rho_{k} \nabla \Psi_{k, t}\right) \cdot \mathbf{n}$, about $\phi=\phi_{i, 0}$ we have:

$$
\begin{aligned}
{\left[\mathbf{S}_{k}+\rho_{k} \nabla \Psi_{k, t}\right] \cdot \mathbf{n} \sim } & S_{k, 0, \phi}+\delta h \frac{\partial}{\partial \phi} S_{k, 0, \phi}+\delta S_{k, 1, \phi} \\
& +\rho_{k}\left[\Psi_{k, 0, \phi t}+\delta h \Psi_{k, 0, \phi \phi t}+\delta \Psi_{k, 1, \phi t}\right] \quad \text { at } \phi=\phi_{i, 0} \\
\left(1, \tan \beta \sin \pi \phi_{i}\right) \cdot \mathbf{n} \sim & 1-\delta h_{\xi} \tan \beta \sin \pi \phi_{i, 0}
\end{aligned}
$$

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Noting that $S_{k, 0, \phi}$ is independent of $\phi$ and $\Psi_{k, 0}$ independent of $t$, at $O(\delta)$, condition (4.20) becomes

$$
\begin{equation*}
\left.\left(S_{k, 1, \phi}+\rho_{k} \Psi_{k, 1, \phi t}\right)\right|_{k=1} ^{k=2}=b h_{\xi} \tan \beta \sin \pi \phi_{i, 0}, \quad \text { at } \phi=\phi_{i, 0} \tag{4.38}
\end{equation*}
$$

We consider first the case where both fluids are yielded and mobile throughout the annulus and subsequently those cases when a fluid layer contains static regions.

## Mobile fluids

In this case $\mathbf{S}_{k, 1}$ is defined by (4.23) and the stability problem, consisting of (4.33-4.38), is posed on the entire fluid domain. We assume a normal mode solution:

$$
\Psi_{1, k} \sim f_{k}(\phi) e^{i(\alpha \xi-s t)}, \quad h \sim h_{0} e^{i(\alpha \xi-s t)}
$$

and substitute into (4.33)-(4.38) to give:

$$
\begin{align*}
D\left[\chi_{1}^{\prime}\left(\left|\Psi_{1,0, \phi}\right|\right) D f_{1}\right]-\alpha^{2}\left[\frac{\chi_{1}\left(\left|\Psi_{1,0, \phi}\right|\right)+\frac{\tau_{1, Y}}{H}}{\left|\Psi_{1,0, \phi}\right|}\right] f_{1}, & =i s \rho_{1}\left(D^{2}-\alpha^{2}\right) f_{1} \\
\phi \in\left(0, \phi_{i, 0}\right) &  \tag{4.39}\\
D\left[\chi_{2}^{\prime}\left(\left|\Psi_{2,0, \phi}\right|\right) D f_{2}\right]-\alpha^{2}\left[\frac{\chi_{2}\left(\left|\Psi_{2,0, \phi}\right|\right)+\frac{\tau_{2, Y}}{H}}{\left|\Psi_{2,0, \phi}\right|}\right] f_{2}, & =i s \rho_{2}\left(D^{2}-\alpha^{2}\right) f_{2} \\
\phi \in\left(\phi_{i, 0}, 1\right) &  \tag{4.40}\\
f_{1}(0) & =0  \tag{4.41}\\
f_{2}(1) & =0 \tag{4.42}
\end{align*}
$$

and at $\phi=\phi_{i, 0}$.

$$
\begin{align*}
-\frac{H}{\epsilon} s h_{0}+\alpha\left(f_{k}+\Psi_{k, 0, \phi} h_{0}\right) & =0  \tag{4.43}\\
\left.\left(h_{0} \Psi_{0, k, \phi}+f_{k}\right)\right|_{1} ^{2} & =0  \tag{4.44}\\
{\left.\left[\left(\chi_{k}^{\prime}\left(\left|\Psi_{0, k, \phi}\right|\right)-i s \rho_{k}\right) D f_{k}\right]\right]_{k=1}^{k=2} } & =i \alpha b h_{0} \tan \beta \sin \pi \phi_{i, 0} \tag{4.45}
\end{align*}
$$

In general, it is necessary to solve the eigenvalue problem (4.39)-(4.45) computationally. We discuss results for this in $\S 4.6$.

We note that for temporal stability we may consider $\alpha \neq 0$, since if $\alpha=0$ then (4.43) implies that either $s=0$ (in which case $f_{k}=0$ and the perturbation is trivial), or instead $h_{0}=0$. For $h_{0}=0$, following the procedure in $\S 4.3$, we may multiply by $f_{k}^{*}$, integrate across each fluid layer, sum the resulting expressions and show that $s=\mathrm{i} s_{I}$ with $s_{I}<0$. The interfacial terms cancel by virtue of (4.44) \& (4.45), which simplify considerably.

## Static mud channel problems: $\phi_{i}>\phi_{i, m i n}$

In this case fluid 2 , on the narrow side, is completely unyielded and hence static. Therefore, for $\phi>\phi_{i}$ we have

$$
\left|\mathbf{S}_{2,0}\right|<\frac{\tau_{2, Y}}{H\left(\phi_{i}\right)} .
$$

By choosing $\delta$ sufficiently small we may ensure that

$$
\left|\mathbf{S}_{2,0}+\delta \mathbf{S}_{2,1}\right|<\frac{\tau_{2, Y}}{H\left(\phi_{i}\right)}
$$

so that that fluid 2 remains unyielded. In this case, the interface does not deform under a linear perturbation and we have again a single fluid problem:

$$
\begin{align*}
D\left[\chi_{1}^{\prime}\left(\left|\Psi_{1,0, \phi}\right|\right) D f_{1}\right]-\alpha^{2}\left[\frac{\chi_{1}\left(\left|\Psi_{1,0, \phi}\right|\right)+\frac{\tau_{1, Y}}{H}}{\left|\Psi_{1,0, \phi}\right|}\right] f_{1}, & =i s \rho_{1}\left(D^{2}-\alpha^{2}\right) f_{1} \\
\phi \in\left(0, \phi_{i, 0}\right) &  \tag{4.46}\\
f_{1}(0) & =0  \tag{4.47}\\
f_{1}\left(\phi_{i, 0}\right) & =0 \tag{4.48}
\end{align*}
$$

As in $\S 4.3$, we can multiply through by $f_{1}^{*}$, integrate over $\left(0, \phi_{i, 0}\right)$, and establish that these flows are linearly stable to temporal perturbations, but unstable to any non-trivial spatial perturbation.

## Partly static fluids

In this case we first suppose that fluid 1 is fully yielded and fluid 2 is part yielded and part unyielded. We discuss afterwards the other possibilities. Therefore, we have a yield surface at say $\phi_{2, Y}$ where

$$
\phi_{2, Y}=\phi_{2, Y, 0}+\delta \phi_{2, Y, 1}(\xi, t)
$$

with $\phi_{2, Y, 0}$ the yield surface of fluid 2 in the base multi-layer flow and $\delta \phi_{2, Y, 1}(\xi, t)$ is the linear perturbation of the yield surface. Note that $\phi_{i}<\phi_{2, Y} \leq 1$ and $\phi_{i}<\phi_{i, \min }$.

As $\phi \rightarrow \phi_{2, Y}^{+}$fluid 2 becomes unyielded across the annular gap and static, i.e.

$$
\left|\nabla \Psi_{2}\right| \rightarrow 0 \quad \text { as } \quad \phi \rightarrow \phi_{2, Y}^{+}
$$

Expanding each component to $O(\delta)$ and linearising onto $\phi_{2, Y, 0}$, we have that:

$$
\begin{array}{rlrl}
\Psi_{2,1, \phi}+\phi_{2, Y, 1}(\xi, t) \Psi_{2,0, \phi \phi} & =0, & & \text { at } \phi=\phi_{2, Y, 0} \\
\Psi_{2,1, \xi} & =0, & \text { at } \phi=\phi_{2, Y, 0} \tag{4.50}
\end{array}
$$

Equation (4.50) leads to the boundary condition, $f_{2}=0$ at $\phi=\phi_{2, Y, 0}$, for the normal mode. This condition is the same as (4.42). Equation (4.49) simply describes evolution of the yield surface once the stream function perturbation is known. As in other linear stability analyses of yield stress fluids, [27, 28], the yield surface perturbation plays an essentially passive role, in that the same boundary conditions are satisfied with or without the yield surface. On the other hand, its worth noting that the yield stress affects the stability problem; via the basic flow and (here) via definition of the yielded region.

Our eigenvalue problem is identical to (4.39)-(4.45), but replacing (4.40) \& (4.42) with :

$$
\begin{align*}
D\left[\chi_{2}^{\prime}\left(\left|\Psi_{2,0, \phi}\right|\right) D f_{2}\right]-\alpha^{2}\left[\frac{\chi_{2}\left(\left|\Psi_{2,0, \phi}\right|\right)+\frac{\tau_{2, Y}}{H}}{\left|\Psi_{2,0, \phi}\right|}\right] f_{2}, & =i s \rho_{2}\left(D^{2}-\alpha^{2}\right) f_{2} \\
\phi \in\left(\phi_{i, 0}, \phi_{2, Y, 0}\right) &  \tag{4.51}\\
: f_{2}\left(\phi_{2, Y, 0}\right) & =0 \tag{4.52}
\end{align*}
$$

i.e. only the domain of the fluid 2 problem is changed.

The other situation to consider is that in which fluid 1 is partly unyielded. Since $H(\phi)$ decreases, the criterion $\left|\mathbf{S}_{1}\right| \leq \tau_{1, Y} / H$ will be met in some layer $\phi \in\left(\phi_{1, Y, 0}, \phi_{i}\right)$ for the basic flow. It follows that $\left|\mathbf{S}_{1}\right|<\tau_{1, Y} / H$ at the interface, for the base flow, and hence that for sufficiently small $\delta$
the interface does not yield under the perturbation. The stability problem in fluid 1 becomes:

$$
\begin{equation*}
D\left[\chi_{1}^{\prime}\left(\left|\Psi_{1,0, \phi}\right|\right) D f_{1}\right]-\alpha^{2}\left[\frac{\chi_{1}\left(\left|\Psi_{1,0, \phi}\right|\right)+\frac{\tau_{1, Y}}{H}}{\left|\Psi_{1,0, \phi}\right|}\right] f_{1},=i s \rho_{1}\left(D^{2}-\alpha^{2}\right) f_{1} \tag{4.53}
\end{equation*}
$$

with boundary conditions (4.41) and

$$
\begin{equation*}
f_{1}\left(\phi_{1, Y, 0}\right)=0 \tag{4.54}
\end{equation*}
$$

For fluid 2, we have either (4.40) or (4.51), depending on whether the base flow of fluid 2 is fully or partially yielded. In either case, $f_{2}\left(\phi_{i}\right)=0$, since the interface is not deformed, and either (4.42) or (4.52) is satisfied at $\phi=1$ or $\phi=\phi_{2, Y, 0}$, respectively.

Neither of these cases is of any interest. For both, we may follow $\S 4.3$, multiply the fluid $k$ equation by $f_{k}^{*}$, integrate across the domains and establish that the flows are linearly stable.

## Summary

To summarise, the only base flows that may be temporally unstable are: (i) two fully yielded fluid layers; (ii) fluid 1 fully yielded and fluid 2 partly yielded (i.e. a static channel on the narrow side, wholly contained within fluid 2). Stability of the other base flows essentially reduces to the stability of one or more independent single fluid layers.

### 4.5 Analytical simplifications

Analytical progress is possible only in limited cases. One of these cases is when the annulus is concentric and consequently the base velocity is constant in each fluid layer. The simplest case is when there is no yield stress.

### 4.5.1 Newtonian fluids in a concentric annulus

For two Newtonian fluids in a concentric annulus $(H(\phi)=1)$, we find:

$$
\begin{equation*}
\Psi_{0,1, \phi}=\frac{1+\left(1-\phi_{i, 0}\right) \frac{b}{3 \kappa_{2}}}{\phi_{i, 0}+\left(1-\phi_{i, 0}\right) \frac{\kappa_{1}}{\kappa_{2}}}, \quad \Psi_{0,2, \phi}=\frac{\frac{\kappa_{1}}{\kappa_{2}}-\phi_{i, 0} \frac{b}{3 \kappa_{2}}}{\phi_{i, 0}+\left(1-\phi_{i, 0}\right) \frac{\kappa_{1}}{\kappa_{2}}} \tag{4.55}
\end{equation*}
$$

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The two fluid layers are mobile and in fluid $k, \chi_{k}$ and $|\nabla \Psi|$ are related linearly by: $\chi_{k}=$ $3 \kappa_{k}|\nabla \Psi|$. Therefore, (4.39) \& (4.40) become

$$
\begin{array}{ll}
\left(3 \kappa_{1}-i \rho_{1} s\right)\left(D^{2}-\alpha^{2}\right) f_{1}=0, & \phi \in\left(0, \phi_{i, 0}\right) \\
\left(3 \kappa_{2}-i \rho_{2} s\right)\left(D^{2}-\alpha^{2}\right) f_{2}=0, & \phi \in\left(\phi_{i, 0}, 1\right)
\end{array}
$$

We note that if $\left(3 \kappa_{k}-i \rho_{k} s\right)=0$, then $s_{I}<0$ and these modes are stable. Therefore, for instability we may assume that $\left(3 \kappa_{k}-i \rho_{k} s\right) \neq 0$, and interestingly observe that the eigenvalue $s$ does not appear in the field equations, but only the jump conditions. This degeneracy is peculiar to the Newtonian case. The solutions that satisfy (4.41) \& (4.42) are:

$$
\begin{equation*}
f_{1}(\phi)=A_{1} \sinh \alpha \phi, \quad f_{2}(\phi)=A_{2} \sinh \alpha(\phi-1), \tag{4.56}
\end{equation*}
$$

for constants $A_{1} \& A_{2}$. We may assume that $h_{0} \neq 0$, write $A_{1}=h_{0} C_{1}, A_{2}=h_{0} C_{2}$ for constants $C_{1} \& C_{2}$, and divide (4.43)-(4.45) by $h_{0}$. From (4.43) \& (4.44) we find

$$
\begin{aligned}
C_{1} & =\frac{1}{\sinh \alpha \phi_{i, 0}}\left[\frac{s}{\alpha \epsilon}-\frac{1+\frac{b}{3 \kappa_{2}}\left(1-\phi_{i, 0}\right)}{\phi_{i, 0}+\left(1-\phi_{i, 0}\right) \frac{\kappa_{1}}{\kappa_{2}}}\right] \\
C_{2} & =\frac{1}{\sinh \alpha\left(\phi_{i, 0}-1\right)}\left[\frac{s}{\alpha \epsilon}-\frac{\frac{\kappa_{1}}{\kappa_{2}}+\frac{b}{3 \kappa_{2}} \phi_{i, 0}}{\phi_{i, 0}+\left(1-\phi_{i, 0}\right) \frac{\kappa_{1}}{\kappa_{2}}}\right]
\end{aligned}
$$

Finally, substituting into (4.45), we find the following quadratic relation for $s$ :

$$
\begin{align*}
0= & i \frac{b}{3 \kappa_{2}} \tan \beta \sin \pi \phi_{i, 0}  \tag{4.57}\\
& +\operatorname{coth} \alpha \phi_{i, 0}\left[\frac{s}{\alpha \epsilon}-\frac{1+\frac{b}{3 \kappa_{2}}\left(1-\phi_{i, 0}\right)}{\phi_{i, 0}+\left(1-\phi_{i, 0}\right) \frac{\kappa_{1}}{\kappa_{2}}}\right]\left[\frac{\kappa_{1}}{\kappa_{2}}-i s \frac{\rho_{1}}{3 \kappa_{2}}\right] \\
& -\operatorname{coth} \alpha\left(\phi_{i, 0}-1\right)\left[\frac{s}{\alpha \epsilon}-\frac{\frac{\kappa_{1}}{\kappa_{2}}-\frac{b}{3 \kappa_{2}} \phi_{i, 0}}{\phi_{i, 0}+\left(1-\phi_{i, 0}\right) \frac{\kappa_{1}}{\kappa_{2}}}\right]\left[1-i s \frac{\rho_{2}}{3 \kappa_{2}}\right]
\end{align*}
$$

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Figure 4.11: Maximum imaginary part of $s$ vs. $\alpha$. Rheological and physical parameters are: $\kappa_{1}=1, e=0.0, \phi_{i}=0.5$ and $\beta=0$. a) $\rho_{1}=1, \rho_{2}=0.75$. b) $\rho_{1}=1, \rho_{2}=1.25$.


Figure 4.12: Maximum imaginary part of $s$ vs. $\alpha$. Rheological and physical parameters are: $\kappa_{1}=1, \rho_{1}=1, \rho_{2}=1 e=0.0, \phi_{i}=0.5$ and $\beta=0$.

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Figure 4.13: Maximum imaginary part of $s$ vs. $\alpha$. Rheological and physical parameters are: $\kappa_{1}=1, e=0.0, \phi_{i}=0.45$ and $\beta=0$. a) $\rho_{1}=1, \rho_{2}=0.75$. b) $\rho_{1}=1, \rho_{2}=1.25$.


Figure 4.14: Maximum imaginary part of $s$ vs. $\alpha$. Rheological and physical parameters are: $\kappa_{1}=1, e=0.0, \phi_{i}=0.55$ and $\beta=0$. a) $\rho_{1}=1, \rho_{2}=0.75$. b) $\rho_{1}=1, \rho_{2}=1.25$.


Figure 4.15: Maximum imaginary part of $s$ vs. $\alpha$. Rheological and physical parameters are: $\kappa_{1}=1, e=0.0, \phi_{i}=0.45$ and $\beta=\pi / 8$. a) $\rho_{1}=1, \rho_{2}=0.75$. b) $\rho_{1}=1, \rho_{2}=1.25$.

Figures 4.11 and 4.12 show the maximum imaginary part of $s$ as a function of the wave length $\alpha$. As we can see, buoyancy is a key parameter for instability to arise. In the absence of any density difference, it appears that viscosity differences are not strong enough to drive an unstable behaviour. Thus, it seems that we need the velocity difference at the interface to be sufficiently large for instabilities to appear. Note the difference between the results found by Raghavan \& Marsden, [62], and ours. They showed that for a horizontal multi-layer parallel flow in a porous media in the presence of density difference, if the heavier fluid was in top of the lighter one, the flow was unstable for all viscosity ratios, if the case is reversed, the lighter on top of the heavier, they claim that the flow will be unstable only for large viscosity ratios (the more viscous on top of the less viscous). For the vertical parallel flow, buoyant forces play a key role in defining the jump in the tangential velocity.

Figures 4.13 and 4.14 show the effect of the interface position on the stability of the flow. Clearly, effects of this position are not of great importance, for a concentric flow.

Figure 4.15 shows the effect of inclination of the annulus. As expected, in the presence of
buoyancy, the flow may be unstable if the heavier fluid is on top of the lighter one, i.e. $\rho_{1}>\rho_{2}$. If we have positive buoyancy, i.e. $\rho_{1}<\rho_{2}$, then the inclination has stabilizing effects.

Another case for analytical progress is when there is no yield stress but the power law index $n$ is less than one. Again we will assume that the annulus is concentric and that the base velocity is constant in each fluid layer.

### 4.5.2 Power-Law fluids in a concentric annulus

Because the base velocity is constant in each fluid layer, we find:

$$
\Psi_{0,1, \phi}=\operatorname{sgn}(A) \frac{|A|^{m_{1}}}{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}, \quad \Psi_{0,2, \phi}=\operatorname{sgn}(A-b) \frac{|A-b|^{m_{2}}}{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)}
$$

where $A$ is the modified pressure gradient defined in (2.57) and $b$ is the buoyancy. We can find $A\left(\phi_{i}\right)$ from:

$$
\begin{equation*}
1=\phi_{i} \operatorname{sgn}(A) \frac{|A|^{m_{1}}}{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}+\left(1-\phi_{i}\right) \operatorname{sgn}(A-b) \frac{|A-b|^{m_{2}}}{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)} \tag{4.58}
\end{equation*}
$$

In general (4.58) has to be solved numerically. The two fluid layers are mobile and in fluid $k$, from the definition of $\chi_{k}$, and having no yield stress we see that:

$$
\begin{equation*}
\chi_{1}\left(\left|\Psi_{0, \phi, 1}\right|\right)=|A| \quad \chi_{2}\left(\left|\Psi_{0, \phi, 2}\right|\right)=|A-b| \tag{4.59}
\end{equation*}
$$

Using (4.59) we can find an explicit expression for $\chi_{k}^{\prime}$ in terms of $A$,

$$
\begin{equation*}
\chi_{1}^{\prime}\left(\left|\Psi_{0, \phi, 1}\right|\right)=\frac{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}{m_{1}|A|^{m_{1}-1}} \quad \chi_{2}^{\prime}\left(\left|\Psi_{0, \phi, 2}\right|\right)=\frac{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)}{m_{2}|A-b|^{m_{2}-1}}, \tag{4.60}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{\chi_{1}\left(\left|\Psi_{0, \phi, 1}\right|\right)}{\left|\Psi_{0, \phi, 1}\right|}=\frac{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}{|A|^{m_{1}-1}} \quad \frac{\chi_{2}\left(\left|\Psi_{0, \phi, 2}\right|\right)}{\left|\Psi_{0, \phi, 2}\right|}=\frac{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)}{|A-b|^{m_{2}-1}} \tag{4.61}
\end{equation*}
$$

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Substituting (4.60) \& (4.61), into (4.39) \& (4.40), we get

$$
\begin{array}{rlr}
D^{2} f_{1}-\alpha^{2}\left(\frac{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)-i|A|^{m_{1}-1} \rho_{1} s}{\frac{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}{m_{1}}-i|A|^{m_{1}-1} \rho_{1} s}\right) & =0, & \phi \in\left(0, \phi_{i, 0}\right) \\
D^{2} f_{2}-\alpha^{2}\left(\frac{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)-i|A-b|^{m_{2}-1} \rho_{2} s}{\frac{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)}{m_{2}}-i|A-b|^{m_{2}-1} \rho_{2} s}\right) & =0, & \phi \in\left(\phi_{i, 0}, 1\right) \tag{4.63}
\end{array}
$$

with boundary and jump conditions:

$$
\begin{array}{rlr}
f_{1}(0) & =0 \\
f_{2}(1) & =0 \\
\frac{f_{2}-f_{1}}{h_{0}} & =\operatorname{sgn}(A) \frac{|A|^{m_{1}}}{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}-\operatorname{sgn}(A-b) \frac{|A-b|^{m_{2}}}{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)} \\
-i \alpha b \tan \beta \sin \pi \phi_{i, 0} & =\left(\frac{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}{m_{1}|A|^{m_{1}-1}}-\rho_{1} s\right) D f_{1}-\left(\frac{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)}{m_{2}|A|^{m_{2}-1}}-\rho_{2} s\right) D f_{2} & \text { at } \phi=\phi_{i, 0} \tag{4.66}
\end{array}
$$

and the kinematic condition

$$
\begin{equation*}
-\frac{s h_{0}}{\epsilon}+\alpha\left(f_{1}+h_{0} \operatorname{sgn}(A) \frac{|A|^{m_{1}}}{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}\right)=0 \quad \text { at } \phi=\phi_{i, 0} \tag{4.68}
\end{equation*}
$$

Solving (4.62) \& (4.63) with (4.64) \& (4.65) respectively we have:

$$
\begin{equation*}
f_{1}(\phi)=C_{1} h_{0} \sinh \lambda_{1} \phi, \quad f_{2}(\phi)=C_{2} h_{0} \sinh \lambda_{2}(\phi-1) \tag{4.69}
\end{equation*}
$$

for constants $C_{1} \& C_{2}$ and

$$
\begin{aligned}
& \lambda_{1}^{2}=\alpha^{2}\left(\frac{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)-i|A|^{m_{1}-1} \rho_{1} s}{\frac{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}{m_{1}}-i|A|^{m_{1}-1} \rho_{1} s}\right) \\
& \lambda_{2}^{2}=\alpha^{2}\left(\frac{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)-i|A-b|^{m_{2}-1} \rho_{2} s}{\frac{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)}{m_{2}}-i|A-b|^{m_{2}-1} \rho_{2} s}\right)
\end{aligned}
$$

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Assuming that $h_{0} \neq 0$, and using (4.68) we can find an expression for $C_{1}$,

$$
C_{1}=\frac{1}{\sinh \lambda_{1} \phi_{i}}\left(\frac{s}{\alpha \epsilon}-\frac{\operatorname{sgn}(A)|A|^{m_{1}}}{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}\right) .
$$

In a similar way, using the first jump condition, (4.66), and the definition of $C_{1}$, we can find an expression for $C_{2}$,

$$
C_{2}=\frac{1}{\sinh \lambda_{2} \phi_{i}}\left(\frac{s}{\alpha \epsilon}-\frac{\operatorname{sgn}(A-b)|A-b|^{m_{2}}}{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)}\right) .
$$

Using the definitions of $C_{1} \& C_{2}$ and (4.67) we find a non-linear equation for $s$,

$$
\begin{align*}
i \alpha b h_{0} \tan \beta \sin \pi \phi_{i, 0} & +\left(\frac{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}{m_{1}|A|^{m_{1}-1}}-\rho_{1} s\right) \frac{\lambda_{1}}{\tanh \lambda_{1} \phi_{i}}\left(\frac{s}{\alpha \epsilon}-\frac{\operatorname{sgn}(A)|A|^{m_{1}}}{\kappa_{1}^{m_{1}}\left(m_{1}+2\right)}\right) \\
& -\left(\frac{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)}{m_{2}|A|^{m_{2}-1}}-\rho_{2} s\right) \frac{\lambda_{2}}{\tanh \lambda_{2} \phi_{i}}\left(\frac{s}{\alpha \epsilon}-\frac{\operatorname{sgn}(A-b)|A-b|^{m_{2}}}{\kappa_{2}^{m_{2}}\left(m_{2}+2\right)}\right)=0 \tag{4.70}
\end{align*}
$$


(a)

(b)

Figure 4.16: Maximum imaginary part of $s$ vs. $\alpha$. Rheological and physical parameters are: $\kappa_{1}=1.5, \kappa_{2}=1, e=0.0$ and $\beta=0, m_{1}=1.5, m_{2}=2$. a) $\phi_{i}=0.5$. b) $\phi_{i}=0.6$. For $b<0$, $\rho_{1}=1$ and $\rho_{2}=0.75$, for $b=0, \rho_{1}=1$ and $\rho_{2}=1$ and for $b>0, \rho_{1}=1$ and $\rho_{2}=1.25$.


Figure 4.17: Maximum imaginary part of $s$ vs. $\alpha$. Rheological and physical parameters are: $\kappa_{1}=1.5, \kappa_{2}=1, e=0.0, \phi_{i}=0.5$ and $\beta=\pi / 8, m_{1}=1.5, m_{2}=2$. For $b<0, \rho_{1}=1$ and $\rho_{2}=0.75$, for $b=0, \rho_{1}=1$ and $\rho_{2}=1$ and for $b>0, \rho_{1}=1$ and $\rho_{2}=1.25$.

Figure 4.16 shows the maximum imaginary part of the spectrum as a function of the wave number $\alpha$. We have considered two interface positions, as in the Newtonian case, $\phi_{i}$ does not play an important role in the stability of the flow. For Newtonian fluids, we observed that if fluid 1 was more viscous than fluid 2, but fluid 2 heavier than fluid 1, i.e. buoyancy effects dominating viscous effects, the flow was unstable. We have the same phenomena for power law fluids and as for the Newtonian case, it seems that in the absence of density difference the viscous effects are not strong enough to drive an unstable behaviour.

Finally, Figure 4.17 shows the destabilizing effects of inclination if heavier fluid 1 is on top of lighter fluid 2. Even if fluid 1 is more viscous and with lower power law index than fluid 2 , stable cases from previous figures, instability arises. In this type of flow, buoyancy and inclination are the key factors for unstable behaviour.

Now we can proceed with the investigation of the full problem, i.e. yield stress fluids in an inclined eccentric annulus.

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### 4.6 Numerical Results

In this section we present numerical results for the eigenvalue problem,

$$
\begin{align*}
D\left[\chi_{1}^{\prime}\left(\left|\Psi_{1,0, \phi}\right|\right) D f_{1}\right]-\alpha^{2}\left[\frac{\chi_{1}\left(\left|\Psi_{1,0, \phi}\right|\right)+\frac{\tau_{1, Y}}{H}}{\left|\Psi_{1,0, \phi}\right|}\right] f_{1}, & =i s \rho_{1}\left(D^{2}-\alpha^{2}\right) f_{1}  \tag{4.71}\\
\phi \in\left(0, \phi_{i, 0}\right) & \\
D\left[\chi_{2}^{\prime}\left(\left|\Psi_{2,0, \phi}\right|\right) D f_{2}\right]-\alpha^{2}\left[\frac{\chi_{2}\left(\left|\Psi_{2,0, \phi}\right|\right)+\frac{\tau_{2, Y}}{H}}{\left|\Psi_{2,0, \phi}\right|}\right] f_{2}, & =i s \rho_{2}\left(D^{2}-\alpha^{2}\right) f_{2}  \tag{4.72}\\
\phi \in\left(\phi_{i, 0}, 1\right) &  \tag{4.73}\\
f_{1}(0) & =0,  \tag{4.74}\\
f_{2}(1) & =0,
\end{align*}
$$

and at $\phi=\phi_{i, 0}$ :

$$
\begin{align*}
-\frac{H}{\epsilon} s h_{0}+\alpha\left(f_{k}+\Psi_{k, 0, \phi} h_{0}\right) & =0  \tag{4.75}\\
\left.\left(h_{0} \Psi_{0, k, \phi}+f_{k}\right)\right|_{1} ^{2} & =0  \tag{4.76}\\
{\left.\left[\left(\chi_{k}^{\prime}\left(\left|\Psi_{0, k, \phi}\right|\right)-i s \rho_{k}\right) D f_{k}\right]\right|_{k=1} ^{k=2} } & =i \alpha b h_{0} \tan \beta \sin \pi \phi_{i, 0} \tag{4.77}
\end{align*}
$$

### 4.6.1 Numerical Method

In this section we give an outline of the method used to solve system (4.15)-(4.20). We use a spectral method, which is a global method that uses the fully discretized stability operator that later on can be supplied to a matrix eigenvalue solver which gives us the spectrum. We have chosen to use Chebyshev expansions to discretize our problem. The use of Chebyshev polynomials, especially in bounded domains, have been very effective and accurate for the discretization of initial and boundary value problems, see [66].

## Chevyshev polynomials

Chebyshev polynomials may be defined in several ways, e.g.

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- In terms of trigonometric functions:

$$
T_{n}(y)=\cos (n \arccos (y)) .
$$

- As solutions of the following Sturm-Liouville problem,

$$
\frac{d}{d y}\left(\sqrt{1-y^{2}} \frac{d}{d y} T_{n}(y)\right)+\frac{\dot{n}^{2}}{\sqrt{1-y^{2}}} T_{n}(y)=0 .
$$

- As a recurrence relation,

$$
T_{0}(y)=1, \quad T_{1}(y)=y, \quad \ldots \quad, T_{n+1}(y)=2 y T_{n}(y)-T_{n-1}(y) .
$$

For numerical reasons, we will use the trigonometric representation which is the most practical.
The method is summarized as follows:

- Approximate $f_{1}(y)$ and $f_{2}(y)$ using a Chebyshev expansion,

$$
f_{k}(y)=\sum_{n=0}^{N} a_{n} T_{n}(y),
$$

and evaluate the Chebyshev polynomials at the extrema of the $N$-th Chebyshev polynomial, given as

$$
y_{j}=\cos \left(\frac{j \pi}{N}\right)
$$

which are known as the Gauss-Lobatto points. Note that $y \in[-1,1]$ and that $\phi \in[0,1]$ so we use the following change of variables so we can work with the Gauss-Lobatto points,

$$
\begin{array}{ll}
\phi=\frac{-\phi_{i} y+\phi_{i}}{2} & \text { for } \phi \in\left[0, \phi_{i}\right], \\
\phi=\frac{-\left(-\phi_{i}+\phi_{Y}\right) y+\left(\phi_{i}+\phi_{Y}\right)}{2} & \text { for } \phi \in\left(\phi_{i}, 1\right],
\end{array}
$$

where $\phi_{i}$ is the position of the interface between the two fluids and $\phi_{Y}$ is the position of the yield surface, if any. Observe that if fluid 2 is completely yielded, then $\phi_{Y}=1$.

- After discretizing in this way, we have as a final result the following eigenvalue problem in discrete form,

$$
\boldsymbol{A} a=c \boldsymbol{B} a .
$$

We chose to implement the boundary, jump and kinematic conditions on the first, last and next-to last row for $f_{1}(\phi)$ matrices and on the first, second and last row for $f_{2}(\phi)$ matrices. In the case of boundary conditions, we implement them only on the first and last row of $\boldsymbol{B}$. These same rows in matrix $\boldsymbol{A}$ are chosen to be a complex multiple of the corresponding rows in $\boldsymbol{B}$. In this way, the resulting spurious modes associated with the boundary conditions can be mapped to an arbitrary location in the complex plane.

## Code validation


(a)

(b)

Figure 4.18: Maximum imaginary part of $s$ vs. $\alpha$. Inclined annulus. Rheological and physical parameters are: $\kappa_{1}=1.5, \kappa_{2}=1, m_{1}=1, m_{2}=1, b=-1$ and $\beta=0$. a) Exact solution. b) Numerical solution.

We validate our code with the exact value of the maximum imaginary part of the spectrum, for 2 Newtonian fluids, which is given by (4.57). The absolute value of the error values at each $\alpha$ is presented in Table 4.1 and the numerical and analytical values are compared in Figure (4.18). Following, we present our numerical results for the full problem.

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| $\alpha$ | $\max s_{i}$ exact | $\max s_{i}$ numerical | Error |
| ---: | ---: | ---: | ---: |
| 0.001 | $-1.327407374138356 \mathrm{E}-8$ | $-2.151373954828488 \mathrm{E}-8$ | $8.239665806901317 \mathrm{E}-9$ |
| 0.501 | $-3.284347465794034 \mathrm{E}-3$ | $-3.284348381546957 \mathrm{E}-3$ | $9.157529227109273 \mathrm{E}-10$ |
| 1.001 | $-1.256769763479928 \mathrm{E}-2$ | $-1.256769663542340 \mathrm{E}-2$ | $9.993758789267337 \mathrm{E}-10$ |
| 1.501 | $-2.638888993138258 \mathrm{E}-2$ | $-2.638889188681512 \mathrm{E}-2$ | $1.955432531175472 \mathrm{E}-9$ |
| 2.001 | $-4.280735057269110 \mathrm{E}-2$ | $-4.280735611097798 \mathrm{E}-2$ | $5.538286877715404 \mathrm{E}-9$ |
| 2.501 | $-5.993596650918038 \mathrm{E}-2$ | $-5.993596650695535 \mathrm{E}-2$ | $2.225025719226892 \mathrm{E}-12$ |
| 3.001 | $-7.634083070570413 \mathrm{E}-2$ | $-7.634083068837617 \mathrm{E}-2$ | $1.732795851250302 \mathrm{E}-11$ |
| 3.501 | $-9.117814039039307 \mathrm{E}-2$ | $-9.117814032189228 \mathrm{E}-2$ | $6.850078837494777 \mathrm{E}-11$ |
| 4.001 | $-1.041107832784735 \mathrm{E}-1$ | $-1.041107832084426 \mathrm{E}-1$ | $7.003088386969836 \mathrm{E}-11$ |
| 4.501 | $-1.151352471298047 \mathrm{E}-1$ | $-1.151352470885577 \mathrm{E}-1$ | $4.124704744423724 \mathrm{E}-11$ |
| 6 | $-1.387140792107665 \mathrm{E}-1$ | $-1.387140792299231 \mathrm{E}-1$ | $1.915664848972654 \mathrm{E}-11$ |
| 8 | $3-1.560598543973128 \mathrm{E}-1$ | $-1.560598543900940 \mathrm{E}-1$ | $7.218753372839615 \mathrm{E}-12$ |
| 10 | $-1.653734816055713 \mathrm{E}-1$ | $-1.653734815152703 \mathrm{E}-1$ | $9.030104441976050 \mathrm{E}-11$ |
| 12 | $-1.708225692269705 \mathrm{E}-1$ | $-1.708225692174500 \mathrm{E}-1$ | $9.520467747492489 \mathrm{E}-12$ |
| 37 | $-1.828331366144432 \mathrm{E}-1$ | $-1.828331366346426 \mathrm{E}-1$ | $2.019936995445448 \mathrm{E}-11$ |
| 14 | $3-1.742498623405992 \mathrm{E}-1$ | $-1.742498624152761 \mathrm{E}-1$ | $7.467687579421067 \mathrm{E}-11$ |
| 16 | $-1.765335018738152 \mathrm{E}-1$ | $-1.765335016295237 \mathrm{E}-1$ | $2.442915036926507 \mathrm{E}-10$ |
| 18 | $-1.781267104457814 \mathrm{E}-1$ | $-1.781267104777813 \mathrm{E}-1$ | $3.199993048319527 \mathrm{E}-11$ |
| 20 | $-1.792802857595362 \mathrm{E}-1$ | $-1.792802855209202 \mathrm{E}-1$ | $2.386159325684645 \mathrm{E}-10$ |
| 22 | $-1.801413773516178 \mathrm{E}-1$ | $-1.801413771880445 \mathrm{E}-1$ | $1.635732937987910 \mathrm{E}-10$ |
| 25 | $-1.810736771505876 \mathrm{E}-1$ | $-1.810736770116384 \mathrm{E}-1$ | $1.389492687575711 \mathrm{E}-10$ |
| 28 | $-1.817271639851303 \mathrm{E}-1$ | $-1.817271639827821 \mathrm{E}-1$ | $2.348121697082206 \mathrm{E}-12$ |
| 31 | $-1.822025801255159 \mathrm{E}-1$ | $-1.822025804615425 \mathrm{E}-1$ | $3.360265676821683 \mathrm{E}-10$ |
| 34 | $-1.825590619823075 \mathrm{E}-1$ | $-1.825590618802357 \mathrm{E}-1$ | $1.020717121935633 \mathrm{E}-10$ |
|  | $-1.830483481665764 \mathrm{E}-1$ | $1.170986885234981 \mathrm{E}-8$. |  |

Table 4.1: Code validation. Absolute value of the error between analytical and numerical results.

### 4.6.2 Yield stress fluids

Figure 4.19 shows the effects of yield stress and interface position on the flow in a concentric annulus. As expected, because both fluids are yielded at the interface, yield stress does not appear to have a large influence on the stability of the flow, and nor does interface position. Only viscous and buoyant forces govern the unstable behaviour. The results are very similar to those of the Newtonian fluids in Figures (4.11)-(4.14).

Figure 4.20a) shows similar behaviour as for Newtonian fluids, compare with Figure (4.15). As above, both fluids are yielded at the interface. Figure 4.20b) shows the effects of inclination of the annulus. As in the Newtonian case, if the heavier fluid1 is on top of the lighter fluid 2 the flow will be unstable.


Figure 4.19: Maximum imaginary part of $s$ vs. $\alpha$. Rheological and physical parameters are: $\kappa_{1}=1.5, \kappa_{2}=1, e=0.0$ and $\beta=0, m_{1}=1, m_{2}=1, \tau_{Y, 1}=0.9$ and $\tau_{Y, 2}=0.7$. a) $\phi_{i}=0.5$. b) $\phi_{i}=0.6$. For $b<0, \rho_{1}=1$ and $\rho_{2}=0.75$, for $b=0, \rho_{1}=1$ and $\rho_{2}=1$ and for $b>0, \rho_{1}=1$ and $\rho_{2}=1.25$.


Figure 4.20: Maximum imaginary part of $s$ vs. $\alpha$. Rheological and physical parameters are: a) $\kappa_{1}=1.5, \kappa_{2}=1, e=0.0, \phi_{i}=0.5$ and $\beta=0, m_{1}=1, m_{2}=1, \tau_{Y, 1}=0.7$ and $\tau_{Y, 2}=0.9$. b) $\kappa_{1}=1.5, \kappa_{2}=1, e=0.0, \phi_{i}=0.5, \beta=\pi / 8, m_{1}=1, m_{2}=1, \tau_{Y, 1}=0.7$ and $\tau_{Y, 2}=0.9$., $m_{1}=2, m_{2}=1.5$. For $b<0, \rho_{1}=1$ and $\rho_{2}=0.75$, for $b=0, \rho_{1}=1$ and $\rho_{2}=1$ and for $b>0$, $\rho_{1}=1$ and $\rho_{2}=1.25$.

Following this line of reasoning, if we have a concentric annulus and both fluids are yielded at the interface, changes in power law index will lead us to qualitatively similar results as for power law fluids, as examined in $\S 4.5$. Next we will study the effects of eccentricity in the stability of the flow.

Figure 4.21 shows the effects of eccentricity on the flow. As we can see, eccentricity has destabilizing effects if $b \leq 0$ and stabilizing effects if $b>0$. This may be explained as follow. We are considering flows for which both fluids are yielded at the interface. Now, when we increase the eccentricity, we need a higher pressure gradient to overcome the yield stress of fluid 2 , note that we assume that fluid 2 is on the narrow side of the annulus. If fluid 2 is less viscous but heavier than fluid 1 , the velocity difference between both fluids will decrease due to the fact that the heavier fluid is flowing in the narrow part. Having the opposite situation, the less viscous and lighter fluid in the narrow side, the velocity difference will increase and instabilities will arise.

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Figure 4.21: Maximum imaginary part of $s$ vs. $\alpha$. Different positions of the interface between the fiuids. Rheological and physical parameters are: $\kappa_{1}=1.5, \kappa_{2}=1, e=0.0$ and $\beta=0$, $m_{1}=1, m_{2}=1, \tau_{Y, 1}=0.9$ and $\tau_{Y, 2}=0.7$. a) $e=0.2$. b) $e=0.4$. c) $e=0.6$. d) $e=0.8$. For . $b<0, \rho_{1}=1$ and $\rho_{2}=0.75$, for $b=0, \rho_{1}=1$ and $\rho_{2}=1$ and for $b>0, \rho_{1}=1$ and $\rho_{2}=1.25$.


Figure 4.22: Spectrum and eigenfunctions for system (4.39)-(4.40), $\alpha=0.5$. Rheological and physical parameters are: $\kappa_{1}=1.5, \kappa_{2}=1, m_{1}=1, m_{2}=1, \tau_{Y, 1}=0.9, \tau_{Y, 2}=0.7$, negative buoyancy, $e=0.4$ and $\beta=0$. a) Spectrum of the eigenvalue problem. b) Eigenfunctions of the interfacial mode, solid line shows the real part of $f_{1}$ and $f_{2}$, dashed line shows the imaginary part of $f_{1}$ and $f_{2}$ and dashed thick line shows the magnitude of $f_{1}$ and $f_{2}$. c) Eigenfunctions of an interior mode, solid line shows the real part of $f_{1}$ and $f_{2}$, dashed line shows the imaginary part of $f_{1}$ and $f_{2}$ and dashed thick line shows the magnitude of $f_{1}$ and $f_{2}$.

Lastly in Figure 4.22a) we show the spectrum of the problem with rheological and physical parameters $\kappa_{1}=1.5, \kappa_{2}=1, m_{1}=1, m_{2}=1, \tau_{Y, 1}=0.9, \tau_{Y, 2}=0.7$, negative buoyancy, $e=0.4$ and $\beta=0$ for wave length $\alpha=0.5$. Clearly, we have only one unstable mode, the interfacial one. We proved in $\S 4.3$ that the problem of a single fluid is always stable, thus if we are away from the interface all the modes are negative. In Figures 4.22 b) and c) we show the eigenfunctions for the interfacial mode and an interior mode respectively, note the big jump at the interface for the interfacial eigenfunction this may be due to the large difference in the base velocities when instability is present.

Another interesting point that it is worth to mention is that all significant changes occur for long wave lengths. This is also in agreement with our model derivation. Recall that our model neglects all the behaviour at $O\left(\delta^{*}\right)$, thus $\alpha \gg 1$ will fall into this regime and our model loses validity. Therefore we will concentrate in the limit $\alpha \rightarrow 0$ by studying the long wave length asymptotic behaviour of our problem.

### 4.7 Long wavelength asymptotics

First, we consider the flow of two Newtonian fluids in a vertical or near vertical concentric annulus.

### 4.7.1 Newtonian fluids

We are interested in the behaviour of the solution of (4.57) for small $\alpha$, and we have seen that $s \rightarrow 0$ as $\alpha \rightarrow 0$. Thus we assume that,

$$
s \sim \alpha s_{1}+\alpha^{2} s_{2}+\ldots
$$

Let us use the series expansion for $\operatorname{coth}(x)$ around zero, substitute all these into (4.57) and collect first order terms, then

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$O(1)$ problem

$$
\left(\phi_{i, 0}-\frac{\kappa_{1}}{\kappa_{2}}\left(\phi_{i, 0}-1\right)\right) s_{1}+\epsilon\left(\Psi_{0,1, \phi} \frac{\kappa_{1}}{\kappa_{2}}\left(\phi_{i, 0}-1\right)-\Psi_{0,2, \phi} \phi_{i, 0}\right)=0
$$

Thus,

$$
\begin{equation*}
s_{1}=\frac{\epsilon\left(\Psi_{0,1, \phi} \frac{\kappa_{1}}{\kappa_{2}}\left(1-\phi_{i, 0}\right)+\Psi_{0,2, \phi} \phi_{i, 0}\right)}{\left(\phi_{i, 0}+\frac{\kappa_{1}}{\kappa_{2}}\left(1-\phi_{i, 0}\right)\right)} \tag{4.78}
\end{equation*}
$$

which is a real value. To determine the sign of $s_{1}$ in general is hard, since (4.78) depends on 6 parameters, $\kappa_{1}, \kappa_{2}, \rho_{1}, \rho_{2}, S t^{*}$ and $\phi_{i, 0}$. Now, we continue to the $O(\alpha)$ problem.

## $O(\alpha)$ problem

Collecting $O(\alpha)$ terms, we have the following equation for $s_{2}$,

$$
\begin{aligned}
\frac{1}{\epsilon}\left(\frac{1}{\phi_{i, 0}-1}-\frac{\frac{\kappa_{1}}{\kappa_{2}}}{\phi_{i, 0}}\right) s_{2} & =i s_{1}^{2}\left(\frac{\rho_{2}}{3 \kappa_{2}} \frac{1}{\epsilon\left(\phi_{i, 0}-1\right)}-\frac{\rho_{1}}{3 \kappa_{2}} \frac{1}{\epsilon \phi_{i, 0}}\right) \\
& +i s_{1}\left(\frac{\rho_{1}}{3 \kappa_{2}} \frac{1}{\phi_{i, 0}} \Psi_{0,1, \phi}-\frac{\rho_{2}}{3 \kappa_{2}} \frac{1}{\left(\phi_{i, 0}-1\right)} \Psi_{0,2, \phi}\right) \\
& +i \frac{b}{3 \kappa_{2}} \tan \beta \sin \pi \phi_{i, 0}
\end{aligned}
$$

Therefore,

$$
\begin{align*}
s_{2} & =i s_{1}\left(s_{1}\left(\frac{\rho_{2}}{3 \kappa_{2}} \phi_{i, 0}+\frac{\rho_{1}}{3 \kappa_{2}}\left(1-\phi_{i, 0}\right)\right)-\epsilon\left(\frac{\rho_{1}}{3 \kappa_{2}}\left(1-\phi_{i, 0}\right) \Psi_{0,1, \phi}+\frac{\rho_{2}}{3 \kappa_{2}} \phi_{i, 0} \Psi_{0,2, \phi}\right)\right) \\
& -\frac{i \frac{b}{3 \kappa_{2}}\left(1-\phi_{i, 0}\right) \phi_{i, 0} \tan \beta \sin \pi \phi_{i, 0}}{\phi_{i, 0}+\frac{\kappa_{2}}{\kappa_{2}}\left(1-\phi_{i, 0}\right)} \tag{4.79}
\end{align*}
$$

Clearly, $s_{2}$ is pure imaginary and in the presence of negative buoyancy, i.e. fluid 1 heavier and in top of fluid 2 , and some inclination, i.e. $\beta \neq 0$ the flow may be unstable. On the other hand,

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if we consider a vertical parallel flow, then the stability curves will be determined by the sign of $s_{1}$. For example, let us assume for the moment that $s_{1}>0$, thus

$$
\left(s_{1}\left(\frac{\rho_{2}}{3 \kappa_{2}} \phi_{i, 0}+\frac{\rho_{1}}{3 \kappa_{2}}\left(1-\phi_{i, 0}\right)\right)-\epsilon\left(\frac{\rho_{1}}{3 \kappa_{2}}\left(1-\phi_{i, 0}\right) \Psi_{0,1, \phi}+\frac{\rho_{2}}{3 \kappa_{2}} \phi_{i, 0} \Psi_{0,2, \phi}\right)\right)>0
$$

Substituting (4.78) and after some algebra, we have

$$
\begin{equation*}
\epsilon \dot{\phi_{i, 0}}\left(\frac{\rho_{2}}{\rho_{1}} \frac{\kappa_{1}}{\kappa_{2}}\left(\Psi_{0,1, \phi}-\Psi_{0,2, \phi}\right)+\left(\Psi_{0,2, \phi}-\Psi_{0,1, \phi}\right)\right)>0 \tag{4.80}
\end{equation*}
$$

the inequality is reversed if $s_{1}<0$.
This will lead us to 2 cases,
1.) Case 1 ,

$$
\left(\Psi_{0,2, \phi}-\Psi_{0,1, \phi}\right)>0,
$$

thus

$$
\begin{align*}
\kappa_{1}-\kappa_{2} & <\frac{b}{3} \\
\frac{\kappa_{1}}{\kappa_{2}} & <\frac{\rho_{1}}{\rho_{2}} . \tag{4.81}
\end{align*}
$$

2.) Case 2 ,

$$
\left(\Psi_{0,2, \phi}-\Psi_{0,1, \phi}\right)<0,
$$

thus

$$
\begin{align*}
\kappa_{1}-\kappa_{2} & >\frac{b}{3} \\
\frac{\kappa_{1}}{\kappa_{2}} & >\frac{\rho_{1}}{\rho_{2}} . \tag{4.82}
\end{align*}
$$

If $s_{1}>0$, the cases are analogous with the inequalities reversed.
From these expressions we can see that it is quite difficult to determine the sign of $s_{1}$ in a general way, the 5 parameter dependency makes the analysis very difficult. In spite of this difficulty, we
can infer that the jump in the base velocity is the principal physical cause for the interface to be unstable, i.e. if the jump is sufficiently large, instabilities arise. This translates as follows, the jump in the base flow is governed by viscosity difference, buoyancy and interface position, thus, both viscosity and density ratio play a decisive role in the instability of the flow. It is worth to mention that this phenomena is not comparable to Kelvin-Helmholtz instabilities, see e.g. [18]. This classical instability involves two inviscid fluids one in top of the other, here only inertia, which we have neglected, governs the instability. In [33] the authors describe what they call Kelvin-Helmholtz-Darcy instability. They consider parallel flow in a horizontal Hele-Shaw cell, one of the fluids is a gas, the other is a liquid. They report that the instability is only governed by inertia and that the viscous effects govern the wave velocity and the growth rate only. Our problem concerns the parallel flow of two viscous fluids in a vertical Hele-Shaw cell and even though the nonlinear inertial terms are absent, instabilities arise and are governed by density and viscosity differences. The only possible comparison between the Kelvin-Helmholtz instability and the Kelvin-Helmholtz-Darcy instability with the one found in the previous work is that without difference in velocity the flow is neutrally stable. When we move to the problem of a near vertical Hele-Shaw cell, we see that the viscosity ratio no longer plays an important role on the stability of the flow if the heavier fluid is on top of the lighter one, $b<0$ case. This behaviour is seen in (4.79). If $b<0$ the last term of the right hand side of (4.79) will be always positive.

## Numerical example

Figure 4.23 shows the contours of the imaginary part of $s_{2}$ with respect to $\rho_{2}$ and $\kappa_{2}$. We have fixed $\kappa_{1}=1, \rho_{1}=1, e=0.0, \phi_{i}=0.5$ and $\beta=0$. This is in full agreement with our analytic results. One thing is worth to mention is that when we have the isodensity case, i.e. $b=0$, the flow will be always stable. This is easily seen, if $b=0$ and for example, $\left(\Psi_{0,2, \phi}-\Psi_{0,1, \phi}\right)>0$, from (4.55) we have,

$$
\frac{\kappa_{1}}{\kappa_{2}}>1
$$

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and for instability to arise, from Case 1 we have,

$$
\frac{\kappa_{1}}{\kappa_{2}}<\frac{\rho_{1}}{\rho_{2}}=1
$$

Therefore the flow will always be stable.


Figure 4.23: Contour plot of $s_{2}$. Rheological and physical parameters are: $\kappa_{1}=1, \rho_{1}=1$, $e=0.0, \phi_{i}=0.5$ and $\beta=0$.

### 4.7.2 Yield stress fluids

Now, let us consider the asymptotic behaviour of the full problem when $\alpha \rightarrow 0$. Assume that,

$$
\begin{aligned}
f_{k} & \sim f_{k, 0}+\alpha f_{k, 1}+\alpha^{2} f_{k, 2}+\ldots \\
s & \sim \alpha s_{1}+\alpha^{2} s_{2}+\ldots \\
h & \sim h_{0}+\alpha h_{1}+\alpha^{2} h_{2}+\ldots
\end{aligned}
$$

substituting this expansion into (4.39)-(4.45) and collecting the first order terms we have:

## $O(1)$ problem

$$
\begin{align*}
D\left(\chi_{k}^{\prime} D f_{k, 0}\right) & =0  \tag{4.83}\\
f_{k, 0} & =0 \quad \phi=0,1 \tag{4.84}
\end{align*}
$$

and at $\phi=\phi_{i}$

$$
\begin{align*}
-\frac{H}{\epsilon} s_{1} h_{0}+\left(f_{k, 0}+\Psi_{k, 0, \phi} h_{0}\right) & =0  \tag{4.85}\\
\left.\left(f_{k, 0}+\Psi_{k, 0, \phi} h_{0}\right)\right|_{1} ^{2} & =0  \tag{4.86}\\
\left.\chi_{k}^{\prime} D f_{k, 0}\right|_{1} ^{2} & =0 . \tag{4.87}
\end{align*}
$$

Solving (4.83) we have:

$$
\begin{align*}
f_{1,0}(\phi) & =\int_{0}^{\phi} \frac{C_{1}}{\chi_{1}^{\prime}} \mathrm{d} \phi  \tag{4.88}\\
f_{2,0}(\phi) & =f_{2,0}\left(\phi_{i}\right)+\int_{\phi_{i}}^{\phi} \frac{C_{1}}{\chi_{2}^{\prime}} \mathrm{d} \phi \tag{4.89}
\end{align*}
$$

where

$$
f_{2,0}\left(\phi_{i}\right)=\int_{\phi_{i}}^{\phi} \frac{C_{1}}{\chi_{1}^{\prime}} \mathrm{d} \phi+\left.h_{0} \Psi_{k, 0, \phi}\right|_{2} ^{1}
$$

Applying the boundary condition at $\phi=1$, we can find an expression for $C_{1}$,

$$
\begin{equation*}
C_{1}=\frac{h_{0} \Delta \Psi_{k, 0, \phi}}{I_{1}\left(\phi_{i}\right)+I_{2}(1)} \tag{4.90}
\end{equation*}
$$

where

$$
\begin{equation*}
I_{1}(\phi)=\int_{0}^{\phi} \frac{1}{\chi_{1}^{\prime}} \mathrm{d} \phi \quad I_{2}(\phi)=\int_{\phi_{i}}^{\phi} \frac{1}{\chi_{2}^{\prime}} \mathrm{d} \phi . \tag{4.91}
\end{equation*}
$$

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From (4.85) we get

$$
\begin{equation*}
s_{1} \frac{H\left(\phi_{i}\right)}{\epsilon}=\frac{\Psi_{2,0, \phi} I_{1}\left(\phi_{i}\right)+\Psi_{1,0, \phi} I_{2}(1)}{I_{1}\left(\phi_{i}\right)+I_{2}(1)} \tag{4.92}
\end{equation*}
$$

clearly, $s_{1} \in \mathbf{R}$.

## $O(\alpha)$ problem

$$
\begin{align*}
D\left(\chi_{k}^{\prime} D f_{k, 1}\right) & =\rho_{k} i s_{1} D^{2} f_{k, 0}  \tag{4.93}\\
f_{k, 1} & =0 \quad \phi=0,1 \tag{4.94}
\end{align*}
$$

and at $\phi=\phi_{i}$

$$
\begin{align*}
\frac{H}{\epsilon}\left(s_{2} h_{0}+s_{1} h_{1}\right)+\left(f_{k, 1}+\Psi_{k, 0, \phi} h_{1}\right) & =0  \tag{4.95}\\
\left.\left(f_{k, 1}+\Psi_{k, 0, \phi} h_{1}\right)\right|_{1} ^{2} & =0  \tag{4.96}\\
\left.\left(\chi_{k}^{\prime} D f_{k, 1}-i s_{1} \rho_{k} D f_{k, 0}\right)\right|_{1} ^{2} & =i b h_{0} \tan \beta \sin \pi \phi_{i, 0} \tag{4.97}
\end{align*}
$$

Integrating (4.93) once, we have:

$$
\begin{equation*}
\chi_{k}^{\prime} D f_{k, 1}-\rho_{k} i s_{1} D f_{k, 0}=C_{2, k} \tag{4.98}
\end{equation*}
$$

using (4.97) we get,

$$
C_{2,2}=C_{2,1}+i b h_{0} \tan \beta \sin \pi \phi_{i, 0}
$$

Integrating (4.98) and using (4.96) we can find the solutions to (4.93),

$$
\begin{align*}
f_{1,1}(\phi) & =\int_{0}^{\phi}\left(\frac{C_{2,1}}{\chi_{1}^{\prime}}+i s_{1} \rho_{1} \frac{C_{1}}{\left(\chi_{1}^{\prime}\right)^{2}}\right) \mathrm{d} \phi  \tag{4.99}\\
f_{2,1}(\phi) & =f_{1,1}\left(\phi_{i}\right)-h_{1} \Delta \Psi_{k, 0, \phi}+\int_{\phi_{i}}^{\phi}\left(\frac{C_{2,2}}{\chi_{2}^{\prime}}+i s_{1} \rho_{2} \frac{C_{1}}{\left(\chi_{2}^{\prime}\right)^{2}}\right) \mathrm{d} \phi \tag{4.100}
\end{align*}
$$

where $C_{1}$ is defined in (4.90).

Using the boundary condition at $\phi=1$ we can find an expression for $C_{2,1}$

$$
\begin{equation*}
C_{2,1}=h_{1} \frac{\Delta \Psi_{k, 0, \phi}}{I_{1}\left(\phi_{i}\right)+I_{2}(1)}-C_{1} i s_{1} \frac{J_{1}\left(\phi_{i}\right)+J_{2}(1)}{I_{1}\left(\phi_{i}\right)+I_{2}(1)}-i b h_{0} \tan \beta \sin \pi \phi_{i, 0} \frac{I_{2}(1)}{I_{1}\left(\phi_{i}\right)+I_{2}(1)} . \tag{4.101}
\end{equation*}
$$

where $I_{1} \& I_{2}$ are defined in (4.91) and

$$
\begin{equation*}
J_{1}(\phi)=\int_{0}^{\phi} \frac{\rho_{1}}{\left(\chi_{1}^{\prime}\right)^{2}} \mathrm{~d} \phi \quad J_{2}(\phi)=\int_{\phi_{i}}^{\phi} \frac{\rho_{2}}{\left(\chi_{2}^{\prime}\right)^{2}} \mathrm{~d} \phi . \tag{4.102}
\end{equation*}
$$

Using the kinematic condition (4.95), (4.99), (4.90) and (4.101) we have an expression for $s_{2}$,

$$
\begin{equation*}
-\frac{H\left(\phi_{i, 0}\right)}{\epsilon} s_{2}=i b \tan \beta \sin \pi \phi_{i, 0} \frac{I_{1}\left(\phi_{i}\right) I_{2}(1)}{I_{1}\left(\phi_{i}\right)+I_{2}(1)}+i s_{1} \Delta \Psi_{k, 0, \phi} \frac{J_{1}\left(\phi_{i}\right) I_{2}(1)-J_{2}(1) I_{1}\left(\phi_{i}\right)}{\left(I_{1}\left(\phi_{i}\right)+I_{2}(1)\right)^{2}}, \tag{4.103}
\end{equation*}
$$

clearly, $s_{2}=s_{2, I}$.
Note the first term of the right hand side of (4.103), as in previous cases, if $b<0$, i.e. heavier fluid 1 is on top of lighter fluid 2, the buoyancy has destabilizing effects. Turning to the second term of the right hand side of (4.103), we see that again, the jump in the velocity plays an important role in the stability of the flow. But things get more complicated, note that to determine the sign of $s_{2}$, we have to investigate the effects of 11 parameters on (4.103). Fortunately, from §4.6.2 we know that the position of the interface does not play an important role on the stability of the flow. We also proved that eccentricity has destabilizing effects on
the flow as well as inclination. Therefore we will concentrate in investigating the effects of the 8 rheological and physical parameters on the stability of the flow when we are in the case of a mildly eccentric vertical annulus.

Figure 4.24 shows the level curves of the maximum imaginary part of $s_{2}$ for rheological parameters, $\rho_{1}=1, \rho_{2}=1, m_{1}=1, m_{2}=1, \tau_{Y, 1}=1, \tau_{Y, 2}=1$. The neutral stability level curve, $\kappa_{1}=\kappa_{2}$ implies that if $k_{1}>k_{2}$, i.e. the displacing fluid is less viscous than the displaced one, the flow will be stable.

In Figure 4.25 a) and Figure 4.25 b) we have decreased and increased, respectively, the buoyancy parameter. When the displacing fluid, fluid 1 , is denser than the displaced fluid, fluid 2, fluid 2 is pushed up the annulus by buoyancy. This increases the jump in the velocity and a wider unstable region develops. Increasing the buoyancy parameter has the opposite effect, now the displaced fluid is denser than the displacing one and the jump in the velocity decreases, resulting in a thinner unstable region.

Figure 4.26 shows the effects of the difference in power law index. Figure 4.26 a), $m_{1}=1.2$ and $m_{2}=1$, can be explained as follows, fluid 1 now is more shear-thinning than fluid 2 , resulting in a shift of the level curve for $\kappa_{1}=\kappa_{2}$ to the stable region. The opposite effect happens when $m_{1}=1$ and $m_{2}=1.2$, Figure 4.26 b$)$. Now the level curve for $\kappa_{1}=\kappa_{2}$ has been shifted to the unstable region, clear effect of fluid 2 being more shear-thinning than fluid 1.

Figure 4.27 shows the effects of yield stress on the stability of the flow, for Figure 4.27 a) fluid 1 has a higher yield stress than fluid $2, \tau_{Y, 1}=1.5$ and $\tau_{Y, 2}=1$. Because of this, a higher pressure gradient is required to overcome $\tau_{Y, 1}$ for both fluids to be yielded at the interface. This could be seen as fluid 1 being "more viscous", thus the level curve $\kappa_{1}=\kappa_{2}$ is shifted to the unstable region. If now $\tau_{Y, 2}=1.5$ and $\tau_{Y, 1}=1$, Figure 4.27 b ), we see that fluid 2 is "more viscous" and the level curve $\kappa_{1-}=\kappa_{2}$ is shifted to the stable region.

To the knowledge of the author, this is a phenomena never reported in the literature. For this kind of flow, even in the absence of inertia, the flow can be unstable. All rheological parameters
play a role on the stability of the flow, something not seen in Newtonian and power law fluids. For yield stress fluids, even without the presence of buoyant forces the flow can be unstable.

Following the results found in [59] we can make some comparison of our contour plots and unstable regions for the cases when the displacement is unstable.


Figure 4.24: Contour plot of $s_{2}, \kappa_{1}$ vs. $\kappa_{2}$. Rheological and physical parameters are: $\rho_{1}=1$, $m_{1}=1, m_{2}=1, \tau_{Y, 1}=1, \tau_{Y, 2}=1, \rho_{2}=1, e=0.2, \phi_{i}=0.5$ and $\beta=0$.

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Figure 4.25: Contour plot of $s_{2}, \kappa_{1}$ vs. $\kappa_{2}$. Rheological and physical parameters are: $m_{1}=1$, $m_{2}=1, \tau_{Y, 1}=1, \tau_{Y, 2}=1, e=0.2, \phi_{i}=0.5$ and $\beta=0$. a) $\rho_{1}=1.1$ and $\rho_{2}=1$. b) $\rho_{1}=1$ and $\rho_{2}=1.1$


Figure 4.26: Contour plot of $s_{2}, \kappa_{1}$ vs. $\kappa_{2}$. Rheological and physical parameters are: $\rho_{1}=1$, $\tau_{Y, 1}=1, \tau_{Y, 2}=1, \rho_{2}=1.1, e=0.2, \phi_{i}=0.5$ and $\beta=0$. a) $m_{1}=1.2$ and $m_{2}=1$. b) $m_{1}=1$ and $m_{2}=1.2$,

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Figure 4.27: Contour plot of $s_{2}, \kappa_{1}$ vs. $\kappa_{2}$, buoyancy parameter equal zero. Rheological and physical parameters are: $\kappa_{1}=1, \rho_{1}=1, \kappa_{2}=1, \rho_{2}=1, e=0.2, \phi_{i}=0.5$ and $\beta=0$. a) $\tau_{Y, 1}=1.5$ and $\tau_{Y, 2}=1$. b) $\tau_{Y, 1}=1$ and $\tau_{Y, 2}=1.5$.

### 4.7.3 Industrial application

In this section we make some comparisons with the results found in [59]. In [59] the authors use a lubrication model to predict, for a given set of parameters, whether a good displacement takes place or not. If they have an unsteady displacement, i.e. the interface front moves faster in the wide part than in the narrow part, a long finger develops along the annulus. This situation is when our assumption of parallel flow is likely to occur. Furthermore, they investigate when a viscous fingering instability will occur during the displacement. They showed that viscous fingering happens only in the unsteady displacement regime.

Figure 4.28 (from [59]), shows when a unsteady displacement will occur, U meaning unsteady displacement but no viscous fingering, F meaning unsteady displacement and viscous fingering, $S$ meaning neither unsteady nor viscous fingering. Rheological and physical parameters are, $\tau_{Y, 1}=1.0, \kappa_{1}=1.0, m_{1}=1, m_{2}=2, b=-0.5, \beta=0$. It is clear from this picture that as the authors increase $e, \tau_{Y, 2}$ and $\kappa_{2}$, i.e. viscosity of fluid 2 increasing, an increase of the domain of

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unsteadiness and instability is observed.

In Figure 4.29 we show the level curves of the maximum imaginary part of $s_{2}$ as a function fo $\tau_{Y, 2}$ and $e$. We have the same fixed rheological parameters as in Figure 4.28 with $\phi_{i}=0.5$ and $\beta=0$. We capture the same behavior as in Figure 4.28, as fluid 2 becomes more viscous, the unstable domain increases. Surprisingly, our unstable domain occurs only as a subset of the unsteady and fingering domain in Figure 4.28. This means that if a long finger develops during the displacement, for certain parameter ranges the parallel flow may be unstable. Therefore, apart from having an unsteady displacement, mixing may occur at the interface of the finger due to the instability that we study here. It is not clear why there is apparently no interfacial instability for parameter regimes for which the annular displacement would be steady. Note that for such parameters, a base parallel flow is a perfectly well-defined solution of the two fluid problem.

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Figure 4.28: Example unsteady displacement and local instability/viscous fingering regimes: effects of changing $\tau_{Y, 2}$ and $e$ for different $\kappa_{2}$ : (a) $\kappa_{2}=0.5$; (b) $\kappa_{2}=1$; (c) $\kappa_{2}=1.5$. Fixed parameters are: $\tau_{Y, 1}=1.0, \kappa_{1}=1.0, m_{1}=1, m_{2}=2, b=-0.5, \beta=0$. F, local fingering instability ; U, unsteady, but no local instability; S, neither unsteady. From Pelipenko \& Frigaard (2004):


Figure 4.29: Maximum imaginary part of $s$ vs. $\alpha$. Rheological and physical parameters are: $\kappa_{1}=1, \tau_{Y, 1}=1, m_{1}=1, m_{2}=2, b=-0.5, \phi_{i}=0.5$ and $\beta=0$. a) $\kappa_{2}=0.5$. b) $\kappa_{2}=1$. c) $\kappa_{2}=1.5$.

### 4.8 Spatial stability problem

Finally, we present a collection of results concerning linear spatial instabilities. Although incomplete, the results look interesting and potentially of practical value.

### 4.8.1 Single fluid spatial instability

Let us consider, for the single fluid problem, a spatially developing instability: $s>0$ and $\alpha=\alpha_{R}+\mathrm{i} \alpha_{I}$. In this case, (4.25) constitutes an eigenvalue problem for $\alpha^{2} \in \boldsymbol{C}$. We find:

$$
\begin{gather*}
-\left(\alpha^{2}\right)_{R}=\alpha_{I}^{2}-\alpha_{R}^{2}=\frac{s^{2} J_{1} I_{1}+J_{2} I_{2}}{s^{2} I_{2}^{2}+I_{1}^{2}}=C_{1} \geq 0  \tag{4.104}\\
\left(\alpha^{2}\right)_{I}=2 \alpha_{I} \alpha_{R}=\frac{s\left[J_{1} I_{2}-J_{2} I_{1}\right]}{s^{2} I_{1}^{2}+I_{2}^{2}}=C_{2} \tag{4.105}
\end{gather*}
$$

Solving for $\alpha_{R} \& \alpha_{I}$, we have:

$$
\begin{align*}
& \alpha_{I}= \pm \sqrt{\frac{C_{1}+\sqrt{C_{1}^{2}+C_{2}^{2}}}{2}}  \tag{4.106}\\
& \alpha_{R}= \pm \sqrt{\frac{-C_{1}+\sqrt{C_{1}^{2}+C_{2}^{2}}}{2}} \tag{4.107}
\end{align*}
$$

It appears that any non-trivial complex eigenvalue $\alpha^{2}$ and eigenfunction $f$ must admit spatially growing instabilities.

### 4.8.2 Two fluids spatial instability

In an analogous way as in the previous section, let us assume that: $s>0$ and $\alpha=\alpha_{R}+\mathrm{i} \alpha_{I}$.
We assume small $s$ and study the asymptotic behavior of the problem. Thus, let

$$
\begin{aligned}
f_{k} & \sim f_{k, 0}+s f_{k, 1}+s^{2} f_{k, 2}+\ldots \\
\alpha & \sim \alpha_{0}+s \alpha_{1}+s^{2} \alpha_{2} \ldots \\
h & \sim h_{0}+s h_{1}+s^{2} h_{2}+\ldots
\end{aligned}
$$

substituting this expansion into (4.39)-(4.45) and collecting the first order terms we have:

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$$
\begin{align*}
& D\left(\chi_{k}^{\prime} D f_{k, 0}\right)-\alpha_{0}^{2} f_{k, 0}=0, \\
& f_{k, 0}=0 \quad \text { at } \phi=0,1, \\
& \alpha_{0}\left(f_{k, 0}+\Psi_{k, 0, \phi} h_{0}\right)=0, \\
& \left.\left(f_{k, 0}+\Psi_{k, 0, \phi} h_{0}\right)\right|_{1} ^{2}=0, \\
& \left.\chi_{k}^{\prime} D f_{k, 0}\right|_{1} ^{2}=i \alpha_{0} h_{0} \tan \beta \sin \pi \phi_{i, 0} . \tag{4.108}
\end{align*}
$$

System (4.108) has solutions of the form,

$$
f_{k, 0}(\phi)=f_{k, 0, p}(\phi)+f_{k, 0, h}(\phi),
$$

where

$$
\begin{aligned}
& D\left(\chi_{k}^{\prime} D f_{k, 0, p}\right)=0 \\
& f_{k, 0, p}=0 \quad \text { at } \phi=0,1, \\
& f_{k, 0, p}=-\Psi_{k, 0, \phi} h_{0} .
\end{aligned}
$$

Therefore,

$$
\begin{align*}
f_{1,0, p} & =-h_{0} \Psi_{1,0, \phi} \frac{I_{1}(\phi)}{I_{1}\left(\phi_{i}\right)}  \tag{4.109}\\
f_{1,0, p} & =-h_{0} \Psi_{2,0, \phi} \frac{I_{2}(\phi)}{I_{2}\left(\phi_{i}\right)} \tag{4.110}
\end{align*},
$$

where

$$
I_{1}(\phi)=\int_{0}^{\phi} \frac{1}{\chi_{1}^{\prime}} \mathrm{d} \phi \quad I_{2}(\phi)=\int_{\phi}^{1} \frac{1}{\chi_{2}^{\prime}} \mathrm{d} \phi
$$

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Now, $f_{k, 0, h}$ is an eigen-function of

$$
D\left(\chi_{k}^{\prime} D f_{k, 0, h}\right)=\alpha_{0}^{2} f_{k, 0, h}
$$

with $\alpha_{0}^{2}$ being the eigen-value. Therefore, from Sturm-Liouville theory, $\alpha_{0}^{2}$ is real and less than zero. Thus,

$$
\alpha_{R}=0 \quad \text { and } \quad \alpha_{I} \neq 0
$$

Our solution, $f_{k, 0}(\phi)=f_{k, 0, p}(\phi)+f_{k, 0, h}(\phi)$ has to satisfy the second jump condition, then

$$
\begin{equation*}
h_{0}\left(\Psi_{2,0, \phi} \frac{I_{2}(\phi)}{I_{2}\left(\phi_{i}\right)}+\Psi_{1,0, \phi} \frac{I_{1}(\phi)}{I_{1}\left(\phi_{i}\right)}\right)+\chi_{2}^{\prime} D f_{2,0, h}-\chi_{1}^{\prime} D f_{1,0, h}=-\alpha_{I} h_{0} \tan \beta \sin \pi \phi_{i} \tag{4.111}
\end{equation*}
$$

equation (4.111) defines $h_{0}$.

Therefore, condition for instability is that $\lambda=-\alpha_{I}^{2}$ is an eigen-value of both:

$$
\begin{aligned}
& D\left(\chi_{1}^{\prime} D f\right)=\lambda f \\
& f(0)=f\left(\phi_{i}\right)=0 \\
& D\left(\chi_{2}^{\prime} D f\right)=\lambda f \\
& f\left(\phi_{i}\right)=f(1)=0 .
\end{aligned}
$$

For 2 Newtonian fluids at a fixed viscosity ratio, we can easily show that there are interfacial positions, $\phi_{i}$, for which $\lambda$ satisfies both problems. It is unclear whether in general this occurs and we have not analyzed much further.

## Chapter 5

## Static mud channels

As stated in Chapter 4, for certain combinations of dimensionless parameters, it is possible for one of the fluids to be stationary on the narrow side of the annulus. Physically, this occurs when the modified pressure gradient is not large enough to overcome the yield stress of the fluid. From an industrial perspective the flows within this category that have most interest are those in which fluid 2 is static, either partly or fully, i.e. there is a static mud channel. Since the stability studies only predict linear stability of the flow, we ask whether a nonlinear perturbation could shift the mud channel. We study the displacement via numerical simulation.

We will solve numerically the full system,

$$
\begin{equation*}
\nabla \cdot\left[\rho \nabla \Psi_{t}\right]=-\nabla \cdot[\mathbf{S}+\mathbf{f}] \tag{5.1}
\end{equation*}
$$

where

$$
\begin{gather*}
\mathbf{S}=\left(\frac{\chi(|\nabla \Psi|)+\tau_{Y} / H}{|\nabla \Psi|}\right) \nabla \Psi \Leftrightarrow|\mathbf{S}|>\tau_{Y} / H  \tag{5.2}\\
|\nabla \Psi|=0 \Leftrightarrow|\mathbf{S}| \leq \tau_{Y} / H  \tag{5.3}\\
\Psi(0, \xi, t)=0  \tag{5.4}\\
\Psi(1, \xi, t)=Q(t)  \tag{5.5}\\
\Psi(\phi, 0, t)=\Psi_{i n}(\phi, t),  \tag{5.6}\\
\Psi(\phi, Z, t)=\Psi_{o u t}(\phi, t),  \tag{5.7}\\
\frac{1}{\epsilon} \frac{\partial}{\partial t}[H \bar{c}]+\frac{\partial}{\partial \phi}[H \bar{v} \bar{c}]+\frac{\partial}{\partial \xi}[H \bar{w} \bar{c}]=0 \tag{5.8}
\end{gather*}
$$

Whereas for the stability studies, we have $Q(t)=1$, here we consider varying flow rate, such as would be possible with pressure pulsing. For the velocity field we use an augmented Lagrangian algorithm, $[26,31]$, which fully describes the unyielded regions of the flow, the method is outlined in §5.1. We will work in a finite element approximation setting for the velocity field, with bilinear basis functions and square elements, for an summary of the finite element method see $\S 5.3$. The concentration equation is fully discretized using a conservative finite volume approximation at the middle point of each element and solved using a Flux Corrected Transport scheme, see §5.4. The FCT scheme is used to minimize both numerical dispersion and diffusion. Finally in $\S 5.5$ we investigate the effect of pulsing the flow rate on the displacement fronts.

### 5.1 Augmented Lagrangian Method

The augmented Lagrangian method, although it accurately predicts unyielded regions in the flow of a visco-plastic fluid, is not widely used for simulating this kind of flow. Instead a regularised version of the constitutive laws of the fluid is used. This consists of considering, at the point of singularity, (i.e. shear rate equal zero), that the fluid has instead a very large Newtonian viscosity. For some recent work where the augmented Lagrangian algorithm has been successfully implemented we refer the reader to $[80,81,82]$. This section mainly gives an outline of the augmented Lagrangian algorithm, we will follow closely $[26,31]$.

### 5.1.1 Principles of the method

For the purpose of this thesis we will limit ourselves to a finite dimensional problem, i.e. since in any case we eventually discretise and solve a finite dimensional problem.

## Augmented Lagrangian

As an example, let $A$ be a symmetric, positive definite $N \times N$ matrix and suppose that $b \in \mathbf{R}$.
We associate the following quadratic functional to these parameters

$$
\begin{equation*}
J(v)=\frac{1}{2}(A v, v)-(b, v) \tag{5.9}
\end{equation*}
$$

## Chapter 5. Static mud channels

where in $(5.9),(\cdot, \cdot)$ is the canonical Euclidian inner product in $\mathbf{R}^{N}$. Let $T$ be a linear mapping from $\mathbf{R}^{N}$ into $\mathbf{R}^{M}$. Now, consider the minimisation problem

$$
\begin{align*}
J(u) & \leq J(v) \quad \forall v \in \operatorname{Ker} T  \tag{5.10}\\
u & \in \operatorname{Ker} T \tag{5.11}
\end{align*}
$$

Because of the properties of $A$ it is clear that (5.10) has a unique solution, but the constraint (5.11) may make the solution difficult to compute.

Introducing a Lagrange multiplier $p$ to (5.10), in order to get an unconstrained problem, we have

$$
\begin{align*}
& \min \{J(v)+(p, T v)\}  \tag{5.12}\\
& v \in \mathbf{R}^{N}
\end{align*}
$$

Now the problem is unconstrained but the Lagrange multiplier $p$ appears as an extra unknown. This can be obtained through the solution of a saddle-point problem.

Thus, we define

$$
\begin{equation*}
L(v, q)=J(v)+(q, T v) \tag{5.13}
\end{equation*}
$$

recall that ( $u, p$ ) will be a saddle-point of $L$ if

$$
\begin{equation*}
L(u, q) \leq L(u, p) \leq L(v, p) \quad \forall v \in \mathbf{R}^{N}, q \in \mathbf{R}^{M} \tag{5.14}
\end{equation*}
$$

thus we have

$$
\begin{equation*}
\min _{v \in \mathbf{R}^{N}} \max _{q \in \mathbf{R}^{M}} L(v, q)=\max _{q \in \mathbf{R}^{M}} \min _{v \in \mathbf{R}^{N}} L(v, q)=L(u, p) \tag{5.15}
\end{equation*}
$$

It can be shown that $L$ admits at least one saddle-point $(u, p)$ on $\mathbf{R}^{N} \times \mathbf{R}^{M}$, where $u$ is the solution of (5.10) and is common to all the saddle-points of $L$ on $\mathbf{R}^{N} \times \mathbf{R}^{M}$.

We define the augmented Lagrangian $L_{r}$, for $r>0$, by

$$
\begin{equation*}
L_{r}(v, q)=J(v)+(q, T v)+\frac{r}{2}|T v|^{2}=L(v, q)+\frac{r}{2}|T v|^{2} \tag{5.16}
\end{equation*}
$$

## Chapter 5. Static mud channels

and note that a saddle-point of $L(v, q)$ for which $u \in \operatorname{Ker} T$, will also be a saddle-point of $L_{r}(v, q)$. However, (5.16) may be better conditioned to solve numerically.

### 5.1.2 ALG2

In $[26,31]$ the authors describe an algorithm for finding the saddle-points of an augmented Lagrangian functional $L_{r}$, they call it ALG2.

Let us consider the minimization problem

$$
\begin{align*}
& \min \{F(T v)+G(v)\}  \tag{5.17}\\
& v \in V
\end{align*}
$$

as our generalized minimization problem, where

- $V$ and $H$ are Hilbert spaces.
- $T: V \mapsto H$.
- where $F$ and $G$ are convex, proper, and lower semi-continuous (l.s.c) on $H$ and $V$, respectively. It is know that (5.17) and

$$
\begin{align*}
& \min \{F(q)+G(v)\}  \tag{5.18}\\
& (v, q) \in W
\end{align*}
$$

with

$$
W=\{(v, q) \in V \times H, T v-q=0\}
$$

are equivalent problems.

Introducing the variable $q$, linked to $v$ through the linear equality equation $T v=q$ we can reformulate (5.18) as a saddle-point problem.

We will in fact take the spaces $V$ and $H$ to be $V=V_{\rho}$ and $H=L^{2}(\Omega) \times L^{2}(\Omega)$ with their respective inner products and associated norms. We define, for $v \in V_{\rho, 0}, q \in H$ and $\mu \in H$ the Lagrangian functional

$$
L(v, q, \mu)=F(q)+G(v)+(\mu, T v-q)
$$

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and then for $r \geq 0$, the augmented Lagrangian functional is defined as

$$
\begin{equation*}
L(v, q, \mu)_{r}=L(v, q, \mu)+\frac{r}{2}\|T v-q\|^{2} \tag{5.19}
\end{equation*}
$$

The iterative method ALG2 that we use below is in fact a method for calculating the saddlepoints $(u, p, \lambda)$ of (5.19).

ALG2

$$
\begin{equation*}
\left(p^{0}, \lambda^{1}\right) \in H \times H \text { arbitrarily specified; } \tag{5.20}
\end{equation*}
$$

with $\left(p^{n-1}, \lambda^{n}\right)$ known, determine succesively $u^{n}, p^{n}, \lambda^{n+1}$ by

$$
\begin{gather*}
G(v)-G\left(u^{n}\right)+\left(\lambda^{n}, T\left(v-u^{n}\right)\right) \\
+r\left(B u^{n}-p^{n-1}, T\left(v-u^{n}\right)\right) \geq 0 \quad \forall v \in V_{\rho},  \tag{5.21}\\
u^{n} \in V_{\rho} \\
\left.F(q)-F\left(p^{n}\right)-\left(\lambda^{n}, q-p^{n}\right)+r\left(p^{n}-T u^{n}, q-p^{n}\right)\right) \geq 0 \\
\forall q \in H \times H, \text { and } p^{n} \in H \times H \\
\lambda^{n+1}=\lambda^{n}+\rho_{n}\left(T u^{n}-p^{n}\right) . \tag{5.23}
\end{gather*}
$$

### 5.1.3 Convergence of ALG2

Results on convergence of ALG2 are presented below, we just give a summary of what conditions $T, F$ and $G$ have to satisfy and state the convergence result. For a full description of the proof we refer to the reader to [26].

## Chapter 5. Static mud channels

## Conditions for $T, F$ and $G$

(i) $T$ should be injective and the image of $T$ should be closed in $H$.
(ii)

$$
\begin{equation*}
\lim _{|q| \rightarrow \infty} \frac{F(q)}{|q|}=\infty \tag{5.24}
\end{equation*}
$$

(iii) Because $G$ is convex, proper and l.s.c and using (i) and (ii) we have that (5.17) admits a unique solution $u$.
(iv) Let us assume that $F=F_{0}+F_{1}$, where $F_{1}$ is convex, proper and l.s.c. on $H$, and $F_{0}$ is convex, single valued, Gateaux differentiable on $H$, and uniformly convex over the bounded subsets of $H$ in the following sense:

For any $M>0$, there exists a continuous function $\delta_{M}:[0,2 M] \rightarrow \boldsymbol{R}$, strictly increasing, with $\delta_{0}=0$, such that for all $p, q \in H,|p| \leq M,|q| \leq M$, we have:

$$
\begin{equation*}
\left(F_{0}^{\prime}(q)-F_{0}^{\prime}(p), q-p\right) \geq \delta_{M}(|q-p|) . \tag{5.25}
\end{equation*}
$$

Now we are in a position to state the main result for convergence of the algorithm.

Theorem 3 We assume that $L_{r}$ admits a saddle-point $(u, p, \lambda)$ on $V_{\rho} \times H \times H$. Under the assumption on $T, F, G(i)-(i v)$ and if $\rho_{n}$ satisfies

$$
\begin{equation*}
0<\rho_{n}=\rho<\left(\frac{1+\sqrt{5}}{2}\right) r, \tag{5.26}
\end{equation*}
$$

we have for $\boldsymbol{A L G 2}$ the following convergence results

$$
\begin{align*}
& u^{n} \rightarrow u \text { strongly in } V,  \tag{5.27}\\
& p^{n} \rightarrow p \text { strongly in } H, \tag{5.28}
\end{align*}
$$

$$
\begin{equation*}
\lambda^{n+1}-\lambda^{n} \rightarrow 0 \text { strongly in } H \tag{5.29}
\end{equation*}
$$

$$
\begin{equation*}
\lambda^{n} \text { is bounded in } H . \tag{5.30}
\end{equation*}
$$

Proof

For a proof see [26].

### 5.2 Application of ALG2 to the displacement problem.

### 5.2.1 Steady model

Let us take

$$
\begin{aligned}
& T(u)=\nabla u, \\
& G(u)=0, \\
& F(\boldsymbol{q})=F_{0}(\boldsymbol{q})+F_{1}(\boldsymbol{q}),
\end{aligned}
$$

where

$$
\begin{align*}
& F_{0}(\boldsymbol{q})=\int_{\Omega}\left(\frac{1}{2} \int_{0}^{\left|\nabla \Psi^{*}+\boldsymbol{q}\right|^{2}} \frac{\chi\left(s^{1 / 2}\right)}{s^{1 / 2}} \mathrm{~d} s+\boldsymbol{q} \cdot \mathbf{f}\right) \mathrm{d} \Omega, \quad \boldsymbol{q} \in H \times H \\
& F_{1}(\boldsymbol{q})=\int_{\Omega} \frac{\tau_{Y}}{H}\left|\nabla \Psi^{*}+\boldsymbol{q}\right| \mathrm{d} \Omega, \quad \boldsymbol{q} \in H \times H \tag{5.31}
\end{align*}
$$

Let us stop for a moment to make a couple of remarks.

We have homogenized the boundary conditions with $\Psi^{*}$ defined in (3.13).

Recall that if $\mu_{\infty} \neq 0$ we are in a Hilbert space setting, for the limit $\mu_{\infty} \rightarrow 0$ we would have to work in the Banach spaces $W^{1,1+1 / m}(\Omega)$ for $\Psi$ and $\left(L^{1,1+1 / m}(\Omega) \times L^{1,1+1 / m}(\Omega)\right)$ for $q$. The
augmented Lagrangian algorithm requires the solution space to be a Hilbert space. Nevertheless, once we have discretized our variational problem using a finite-dimensional numerical method (Finite Element Method for example), the approximate solution space $V_{\rho, h}$ is a finite dimensional Hilbert space. The subscript $h$ denotes the mesh scale and although the technical difficulty remains in the limit $h \rightarrow 0$, in practice we compute the approximate solution with a small finite $h$. Thus it is safe to assume that our finite dimensional solution space to be $V_{\rho, h} \subset H^{1}(\Omega)$ and similarly we can take $\boldsymbol{q} \in H$.

## Outline of ALG2 for a finite dimensional setting

For a given constant $r>0$ and using the definitions in (5.31) our augmented Lagrangian functional is,

$$
\begin{equation*}
L_{r}(v, \boldsymbol{q}, \boldsymbol{\mu})=F(\boldsymbol{q})+\int_{\Omega \Omega} \boldsymbol{\mu} \cdot(\nabla v-\boldsymbol{q}) \mathrm{d} \Omega+\frac{r}{2} \int_{\Omega}|\nabla v-\boldsymbol{q}|^{2} \mathrm{~d} \Omega \tag{5.32}
\end{equation*}
$$

Therefore, if ( $u, \boldsymbol{p}, \boldsymbol{\lambda}$ ), where $u \in V_{0, \rho}$ and $u$ is defined as $\Psi=\Psi^{*}+u$, is a saddle point of $L_{r}$, $\forall r>0$, then $\Psi^{*}+u$ is a solution to (3.25).

Now we are in a position to apply ALG2 to our problem. Each step of the algorithm consists of finding $u^{n}, p^{n}, \lambda^{n+1}$ by solving sequentially:

$$
\begin{align*}
& \int_{\Omega} \lambda^{n} \cdot \nabla\left(v-u^{n}\right)+\left(r \nabla u^{n}-\boldsymbol{p}^{n-1}\right) \cdot \nabla\left(v-u^{n}\right) \mathrm{d} \Omega \geq 0  \tag{5.33}\\
& \forall v \in V_{0, \rho}, u^{n} \in V_{0, \rho} \\
& F(\boldsymbol{q})-F\left(\boldsymbol{p}^{n}\right)+\int_{\Omega} \lambda^{n} \cdot\left(\boldsymbol{q}-\boldsymbol{p}^{n}\right)+r\left(\boldsymbol{p}^{n}-\nabla u^{n}\right) \cdot\left(\boldsymbol{q}-\boldsymbol{p}^{n}\right) \mathrm{d} \Omega \geq 0  \tag{5.34}\\
& \forall \boldsymbol{q} \in H, \boldsymbol{p}^{n} \in H \\
& \lambda^{n+1}=\lambda^{n}+\rho_{n}\left(\nabla u^{n}-\boldsymbol{p}^{n}\right), \quad \rho_{n}>0 \tag{5.35}
\end{align*}
$$

Minimizing equation (5.33) with respect to $v$ is equivalent to solving the following Poisson
equation,

$$
\begin{equation*}
r \Delta u^{n}=r \nabla \cdot \boldsymbol{p}^{n-1}-\nabla \cdot \boldsymbol{\lambda}^{n}, \tag{5.36}
\end{equation*}
$$

where the right hand side is known from the previous step. The second step consists of minimizing equation (5.34) with respect to $\boldsymbol{q}$ in order to obtain $\boldsymbol{p}^{n}$. This second step is the critical one and shows the advantage of this method. For the case of visco-plastic fluids, in this step we will fully describe the yielded and unyielded regions. In minimizing (5.34) with respect to $\boldsymbol{q}$ in a discrete setting, we may locally minimize the integrand of $\boldsymbol{L}_{r}\left(u^{n}, \boldsymbol{q}, \boldsymbol{\lambda}^{n}\right)$ with respect to $\boldsymbol{q} \in H_{h} \times H_{h}$ where $H_{h}$ denotes the restriction of $H$ to a particular discrete element, thus

$$
\begin{equation*}
\boldsymbol{p}^{n}=\inf _{\boldsymbol{q}}\left\{\frac{1}{2} \int_{0}^{\left|\nabla \Psi^{*}+\boldsymbol{q}\right|^{2}} \frac{\chi\left(s^{1 / 2}\right)}{s^{1 / 2}} \mathrm{~d} s+\boldsymbol{q} \cdot \mathbf{f}+\frac{\tau_{Y}}{H}\left|\nabla \Psi^{*}+\boldsymbol{q}\right|+\frac{r}{2}|\boldsymbol{q}|^{2}-\left(\boldsymbol{\lambda}^{n}+r \nabla u^{n}\right) \cdot \boldsymbol{q}\right\} \tag{5.37}
\end{equation*}
$$

equivalently we have,

$$
\begin{equation*}
\boldsymbol{p}^{n}=\inf _{\boldsymbol{q}}\left\{\frac{1}{2} \int_{0}^{\left|\nabla \Psi^{*}+\boldsymbol{q}\right|^{2}} \frac{\chi\left(s^{1 / 2}\right)}{s^{1 / 2}} \mathrm{~d} s+\frac{\tau_{Y}}{H}\left|\nabla \Psi^{*}+\boldsymbol{q}\right|+\frac{r}{2}\left|\nabla \Psi^{*}+\boldsymbol{q}\right|^{2}-\left(\boldsymbol{\lambda}^{n}+r \nabla\left(\Psi^{*}+u^{n}\right)-\boldsymbol{f}\right) \cdot\left(\nabla \Psi^{*}+\boldsymbol{q}\right)\right\} . \tag{5.38}
\end{equation*}
$$

Clearly, this expression will achieve its minimum when

$$
\left(\nabla \Psi^{*}+\boldsymbol{q}\right) \|\left(\boldsymbol{\lambda}^{n}+r \nabla\left(\Psi^{*}+u^{n}\right)-\boldsymbol{f}\right) .
$$

Therefore, letting

$$
\begin{align*}
\left(\nabla \Psi^{*}+\boldsymbol{p}^{n}\right) & =\theta\left(\boldsymbol{\lambda}^{n}+r \nabla\left(\Psi^{*}+u^{n}\right)-\boldsymbol{f}\right),  \tag{5.39}\\
x & =\left|\boldsymbol{\lambda}^{n}+r \nabla\left(\Psi^{*}+u^{n}\right)-\boldsymbol{f}\right| \tag{5.40}
\end{align*}
$$

we have to find the minimizer of

$$
\begin{equation*}
M(\theta)=\frac{1}{2} \int_{0}^{(\theta x)^{2}} \frac{\chi\left(s^{1 / 2}\right)}{s^{1 / 2}} \mathrm{~d} s+\frac{\tau_{Y}}{H}|\theta| x+\frac{r}{2}(\theta x)^{2}-\theta x \tag{5.41}
\end{equation*}
$$

If $x \leq \tau_{Y} / H$ then $\theta=0$ minimizes $M(\theta)$, thus the fluid will be unyielded. If $x>\tau_{Y} / H$ we solve numerically for $\theta$ the following equation,

$$
\begin{equation*}
\chi(\theta x)+\frac{\tau_{Y}}{H}+r \theta x-x=0 \tag{5.42}
\end{equation*}
$$

which has a unique solution and therefore here the fluid is yielded in this iteration. Finally we update $\boldsymbol{\lambda}$ using (5.35).

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### 5.2.2 Transient model

The derivation of the augmented Lagrangian algorithm for the transient case is basically the same as the one for the steady model. The main difference is with the augmented Lagrangian functional defined as,

$$
\begin{align*}
L_{r}\left(v^{n+1}, \boldsymbol{q}^{n+1}, \boldsymbol{\mu}\right)=F\left(\boldsymbol{q}^{n+1}\right) & +\int_{\Omega} \boldsymbol{\mu} \cdot\left(\nabla v^{n+1}-\boldsymbol{q}^{n+1}\right) \mathrm{d} \Omega+\frac{r}{2} \int_{\Omega}\left|\nabla v^{n+1}-\boldsymbol{q}^{n+1}\right|^{2} \mathrm{~d} \Omega \\
& +\int_{\Omega} \rho\left(\frac{\nabla\left(v^{n+1}-v^{n}\right)}{\Delta t}\right) \cdot \nabla v^{n+1} \mathrm{~d} \Omega \tag{5.43}
\end{align*}
$$

where the super index $v^{n+1}$ means $v\left(t^{n+1}\right)$ where $t^{n+1}=t^{n}+\Delta t$.

Thus the transient version of ALG2 will be,

$$
\begin{array}{r}
\int_{\Omega} \lambda_{k}^{n+1} \cdot \nabla\left(v^{n+1}-u_{k}^{n+1}\right)+\left(\left(r+\frac{\rho}{\Delta t}\right) \nabla u_{k}^{n+1}-\frac{\rho}{\Delta t} \nabla u_{k}^{n}-\boldsymbol{p}_{k-1}^{n+1}\right) \cdot \nabla\left(v^{n+1}-u_{k}^{n+1}\right) \mathrm{d} \Omega \geq 0, \\
\forall v^{n+1} \in V_{0, \rho}, u_{k}^{n+1} \in V_{0, \rho} \\
F\left(\boldsymbol{q}^{n+1}\right)-F\left(\boldsymbol{p}_{k}^{n+1}\right)+\int_{\Omega} \lambda_{k}^{n+1} \cdot\left(\boldsymbol{q}^{n+1}-\boldsymbol{p}_{k}^{n+1}\right)+r\left(\boldsymbol{p}_{k}^{n+1}-\boldsymbol{q}^{n+1}\right) \cdot\left(\boldsymbol{q}^{n+1}-\boldsymbol{p}_{k}^{n+1}\right) \mathrm{d} \Omega \geq 0, \\
\forall \boldsymbol{q}^{n+1} \in H, \boldsymbol{p}_{k}^{n+1} \in H
\end{array}
$$

the subscript $k$ is just the step of ALG2. We will update $\lambda^{n+1}$ using (5.46) and the minimization problem (5.45) corresponds to finding the roots of (5.42).

The only difference is in minimizing (5.44), which is equivalent to solve for each time step the following Poisson problem,

$$
\begin{equation*}
\left(r+\frac{\rho}{\Delta t}\right) \Delta u_{k}^{n+1}=r \nabla \cdot \boldsymbol{p}^{n-1}-\nabla \cdot \lambda^{n}-\frac{\rho}{\Delta t} \nabla u_{k}^{n} \tag{5.47}
\end{equation*}
$$

We solve (5.36) \& (5.47) using a finite element method. In the following section we give a brief outline of the method, which closely follows [43].

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### 5.3 Finite Element discretization

In $\S 5.2$ we stated that the minimization problem (5.33) is equivalent to solve the Poisson problem (5.36), thus the variational problem for (5.36) is,

$$
\begin{equation*}
\text { Find } u \in V_{0, \rho} \text { such that } a(u, v)=b(v) \quad \forall v \in V_{0, \rho} \tag{5.48}
\end{equation*}
$$

where

$$
\begin{aligned}
a(u, v) & =\int_{\Omega} \nabla u \cdot \nabla v \mathrm{~d} \Omega \\
b(v) & =\int_{\Omega} f v \mathrm{~d} \Omega
\end{aligned}
$$

with $f$ being the right hand side of (5.36).
The finite element discretization consists in replacing $V_{0, \rho}$ by a finite dimensional subspace $V_{h}$, i.e. instead of solving (5.48), we solve

$$
\begin{equation*}
\text { Find } u_{h} \in V_{h} \text { such that } a\left(u_{h}, v\right)=b(v) \quad \forall v \in V_{h}, \tag{5.49}
\end{equation*}
$$

this approach is also known as the Galerkin method and the finite-dimensional subspace $V_{h}$ is called an ansatz space.

As an example, let us take $T_{h}$ to be a partition of our domain $\Omega$ into closed triangles $K$ with the following properties

1. $\bar{\Omega}=\bigcup_{K \in T_{h}} K$,
2. For $K, K^{\prime} \in T_{h}, K \neq K^{\prime}$ and interior $(K) \bigcap$ interior $\left(K^{\prime}\right)=\emptyset$.
3. If $K \neq K^{\prime}$ but $K \bigcap K^{\prime}=\emptyset$, then $K \bigcap K^{\prime}$ is either a point or a common edge of $K$ and $K^{\prime}$.

Let us define $h$ as

$$
h=\max \left\{\operatorname{diam}(K) \mid K \in T_{h}\right\},
$$

$K$ is called an element of the partition and the vertices of the triangle are called the nodes. We call this a linear finite element approximation of our problem if $V_{h}$ is defined as follows,

$$
V_{h}=\left\{u \in C(\bar{\Omega})|u|_{K} \in P_{1}(K) \quad \forall K \in T_{h}, u=0 \text { on } \partial \Omega,\right\}
$$

where $P_{1}(K)$ is the set of polynomials of first degree (in 2 variables), i.e. $p \in P_{1}(K) \Leftrightarrow p(x, y)=$ $\alpha+\beta x+\gamma y \quad \forall(x ; y) \in K$ and for $\alpha, \beta, \gamma \in \boldsymbol{R}$.

Now, let us consider the local interpolation problem

$$
\begin{equation*}
p\left(a_{i}\right)=u_{i} \text { for } i=1,2,3 \tag{5.50}
\end{equation*}
$$

where $p \in P_{1}(K)$ and $a_{1}, a_{2}, a_{3}$ are the vertices of $K$, thus

$$
p(x, y)=u_{1} \varphi_{1}(x, y)+u_{2} \varphi_{2}(x, y)+u_{3} \varphi_{3}(x, y)
$$

where

$$
\varphi_{i}\left(a_{j}\right)=\delta_{i j}
$$

has a unique solution, see [43] and that $\left.u\right|_{K}$ can be composed continuously, i.e. if $p_{K} \in P_{1}(K)$ is the unique solution of ( 5.50 ) then

$$
u(x)=p_{K}(x) \quad \text { for } x \in K \in T_{h} .
$$

Stability and convergence results for the finite element method for linear triangular elements can be found in [43]. Here we just state the result.

If the weak solution $u$ of (5.36) has weak derivatives of second order, then

$$
\left\|u-u_{h}\right\|_{1} \leq C h
$$

where $C$ depends on $u$ but not on $h$.

We use this type of linear elements with linear basis functions $\varphi_{i}$ for solving numerically system (2.63) in $\S 3.3 .3$. For the purposes of this chapter, we choose to work with quadrilateral elements and bilinear basis functions. All results stated above can be extended to this type of finite element discretization, (see [43] chap. 3 for reference).

### 5.4 Flux Corrected Transport Algorithm

Now we turn our to attention to equation (5.8). Note that this concentration equation is coupled with system (5.1)-(5.7) via the stream function. Thus we will solve the full coupled
system (5.1)-(5.8), using the augemented Lagrangian method in a finite element setting for the stream function and a Flux Corrected Transport (FCT) algorithm in a finite volume setting for the concentration. We describe the FCT method below.

For a full description of the algorithm we refer to the reader to [89]. Here we just give an outline of the method applied to our particular problem.

It is well know that high order schemes present numerical dispersion near discontinuities and regions of high gradients, which translates into the growth of unphysical oscillations in the conserved quantity, (here $\bar{c}$ ). In contrast, low order schemes do not suffer from this dispersion, but numerical diffusion is introduced in to the solution. Flux limiting methods, such as Flux-limiters and FCT have developed over the years to minimize both numerical dispersion and diffusion. We will concentrate on FCT schemes which limit the flux by minimizing the overshoots and undershoots in the higher order numerical solution. The discretization will be made using a finite volume method which is well known to be conservative. The main reason that we use quadrilateral elements with bilinear basis functions for the stream function is that we will calculate the concentration field at the middle of each element, thus the calculation of numerical fluxes across the boundary of each element is easier to obtain than with triangular elements.

Let us write equation (5.8) in the following form,

$$
\begin{equation*}
\frac{1}{\epsilon} \frac{\partial C}{\partial t}+\frac{\partial f}{\partial \phi}+\frac{\partial g}{\partial \xi}=0 \tag{5.51}
\end{equation*}
$$

where

$$
\begin{aligned}
C & =H \bar{c} \\
f & =\bar{v} H \bar{c}=-\frac{\Psi_{\xi}}{H} H \bar{c} \\
g & =\bar{w} H \bar{c}=\frac{\Psi_{\phi}}{H} H \bar{c}
\end{aligned}
$$

Calculating the numerical transport fluxes across each element and using a forward Euler scheme
for time we have:

$$
\begin{equation*}
C_{i, j}^{n+1}=C_{i, j}^{n}-\frac{\Delta t}{\Delta \phi}\left(F_{i+\frac{1}{2}, j}^{n}-F_{i-\frac{1}{2}, j}^{n}\right)-\frac{\Delta t}{\Delta \xi}\left(G_{i, j+\frac{1}{2}}^{n}-G_{i, j-\frac{1}{2}}^{n}\right) \tag{5.52}
\end{equation*}
$$

where $F$ and $G$ are the discrete numerical fluxes. Note that (5.52) is a second order approximation in space for $C$, thus numerical dispersion will be introduced in the solution, thus we use a FCT scheme to minimize this effect.

First we calculate $F_{i+\frac{1}{2}, j}^{L}$ and $G_{i, j+\frac{1}{2}}^{L}$ using a low order scheme in the following way,

$$
F_{i+\frac{1}{2}, j}^{L}= \begin{cases}V_{i+\frac{1}{2}, j}^{n} C_{i, j}^{n} & \text { if } V_{i+\frac{1}{2}, j}^{n} \geq 0  \tag{5.53}\\ V_{i+\frac{1}{2}, j}^{n} C_{i+1, j}^{n} & \text { if } V_{i+\frac{1}{2}, j}^{n}<0\end{cases}
$$

and

$$
G_{i, j+\frac{1}{2}}^{L}= \begin{cases}W_{i, j+\frac{1}{2}}^{n} C_{i, j}^{n} & \text { if } W_{i, j+\frac{1}{2}}^{p} \geq 0  \tag{5.54}\\ W_{i, j+\frac{1}{2}}^{n} C_{i, j+1}^{n} & \text { if } W_{i, j+\frac{1}{2}}^{n}<0\end{cases}
$$

Now we compute the low order time advanced solution,

$$
\begin{equation*}
C_{i, j}^{l o}=C_{i, j}^{n}-\frac{\Delta t}{\Delta \phi}\left(F_{i+\frac{1}{2}, j}^{L}-F_{i-\frac{1}{2}, j}^{L}\right)-\frac{\Delta t}{\Delta \xi}\left(G_{i, j+\frac{1}{2}}^{L}-G_{i, j-\frac{1}{2}}^{L}\right) . \tag{5.55}
\end{equation*}
$$

Then we compute $F_{i+\frac{1}{2}, j}^{H}$ and $G_{i, j+\frac{1}{2}}^{H}$ using a high order scheme in the following way

$$
\begin{align*}
F_{i+\frac{1}{2}, j}^{H} & =\frac{1}{2}\left(F_{i, j}^{n}+F_{i+1, j}^{n}\right)  \tag{5.56}\\
G_{i, j+\frac{1}{2}}^{H} & =\frac{1}{2}\left(G_{i, j}^{n}+G_{i, j+1}^{n}\right) . \tag{5.57}
\end{align*}
$$

We compute the antidiffusive fluxes, to minimize the diffusion introduced by using a low order scheme,

$$
\begin{align*}
A_{i+\frac{1}{2}, j} & =\left(F_{i+\frac{1}{2}, j}^{H}-F_{i+\frac{1}{2}, j}^{L}\right)  \tag{5.58}\\
A_{i, j+\frac{1}{2}} & =\left(G_{i, j+\frac{1}{2}}^{H}+G_{i, j+\frac{1}{2}}^{L}\right) . \tag{5.59}
\end{align*}
$$

We limit this antidiffusive fluxes to minimize the dispersion effect that the high order approximation has on the solution,

$$
\begin{array}{ll}
A_{i+\frac{1}{2}, j}^{l}=A_{i+\frac{1}{2}, j} L_{i+\frac{1}{2}, j} & \text { for } 0 \leq L_{i+\frac{1}{2}, j} \leq 1 \\
A_{i, j+\frac{1}{2}}^{l}=A_{i, j+\frac{1}{2}} L_{i, j+\frac{1}{2}} & \text { for } 0 \leq L_{i, j+\frac{1}{2}} \leq 1 \tag{5.61}
\end{array}
$$

and finally we apply the limited antidiffusive fluxes to the get the final solution,

$$
\begin{equation*}
C_{i, j}^{n+1}=C_{i, j}^{l o}-\frac{\Delta t}{\Delta \phi}\left(A_{i+\frac{1}{2}, j}^{l}-A_{i-\frac{1}{2}, j}^{l}\right)-\frac{\Delta t}{\Delta \xi}\left(A_{i, j+\frac{1}{2}}^{l}-A_{i, j-\frac{1}{2}}^{l}\right) \tag{5.62}
\end{equation*}
$$

The only thing that remains unclear is how the corrections $L_{i+\frac{1}{2}, j}$ and $L_{i, j+\frac{1}{2}}$ will be chosen to limit $C_{i, j}^{n+1}$ so that it does not exceed a maxim value $C_{i, j}^{\text {max }}$ nor a minimum value $C_{i, j}^{m i n}$. This will be achieved in the following way.

First we define the following quantities,

- 1. We compute the sum of all antidiffusive fluxes into the grid point $(i, j)$ (the centre of our element),

$$
P_{i, j}^{+}=\max \left(0, A_{i-\frac{1}{2}, j}\right)-\min \left(0, A_{i+\frac{1}{2}, j}\right)+\max \left(0, A_{i, j-\frac{1}{2}}\right)-\max \left(0, A_{i, j+\frac{1}{2}}\right)
$$

2. We compute the sum of all antidiffusive fluxes away from the grid point $(i, j)$,

$$
P_{i, j}^{-}=\max \left(0, A_{i+\frac{1}{2}, j}\right)-\min \left(0, A_{i-\frac{1}{2}, j}\right)+\max \left(0, A_{i, j+\frac{1}{2}}\right)-\max \left(0, A_{i, j-\frac{1}{2}}\right)
$$

3. Define

$$
\begin{aligned}
C_{i, j}^{\max } & =\max \left(C_{i, j}^{n}, C_{i+1, j}^{n}, C_{i-1, j}^{n}, C_{i, j+1}^{n}, C_{i, j-1}^{n}\right) \\
C_{i, j}^{\min } & =\min \left(C_{i, j}^{n}, C_{i+1, j}^{n}, C_{i-1, j}^{n}, C_{i, j+1}^{n}, C_{i, j-1}^{n}\right)
\end{aligned}
$$

4. Define

$$
\begin{aligned}
Q_{i, j}^{+} & =\left(C_{i, j}^{\max }-C_{i, j}^{l o}\right) \Delta \phi \Delta \xi \\
Q_{i, j}^{-} & =\left(C_{i, j}^{l o}-C_{i, j}^{m i n}\right) \Delta \phi \Delta \xi
\end{aligned}
$$

5. Define

$$
\begin{align*}
& R_{i, j}^{+}= \begin{cases}\min \left(1, Q_{i, j}^{+} / P_{i, j}^{+}\right) & \text {if } P_{i, j}^{+}>0 \\
0 & \text { if } P_{i, j}^{+}=0\end{cases}  \tag{5.63}\\
& R_{i, j}^{-}= \begin{cases}\min \left(1, Q_{i, j}^{-} / P_{i, j}^{-}\right) & \text {if } P_{i, j}^{-}>0 \\
0 & \text { if } P_{i, j}^{-}=0\end{cases} \tag{5.64}
\end{align*}
$$

6. Finally, the corrections become

$$
\begin{align*}
& L_{i+\frac{1}{2}, j}= \begin{cases}\min \left(R_{i+1, j}^{+}, R_{i, j}^{-}\right) & \text {if } A_{i+\frac{1}{2}, j} \geq 0 \\
\min \left(R_{i, j}^{+}, R_{i+1, j}^{-}\right) & \text {if } A_{i+\frac{1}{2}, j}<0\end{cases}  \tag{5.65}\\
& L_{i, j+\frac{1}{2}}= \begin{cases}\min \left(R_{i, j+1}^{+}, R_{i, j}^{-}\right) & \text {if } A_{i, j+\frac{1}{2}} \geq 0 \\
\min \left(R_{i, j}^{+}, R_{i, j+1}^{-}\right) & \text {if } A_{i, j+\frac{1}{2}}<0 .\end{cases} \tag{5.66}
\end{align*}
$$

Now, we are in the position to solve the system of equations (5.1)-(5.7) numerically and investigate the effects of pulsation of the flow rate in the displacement flow.

### 5.5 Numerical simulations for the annular displacement with pulsating flow rate

In this section we are interested in the behaviour of the displacement flow in the presence of a pulsating flow rate, thus (5.5) becomes

$$
\begin{equation*}
\Psi(1, \xi, t)=1+\delta_{p} \sin \omega t \tag{5.67}
\end{equation*}
$$

Thus we will solve system (5.1)-(5.7) with (5.67) instead of (5.5), using the augmented Lagrangian algorithm in a finite element setting for the velocity field and for the concentration we will use the FCT scheme described in the previous section.

### 5.5.1 Validation

To benchmark our code we follow the approach in [58]. For small eccentricities and certain combinations of the physical parameters, it is possible to find an analytical solution that describes the shape of a steadily advancing displacement front, see [58]. In [58] it is shown that the $c=0.5$ level set, viewed in a moving frame of reference, converges to the steady profile. Here we have performed similar benchmarking computations and final similar results. Figure 5.1 shows an example of these computations. We solve the steady model for the velocity, so that the only time dependence is through the concentration equation, i.e. $\delta_{p}=0$ in (5.67). The following physical and rheological parameters are used, $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2$, $\rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9, \tau_{Y, 1}=0.7 e=0.0$ and $\beta=0$. Results are consistent

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with the ones found in [58]. The validity of our code for the transient velocity field is shown in §3.3.3. Assuming fixed concentration we have seen the decay of the transient velocity to the steady state, at constant flow rate $\left(\delta_{p}=0\right)$.


Figure 5.1: Convergence of computational solution to stable steady state displacement: a) Successive contours of $\bar{c}(\phi, z, t)=0.5$ at times $t=0,0.2,0.4, \ldots 6$ b) Analytical solution for the steady displacement. c) Convergence of the wide and narrow side interface positions to the steady state. Rheological and physical parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2$, $\rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9, \tau_{Y, 1}=0.7 e=0.0$ and $\beta=0$.

### 5.5.2 Pulsation effects

## Pulsation at the beginning of the displacement

In Figure 5.2 we show the initial condition for all the following displacements. One third of the eccentric Hele-Shaw cell is full of displacing fluid, cement, the remaining two thirds is full of the displaced fluid, mud. We will always begin with a flat interface, perpendicular to the $\xi$ axis between the fluids. For the rest of the section we will fix the rheological and some physical parameters to: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9$, $\tau_{Y, 1}=0.7, \beta=0$. We will study the effects of eccentricity, the amplitude and phase of the pulsation and $\epsilon$ on the displacement fronts.


Figure 5.2: Initial condition for the displacements.

In Figure 5.3 we show the unsteady displacement of cement displacing mud. The eccentricity is $e=0.3$. As we can see from the velocity field plots, both fluids are fully yielded along the annulus. We show times $T / 4, T / 2,3 T / 4$ and $T$, where $T$ is the period of the pulsation. For the steady state velocity model is defined as $T=2 \pi / \omega$, for the transient velocity model is defined as $T=2 \pi / \omega \epsilon$. Recall that velocity changes happen on a fast time scale. Thus, in order to make a comparison of our models, we will work with the pulsation on the same time scale as the concentration. Figure 5.4 shows the effects of the pulsation in the steady model, we increase the amplitude to $\delta_{p}=0.1$. No major effects can be seen on the displacement front and velocity fields, as expected, both fluids are fully mobile along the annulus. In Figures 5.5-5.7 we investigate the effects of the pulsation in the transient model. Also we are interested in the results of varying $\epsilon$, recall that $\epsilon$ is the ration between the advective and the viscous time scales. For Figure 5.5 we set $\epsilon=0.1$ and $\omega=10$, note that if $\epsilon \rightarrow 0$, we recover the pseudo-steady model, this is consistent with our findings where we can see that the effects of the pulsation on the displacement front is minimal as in previous figures. If $\epsilon \sim O(1)$, viscous and advective time scales have the same order, thus the perturbation is picked up by the transient behaviour of the velocity as shown in Figure 5.6. Figures 5.7-5.9 show the reversed case, i.e. where viscous effects happen in a slow time scale, $\epsilon=10$. In Figure 5.7 shows the case when the
perturbation of the flow rate is fast, $\omega=10$. Because we show the displacement front over the full period of the pulsation, time is not large enough for viscous effects to take place. When $\omega=1$, Figure 5.8 , time is large enough to see viscous effects and the unsteady displacement takes place. Figure 5.9 shows the effects of a perturbation of the flow rate over a long time, we see clearly the viscous effects had taken place and a very long finger has developed along the annulus. Note that for Figures $5.8 \& 5.9$ we solved the problem in a moving frame of reference, the finger grows so long that we were unable to consider a domain large enough to capture the full finger.

For Figures 5.10-5.12 we increase the eccentricity to 0.5. As shown in [58], as eccentricity increases we attain displacements that burrow slowly up the narrow side of the annlus, i.e. although the interface is propagating along the annulus, in the far field the velocity field is unyielded in the narrow side. In Figure 5.10 we got the same results as the ones found in [58], this was expected as we solve the steady-state velocity model. Figure 5.11 shows the effects pulsation has on the steady state model. We fix the amplitude of the pulsation, $\delta_{p}=0.1$. As above, we cannot see major changes in the displacement front but an interesting phenomena is the decrease of the unyielded regions on the far field, as is clearly seen in the velocity field plots. On the other hand, Figure 5.12 shows that the transient model actually captures the effects of the pulsation. When we reach the maximum point on the pulsation, Figures 5.12 a)-b) both fluids are fully yielded, even in the far field, as we decrease the intensity of the flow rate stagnant zones appear until we reach the minimum point of the pulsation and both fluids are unyielded on the narrow part of the annulus for the far field. When the intensity of the flow increases again we fully remove the stagnant zones. One interesting phenomena is that in contrast to the steady state model, the velocity field of the transient model does not show the same burrowing motion. In Figures 5.10 and 5.11 it is clear that in the vicinity of the interface the displacing fluid is forced to the narrow side and the displaced fluid is forced away from the narrow side. For the transient model this burrowing motion is expanded along the annulus not only near the interface. This will lead us to investigate the effects of the pulsation technique and transient effects of the velocity field in the removal of a mud channel. Figure
5.13 shows the predicted mud channel width using the steady model without pulsing the flow rate. Figures 5.14 and 5.14 show the width of the mud channel with pulsation amplitudes $\delta_{p}=0.1$ and $\delta_{p}=0.2$ respectively. As in the previous cases there is no clear evidence of a change in the displacement. Figures 5.16 and 5.17 show the effects of a pulsating flow rate on the transient model with amplitudes, $\delta_{p}=0.1$ and $\delta_{p}=0.2$ respectively. Clearly the effects of the pulsation are fully captured by the transient velocity field, as above we have a burrowing motion expanding along the annlus, i.e. even though we have a mud channel in the narrow side of the annulus the far field velocity is clearly yielded outside this region. As we increase the magnitude of the pulsation, this motion expands further to the narrow side of the annulus, and therefore a decrease of the width of the mud channel is achieved.

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Figure 5.3: Displacement flow in eccentric annulus, interface propagation. Unsteady displacement. Steady state velocity model. Period $T=2 \pi / \omega, \omega=10$ and $\delta_{p}=0$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9$, $\tau_{Y, 1}=0.7 e=0.3, \beta=0$.

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Figure 5.4: Displacement flow in eccentric annulus, interface propagation. Unsteady displacement. Steady state velocity model. Period $T=2 \pi / \omega, \omega=10$ and $\delta_{p}=0.1$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9$, $\tau_{Y, 1}=0.7 e=0.3, \beta=0$.

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Figure 5.5: Displacement flow in eccentric annulus, interface propagation. Unsteady displacement. Transient velocity model. Period $T=2 \pi / \omega \epsilon, \omega=10, \delta_{p}=0.1$ and $\epsilon=0.1$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1$, $\tau_{Y, 1}=0.9, \tau_{Y, 1}=0.7 e=0.3, \beta=0$.

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Figure 5.6: Displacement flow in eccentric annulus, interface propagation. Unsteady displacement. Transient velocity model. Period $T=2 \pi / \omega \epsilon, \omega=10, \delta_{p}=0.1$ and $\epsilon=1$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1$, $\tau_{Y, 1}=0.9, \tau_{Y, 1}=0.7 e=0.3, \beta=0$.

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Figure 5.7: Displacement flow in eccentric annulus, interface propagation. Unsteady displacement. Transient velocity model. Period $T=2 \pi / \omega \epsilon, \omega=10, \delta_{p}=0.1$ and $\epsilon=10$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1$, $\tau_{Y, 1}=0.9, \tau_{Y, 1}=0.7 e=0.3, \beta=0$.

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Figure 5.8: Displacement flow in eccentric annulus, interface propagation. Unsteady displacement. Transient velocity model. Period $T=2 \pi / \omega \epsilon, \omega=1, \delta_{p}=0.1$ and $\epsilon=10$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4 . \mathrm{g})-\mathrm{h}) T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1$, $\tau_{Y, 1}=0.9, \tau_{Y, 1}=0.7 e=0.3, \beta=0$.

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Figure 5.9: Displacement flow in eccentric annulus, interface propagation. Unsteady displacement. Transient velocity model. Period $T=2 \pi / \omega \epsilon, \omega=0.1, \delta_{p}=0.1$ and $\epsilon=10$. Interface position and velocity field for times: a)-b) $T / 4$. Physical and rheological parameters: $\kappa_{1}=0.5$, $\kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9, \tau_{Y, 1}=0.7 e=0.3, \beta=0$.

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Figure 5.10: Displacement flow in eccentric annulus, interface propagation. Unsteady displacement. Steady state velocity model. Period $T=2 \pi / \omega, \omega=10$ and $\delta_{p}=0$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9$, $\tau_{Y, 1}=0.7 e=0.5, \beta=0$.

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Figure 5.11: Displacement flow in eccentric annulus, interface propagation. Unsteady displacement. Steady state velocity model. Period $T=2 \pi / \omega, \omega=10$ and $\delta_{p}=0.1$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9$, $\tau_{Y, 1}=0.7 e=0.5, \beta=0$.

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Figure 5.12: Displacement flow in eccentric annulus, interface propagation. Unsteady displacement. Transient velocity model. Period $T=2 \pi / \omega \epsilon, \omega=10, \delta_{p}=0.1$ and $\epsilon=0.6$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1$, $\tau_{Y, 1}=0.9, \tau_{Y, 1}=0.7 e=0.5, \beta=0$.

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Figure 5.13: Displacement flow in eccentric annulus, interface propagation. Mud channel formation. Steady state velocity model. Period $T=2 \pi / \omega, \omega=10$ and $\delta_{p}=0$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9$, $\tau_{Y, 1}=0.7 e=0.8, \beta=0$.

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Figure 5.14: Displacement flow in eccentric annulus, interface propagation. Mud channel formation. Steady state velocity model. Period $T=2 \pi / \omega, \omega=10$ and $\delta_{p}=0.1$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9$, $\tau_{Y, 1}=0.7 e=0.8, \beta=0$.

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Figure 5.15: Displacement flow in eccentric annulus, interface propagation. Mud channel formation. Steady state velocity model. Period $T=2 \pi / \omega, \omega=10$ and $\delta_{p}=0.2$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9$, $\tau_{Y, 1}=0.7 e=0.8, \beta=0$.

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Figure 5.16: Displacement flow in eccentric annulus, interface propagation. Mud channel formation. Transient velocity model. Period $T=2 \pi / \omega \epsilon, \omega=10, \delta_{p}=0.1$ and $\epsilon=0.6$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4 . \mathrm{g})-\mathrm{h}) T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1$, $\tau_{Y, 1}=0.9, \tau_{Y, 1}=0.7 e=0.8, \beta=0$.

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Figure 5.17: Displacement flow in eccentric annulus, interface propagation. Mud channel formation. Transient velocity model. Period $T=2 \pi / \omega \epsilon, \omega=10, \delta_{p}=0.2$ and $\epsilon=0.6$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1$, $\tau_{Y, 1}=0.9, \tau_{Y, 1}=0.7 e=0.8, \beta=0$.

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## Pulsation after mud channel forms

Now we turn our attention to the worst case scenario, when a mud channel forms in the narrow part of the annulus. We tried to simulate the case in which the mud channel has already formed and the flow may be described as a parallel flow in a Hele-Shaw cell with the fluid on the narrow side static. We use the same rheological and physical parameter as the previous section, here the interface between the fluids is placed at $\phi=0.8$. Figure 5.18 shows the effects of the pulsation on the steady state velocity model with an amplitude of $\delta_{p}=0.2$. This model predicts that the mud channel will remain static. In Figure 5.19 we show the effects the pulsation has on the transient model with an amplitude of $\delta_{p}=0.2$. Clearly as we go over the full period of the pulsation the mud channel slowly begins to yield until it is fully moving, then goes back to the static mud channel after the pulsation period is over. Thus in each pulsation the mud channel will yield, move up the narrow side for a short period of time and stop again. Another interesting behaviour is that it seems that the flow is stable, no growing instabilities were to be seen at the interface. Actually, the interface remained unperturbed for all time and thus the removal of the mud channel was unsuccessful. One of the reasons we can think of has to be with the choice of the concentration model to simulate this problem. Perhaps an interface tracking model is a better choice for this type of problems. Actually, we tried to solve the interface tracking model using a conformal mapping technique. Due to our transformation, we lost control over the hyperbolic problem defining the interface, allowing us to use only a low order method to solve the problem. We think that the numerical diffusion was to large and no effects of the perturbed flow rate could be seen. An adaptive technique could be a better approach to attack this problem. We cannot therefore be completely sure if pulsation of the flow rate, after the mud channel has already formed, helps its removal. Until we extend our approach to an adaptive technique we cannot derive any firm conclusion. This remains as future work and will not be investigated in this thesis.

As a conclusion to the chapter, the transient model (5.1)-(5.7) with (5.67) instead of (5.5) fully captures the perturbations of the velocity field. This may lead to a reduction of the mud channel telling us that pulsation of the flow rate at the beginning of the displacement may be

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used as a tool to prevent or remove the mud channels along the annulus.


Figure 5.18: Displacement flow in eccentric annulus, interface propagation. Mud channel. Steady state velocity model. Period $T=2 \pi / \omega, \omega=10$ and $\delta_{p}=0.2$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4$. g)-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9$, $\tau_{Y, 1}=0.7 e=0.8, \beta=0$.

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Figure 5.19: Displacement flow in eccentric annulus, interface propagation. Mud channel. Transient velocity model. Period $T=2 \pi / \omega \epsilon, \omega=10, \delta_{p}=0.2$ and $\epsilon=0.6$. Interface position and velocity field for times: a)-b) $T / 4$. c)-d) $T / 2$. e)-f) $3 T / 4 . \mathrm{g})$-h) $T$. Physical and rheological parameters: $\kappa_{1}=0.5, \kappa_{2}=0.4, m_{1}=1, m_{2}=1.2, \rho_{1}=1, \rho_{2}=0.9, \tau_{Y, 1}=1, \tau_{Y, 1}=0.9$, $\tau_{Y, 1}=0.7 e=0.8, \beta=0$.

## Chapter 6

## Summary and conclusions

This thesis is concerned with modelling and analysis of various transient effects in primary cementing of an oil well. This process involves the displacement of one shear-thinning yield stress fluid by another along a narrow eccentric annular duct.

### 6.1 Main contributions of the thesis

The contributions of this thesis fall into 4 categories:

- We developed a transient version of the model found in [10] and demonstrated its validity for small $\epsilon$.
- The nonlinear evolution equation for the velocity field (stream function) is a third order partial differential equation, in 2 spatial dimensions and in time. We proved that there exists a unique solution to this equation and that the solution changes continuously with the data, i.e. well-posedness. These results are detailed and are summarised in chapter 3.
- In the case that a long finger develops along the annulus during displacement, we were able to analyse the linear stability of the limiting parallel multi-layer flows. Specific results are summarised below in $\S 6.2$.
- In the case that a mud channel develops in the narrow side of the annulus, we demonstrated that by perturbing the flow rate, via pulsation, the width of the mud channel can be decreased. It is necessary to pulse the flow rate during displacement, rather than after the mud channel has already formed. Further results are summarised below in $\S 6.3$


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### 6.2 Interfacial instabilities

When a long finger develops we find essentially a multi-layer parallel flow along the length of the annulus. In chapter 4 we studied the linear stability of these flows for the case when both fluids are yielded at the interface. This stability problem depends on 14 dimensionless parameters in its full generality and hence very difficult to study analytically. Thus, we first consider two simpler cases, parallel flow of two Newtonian fluids and parallel flow of two power law fluids, both in a vertical Hele-Shaw cell with constant height, i.e. a concentric annulus. For both cases we found that buoyancy plays a major role in causing instability. In the absence of density difference it seems that the flow is linearly stable. Inclination of the Hele-Shaw cell has destabilizing effects, if the heavier fluid is on top of the lighter fluid, the flow may be unstable. Again, in the absence of buoyant forces, viscous effects and inclination are not strong enough to drive an instability.

The stability problem for yield stress fluids has to be solved via computation. Here we solve the full problem, i.e. the parallel flow of two yield stress fluids in a vertical or near vertical eccentric Hele-Shaw cell. We consider the case when both fluids are yielded at the interface. If the eccentricity is equal to zero, i.e. a concentric annulus, we find analogous results to the Newtonian and power law fluid cases: buoyancy is the key parameter for instability. Although buoyancy forces are the primary reason for instability to arise, viscosity differences play a secondary role in the instability. For example, for certain values of density and viscosity if $\rho_{1}>\rho_{2}$ and $\kappa_{1}>\kappa_{2}$ the flow may be stable, but if $\rho_{1}>\rho_{2}$ and $\kappa_{1}<\kappa_{2}$ the flow may be linearly unstable. This shows that even though density differences drive the unstable behaviour, viscosity differences have stabilizing or destabilizing effects. This phenomenon is in contrast with the work of Gondret \& Rabaud, [33]. They showed that in a horizontal Hele-Shaw cell, viscosity differences only determine the wave velocity of the growing instability.

An interesting phenomena that is worth to mention is the inclusion of eccentricity in to the stability problem. Eccentricity can have stabilizing or destabilizing effects. For example, in the case of a concentric annulus we show that for a certain range of values of density and viscosity
if $\rho_{1}<\rho_{2}$ and $\kappa_{1}>\kappa_{2}$, the flow was linearly unstable for all wave lengths. If we increase the eccentricity, the flow will be linearly stable for long wave lengths. In contrast, if the more viscous fluid is also heavier, the flow was linearly stable for all wave lengths for the concentric annulus. If we now increase the eccentricity the flow will become linearly unstable for long wave lengths. From such results we inferred that all significant changes seem to happen for long wave lengths. Thus we considered the long wave length asymptotic behaviour of our problem.

We presented several stability maps in $\S 4.7$. As expected for the concentric annulus, in the absence of buoyancy the flow will be linearly stable. When eccentricity is different from zero, all the rheological parameters as well as buoyancy affect the instability regions. Thus, even without the presence of buoyancy forces and inertia, the flow may be unstable. We do not believe that these instabilities have been identified before.

In [59] the authors use a lubrication model to predict whether a good displacement takes place or not. Surprisingly, our unstable domain occurs only as a subset of the unsteady displacement domain predicted by [59]. This means that if a long finger develops during the displacement, for certain parameter ranges the parallel flow may be unstable. Therefore, in unsteady displacements this suggests that mixing may occur at the interface of the finger due to the instability that we study here. It is not clear why there is apparently no interfacial instability for parameter regimes for which the annular displacement would be steady. Note that for such parameters, a base parallel flow is a perfectly well-defined solution of the two fluid problem.

### 6.3 Mud channel removal

In chapter 5 we studied the effects of a pulsatile flow rate for the removal of a mud channel. This is via numerical simulation. In chapter 4 we showed that if one of the fluids is unyielded at the interface the flow will be linearly stable.

We investigated the effects that pulsation has on the mud channel by first using the steady state velocity model. If we begin the pulsation after the mud channel already formed, this model predicts that the mud will remain static. On the other hand, if we use our transient model,

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it predicts that as we go over one period of the pulsation, the mud channel slowly begins to yield until the mud is fully mobile and goes back to be static after the period is over. Thus, in each pulsation, the mud channel will yield, moves up the narrow side of the annulus for a short period of time and then stop again. If instead we pulsate the flow rate from the beginning of the displacement, the transient model predicts that a reduction of the mud channel width takes place. The steady state velocity model with a pulsatile flow rate shows no mayor changes with respect to the steady state velocity model with constant flow rate. Therefore, the transient model is necessary to fully capture the effects of the perturbation of the flow rate.

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